

# Intertheoretic Implications of Non-Relativistic Quantum Field Theories

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1. NQFTs and Particles
2. Newtonian Quantum Gravity
3. Intertheoretic Relations

## ■ 1. NQFTs and Particles

- **Relativistic quantum field theory (RQFT)** = A QFT invariant under the symmetries of a Lorentzian spacetime.
- **Non-relativistic quantum field theory (NQFT)** = A QFT invariant under the symmetries of a classical spacetime.

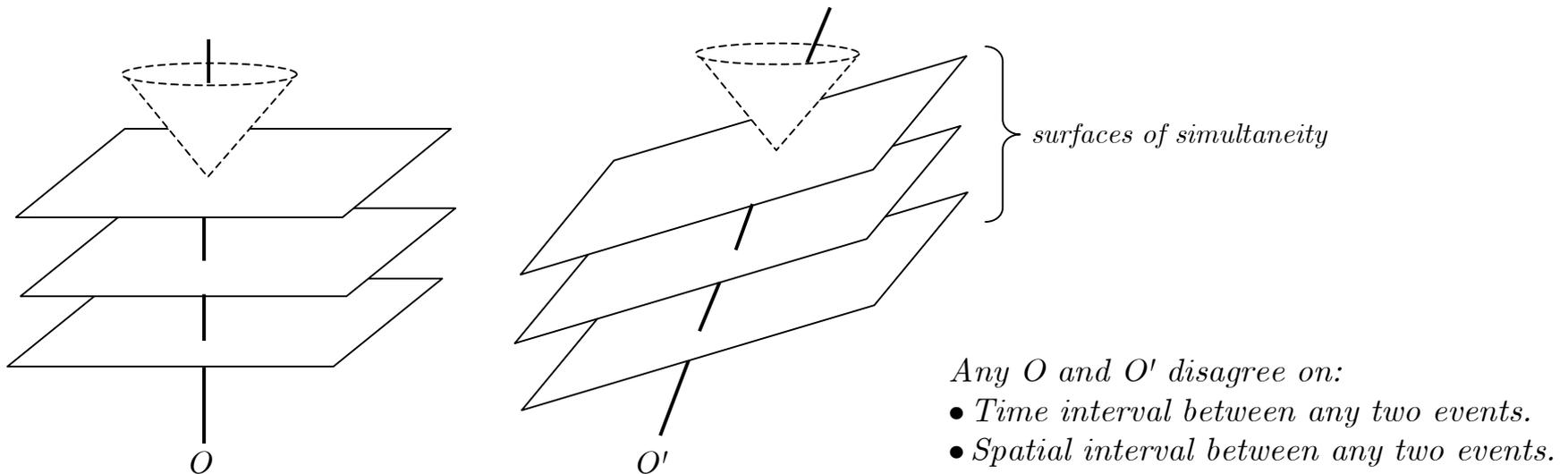
# 1. NQFTs and Particles

*Arena for RQFTs:* Lorentzian spacetime  $(M, g_{ab})$ .

- $g_{ab}$  - pseudo-Riemannian metric with Lorentzian signature  $(1, 3)$ .
- $\nabla_a g_{bc} = 0$  for unique  $\nabla_a$  (*compatibility*)

*Ex. 1:* Minkowski spacetime (*spatiotemporally flat*):  $R^a{}_{bcd} = 0$ .

- No unique way to separate time from space:



- Symmetry group generated by  $\mathcal{L}_x g_{ab} = 0$ . (*Poincaré group*)

## ■ 1. NQFTs and Particles

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*Ex. 1:* Minkowski spacetime (*spatiotemporally flat*):  $R^a{}_{bcd} = 0$ .

*Ex. 2:* Vacuum Einstein spacetime (*Ricci flat*):  $R_{ab} = 0$ .

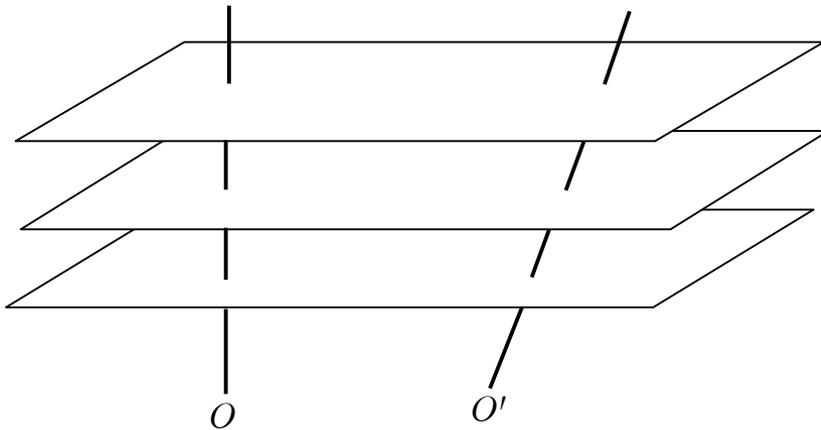
### *Comparison:*

- *Different* metrical structure, *different* curvature, same metric signature (*i.e.*, "in the small", isomorphic to Minkowski spacetime).
- Different types of RQFTs, in flat (Minkowski) and curved Lorentzian spacetimes.

# ■ 1. NQFTs and Particles

**Arena for NQFTs:** Classical spacetime  $(M, h^{ab}, t_{ab}, \nabla_a)$ .

- $h^{ab}, t_{ab}$  - degenerate metrics with signatures  $(0, 1, 1, 1)$  and  $(1, 0, 0, 0)$ .
- $h^{ab}t_{ab} = 0$  (*orthogonality*)
- $\nabla_c h^{ab} = 0 = \nabla_c t_{ab}$  (*compatibility*)  $\Rightarrow$  fails to uniquely determine  $\nabla_a$
- Unique way exists to separate time from space:



*Any  $O$  and  $O'$  agree on:*

- *Time interval between any two events.*
- *Spatial interval between any two simultaneous events.*

- Symmetry group generated by  $\mathcal{L}_x h^{ab} = \mathcal{L}_x t_{ab} = 0$ .

# ■ 1. NQFTs and Particles

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**Ex. 1:** Neo-Newtonian spacetime (*spatiotemporally flat*):  $R^a_{bcd} = 0$ .

- Symmetry group generated by  $\mathfrak{L}_x h_{ab} = \mathfrak{L}_x t_{ab} = \mathfrak{L}_x \Gamma^a_{bc} = 0$ . (*Galilei group*)

**Ex. 2:** Maxwellian spacetime (*rotationally flat*):  $R^{ab}_{cd} = 0$ .

- Symmetry group generated by  $\mathfrak{L}_x h_{ab} = \mathfrak{L}_x t_{ab} = \mathfrak{L}_x \Gamma^{ab}_c = 0$ . (*Maxwell group*)

## **Comparison:**

- *Same* metrical structure, *different* curvature.
- Different types of NQFTs, in flat (Neo-Newtonian) and curved classical spacetimes.

# ■ 1. NQFTs and Particles

## Received View on Particles:

(Arageorgis, Earman, Ruetsche 2003; Halvorson 2007; Halvorson and Clifton 2002; Fraser 2008)

## Necessary conditions for a particle interpretation:

- (A) The QFT must admit a Fock space formulation in which *local number operators* appear that can be interpreted as acting on a state of the system associated with a bounded region of spacetime and returning the number of particles in that region.
  
- (B) The QFT must admit a *unique* Fock space formulation in which a *total number operator* appears that can be interpreted as acting on a state of the system and returning the total number of particles in that state.

# ■ 1. NQFTs and Particles

Claim 1: Conditions (A) and (B) fail in RQFTs.

## Against (B) in RQFTs:

- Problem of Privilege: RQFTs admit unitarily inequivalent Fock space representations of their CCRs.
- Minkowski spacetime exemption? Kay (1979): Minkowski quantization is unique up to unitary equivalence.
- But: The Unruh Effect (in one guise) says: "No!" (at least to some authors).
- In any event: Haag's Theorem says "No!" for realistic (interacting) RQFTs.

*Haag's Theorem*  $\Rightarrow$   $\left( \begin{array}{l} \text{Representations of the CCRs for both a} \\ \text{non-interacting and an interacting} \\ \text{RQFT cannot be constructed so that they} \\ \text{are unitarily equivalent at a given time.} \end{array} \right)$

- Free particle total number operators cannot be used in interacting RQFTs.
- No consistent method for constructing "interacting" total number operators.

# 1. NQFTs and Particles

**Claim 1:** Conditions (A) and (B) fail in RQFTs.

**Against (A) in RQFTs:**

- Separability Corollary (Streater & Wightman 2000): Let  $\mathcal{A}$  be a local algebra of operators associated with a bounded region  $\mathcal{O}$  of spacetime. If

- (i) the vacuum state is cyclic for  $\mathcal{A}$  ("local cyclicity");
- (ii)  $\mathcal{O}$  has non-trivial causal complement;
- (iii) relativistic local commutativity holds;

then the vacuum state is *separating* for  $\mathcal{A}$ .

For any  $A \in \mathcal{A}$ , if  $A\Omega = 0$ ,  
then  $A = 0$ .

- Reeh-Schlieder theorem secures (i) for Minkowski spacetime.
- Structure of Minkowski spacetime secures (ii).
- RQFTs satisfy (iii).
- Thus: Annihilation operators, hence number operators, cannot be defined in  $\mathcal{A}$  for RQFTs in Minkowski spacetime.

# ■ 1. NQFTs and Particles

*To what extent does the Separability Corollary hold for RQFTs in Lorentzian spacetimes in general?*

- Local cyclicity holds for RQFTs in ultrastatic and stationary Lorentzian spacetimes (Verch 1993, Bar 2000, Strohmaier 1999, 2000).

As soon as a classical field satisfies a certain hyperbolic partial differential equation, a state over the field algebra of the quantized theory, which is a ground- or KMS-state with respect to the group of time translations, has the Reeh-Schlieder property [*i.e.*, local cyclicity]. (Strohmaier 2000, pg. 106.)

- Is local cyclicity a generic feature of globally hyperbolic Lorentzian spacetimes?
- If so, then local cyclicity is not a generic feature of RQFTs in Lorentzian spacetimes:
  - Global hyperbolicity is not a necessary condition for the existence of an RQFT in a Lorentzian spacetime. (Fewster and Higuchi 1996.)

# ■ 1. NQFTs and Particles

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- Is local cyclicity a generic feature of states analytic in the energy?
- Perhaps for RQFTs in Lorentzian spacetimes, but not for NQFTs in classical spacetimes:
  - Vacuum states for NQFTs are analytic but not locally cyclic for local algebras defined on spatial regions.

## ■ 1. NQFTs and Particles

*Claim 2:* Conditions (A) and (B) hold in NQFTs due to the absolute temporal metric of classical spacetimes.

### *Condition (A) in NQFTs:*

- *Non-relativistic local commutivity*  $\Rightarrow$  distinction between spatiotemporal local algebras and spatial local algebras.
- For spatiotemporal local algebra:
  - Requardt (1982)  $\Rightarrow$  Vacuum is locally cyclic.
  - *But:* Absolute temporal structure  $\Rightarrow$  Causal complement of  $\mathcal{O}$  is trivial.
  - *Hence:* Vacuum is not separating.
- For spatial local algebras:
  - No local cyclicity result.
  - *Hence:* Vacuum is not separating.

# 1. NQFTs and Particles

*Why does local cyclicity fail for local algebras associated with spatial regions of a classical spacetime?*

- Let  $\phi(t, \mathbf{x})$  be a positive-frequency solution to a well-posed PDE.
  - $\phi(t, \mathbf{x})$  is a boundary value of a holomorphic function.
- Let  $\mathcal{S}$  be an open spatial region of spacetime.
  - If  $\phi(t, \mathbf{x})$  vanishes on  $\mathcal{S}$ , then it vanishes in  $D(\mathcal{S})$ .
- Case 1: Hyperbolic PDE in Lorentzian spacetime.
  - $D(\mathcal{S})$  has non-zero temporal extent.
  - If  $\phi$  vanishes on  $\mathcal{S}$ , then it vanishes in an open set in time, and thus everywhere (Edge of the Wedge theorem).
  - Thus: If  $\phi \neq 0$ , then it cannot vanish on  $\mathcal{S}$ . *Anti-locality for spatial regions.*  
Segal and Goodman (1965)
- Case 2: Parabolic PDE in classical spacetime.
  - $D(\mathcal{S})$  has zero temporal extent.
  - If  $\phi$  vanishes on  $\mathcal{S}$ , then it need not vanish in an open set in time.
  - Thus: If  $\phi \neq 0$ , then it can vanish on  $\mathcal{S}$ . *Anti-locality fails for spatial regions.*

## ■ 1. NQFTs and Particles

*Claim 2:* Conditions (A) and (B) hold in NQFTs due to the absolute temporal metric of classical spacetimes.

*Condition (B) in NQFTs:*

- *No Problem of Privilege:* The absolute temporal metric guarantees a unique global time function on the spacetime, and this guarantees a unique means to construct a one-particle structure over the classical phase space (barring topological mutants).

## ■ 1. NQFTs and Particles

### *General Moral:*

To the extent that Conditions (A) and (B) require the existence of an absolute temporal metric, they are informed by a non-relativistic concept of time, and thus are inappropriate in informing interpretations of RQFTs.

## ■ 2. Newtonian Quantum Gravity

I. Theories of Newtonian Gravity (NG) with a grav. potential field  $\Phi$ .

$$(M, h^{ab}, t_{ab}, \nabla_a, \Phi, \rho)$$

$$h^{ab}t_{ab} = 0 = \nabla_c h^{ab} = \nabla_c t_{ab} \quad \textit{Orthogonality/compatibility}$$

$$h^{ab}\nabla_a\nabla_b\Phi = 4\pi G\rho \quad \textit{Poisson equation}$$

$$\xi^a\nabla_a\xi^b = -h_{ab}\nabla_a\Phi \quad \textit{Equation of motion}$$

**Ex. 1:** Neo-Newtonian NG

$$R^a{}_{bcd} = 0$$

**Ex. 2:** "Island Universe" Neo-Newtonian NG

$$R^a{}_{bcd} = 0, \quad \Phi \rightarrow 0 \text{ as } x^i \rightarrow \infty$$

**Ex. 3:** Maxwellian NG

$$R^{ab}{}_{cd} = 0$$

## 2. Newtonian Quantum Gravity

II. Theories of Newton-Cartan Gravity (NCG) that subsume  $\Phi$  into connection.  $(M, h^{ab}, t_{ab}, \nabla_a, \rho)$

$$h^{ab}t_{ab} = 0 = \nabla_c h^{ab} = \nabla_a t_{ab} \quad \textit{Orthogonality/compatibility}$$

$$R_{ab} = 4\pi G\rho t_{ab} \quad \textit{Generalized Poisson equation}$$

$$\xi^a \nabla_a \xi^b = 0 \quad \textit{Equation of motion}$$

**Ex. 1:** Weak NCG ( $1/c \rightarrow 0$  limit of GR)

$$R^{[a}_{[b}{}^{c]}_{d]} = 0$$

**Ex. 2:** Asymptotically spatially flat weak NCG (recovers Poisson equ.)

$$R^{[a}_{[b}{}^{c]}_{d]} = 0, \quad R^{abcd} = 0 \text{ at spatial infinity}$$

**Ex. 3:** Strong NCG (recovers Poisson equ.)

$$R^{[a}_{[b}{}^{c]}_{d]} = 0, \quad R^{ab}{}_{cd} = 0$$

## ■ 2. Newtonian Quantum Gravity

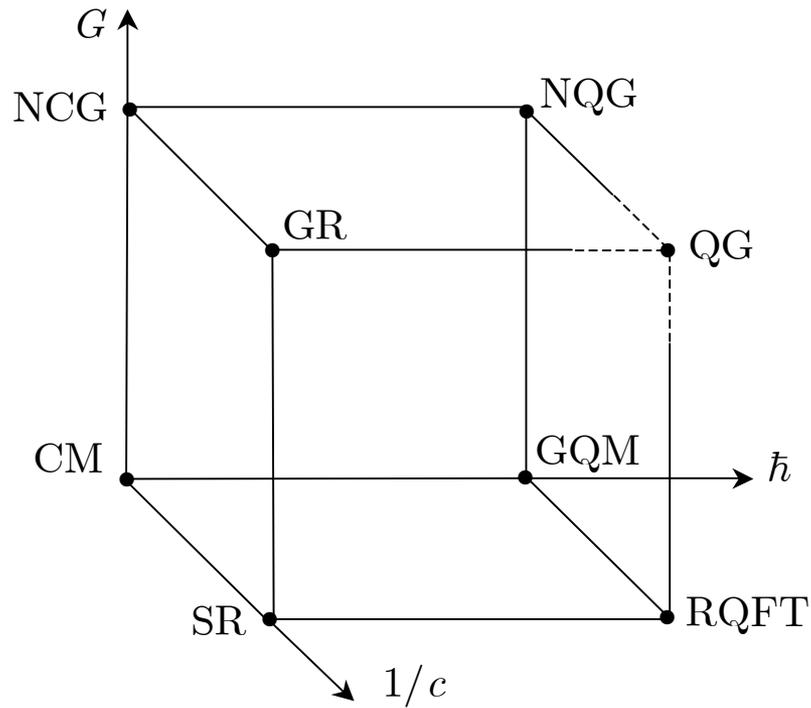
### *Strong NCG*

- Christian (1997): constrained Hamiltonian system, reduced phase space.
- Unique one-parameter family of time evolution maps  $\Rightarrow$  Unique Fock space quantization

### *Newtonian Quantum Gravity (NQG)*

- Interacting (extended) Maxwell-invariant QFT of gravity in curved classical spacetime ("strong Newton-Cartan" spacetime).
- Satisfies Conditions (A) and (B).
- Gravitational degrees of freedom are *dynamic*: Compare with RQFTs in curved Lorentzian spacetimes.
- Gravitational degrees of freedom are *quantized*: Compare with semi-classical quantum gravity.

### 3. Intertheoretic Relations

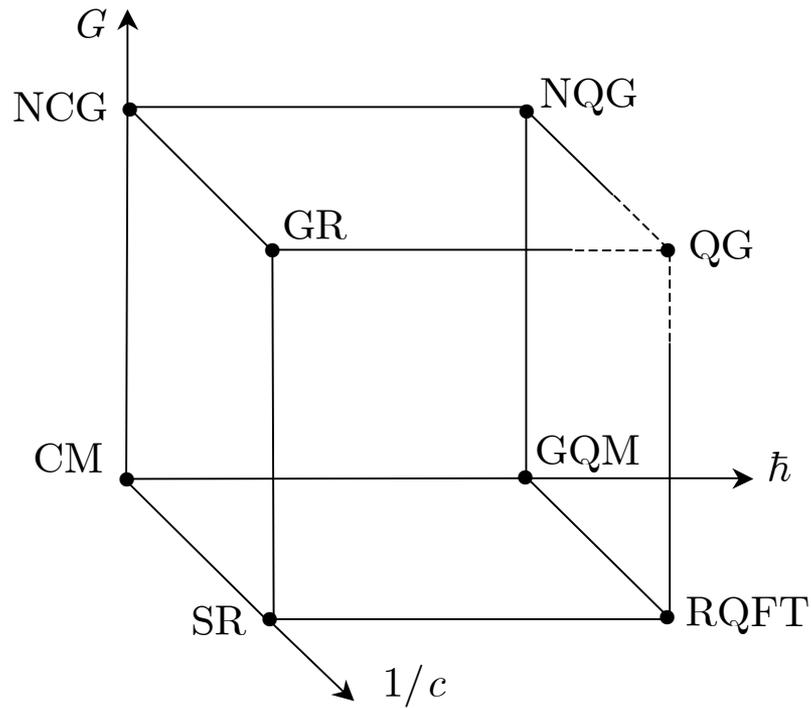


Christian (1997)

#### $1/c \rightarrow 0$ limit

- Contraction of Poincaré Group? (Bacry & Levy-Leblond 1968)
- SR  $\rightarrow$  CM, RQFT  $\rightarrow$  GQM: Depends on dynamics. (Brown & Holland 2003)
- GR  $\rightarrow$  NCG: No.

### 3. Intertheoretic Relations

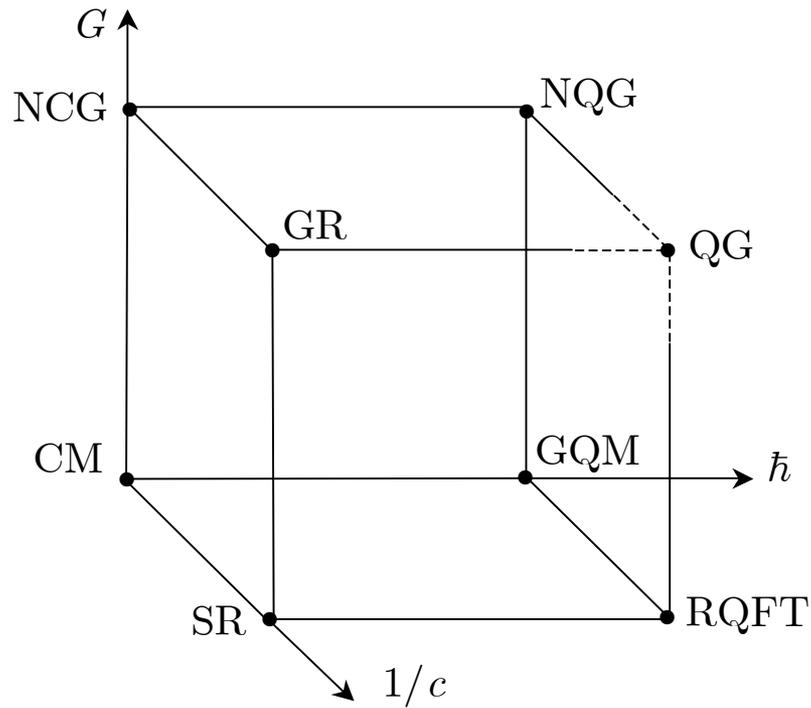


Christian (1997)

#### $G \rightarrow 0$ limit: Ricci vs Riemann flatness

- GR  $\rightarrow$  SR: Vacuum Einstein spacetime vs Minkowski spacetime
- NCG  $\rightarrow$  CM, NQG  $\rightarrow$  GQM: Ricc-flat classical spacetime vs Neo-Newtonian spacetime

### 3. Intertheoretic Relations

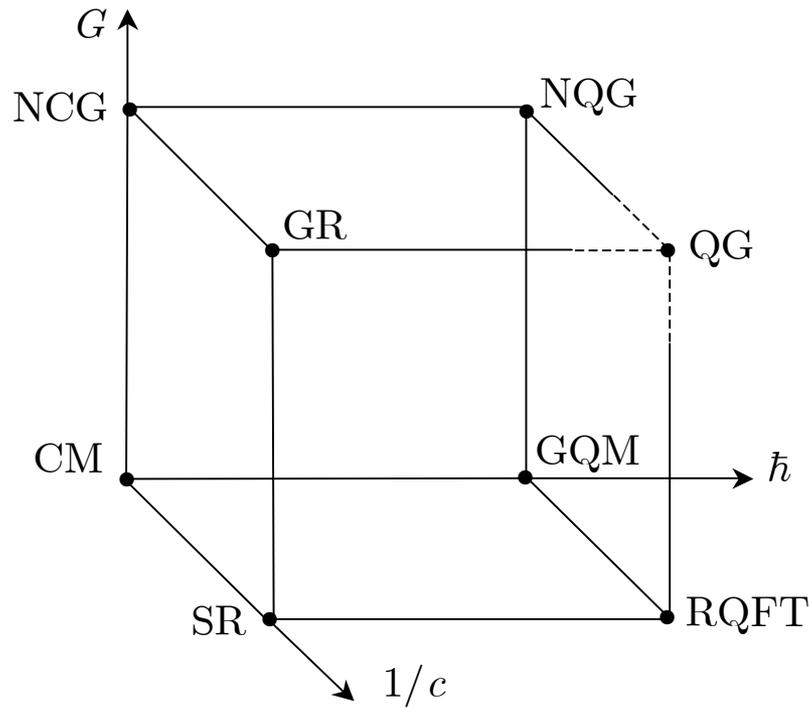


Christian (1997)

#### $\hbar \rightarrow 0$ limit: Problem of Privilege

- RQFT  $\rightarrow$  SR: No unique (up to unitary equivalence) representation of CCRs.
- GQM  $\rightarrow$  CM, NQG  $\rightarrow$  NCG: No problem (barring topological mutants).

### 3. Intertheoretic Relations



Christian (1997)

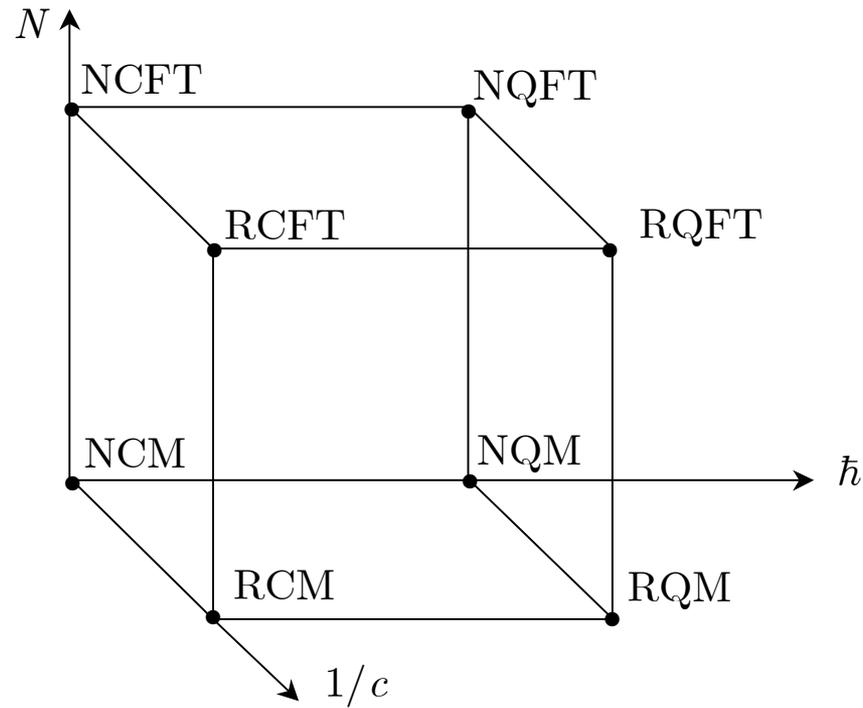
#### Structural Problem

- What is the referent of "GQM"? Where do NQFTs fit in?

#### Proposal: Add another axis for $N = \text{degrees of freedom}$

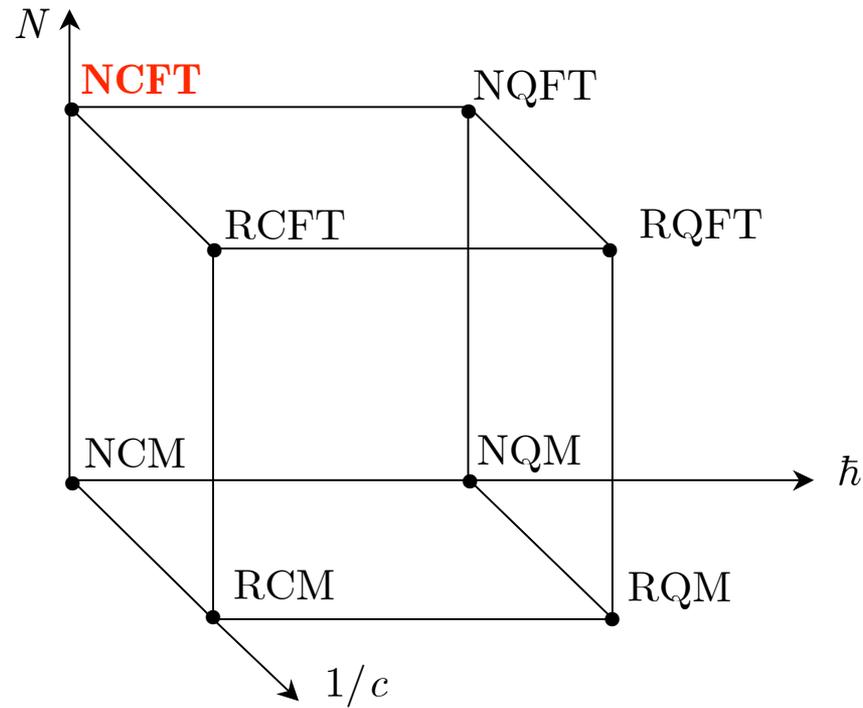
- Let "NQM" refer to non-relativistic finite-dimensional quantum theories of particle dynamics.
- Consider NQMs to be the  $N \rightarrow 0$  limit of NQFTs.

### 3. Intertheoretic Relations



- Particle *vs* field theories ( $N$  axis).
- Relativistic *vs* non-relativistic theories ( $1/c$  axis).
- Gravitational *vs* non-gravitational theories ( $G$  axis).
- Classical *vs* quantum theories ( $\hbar$  axis).

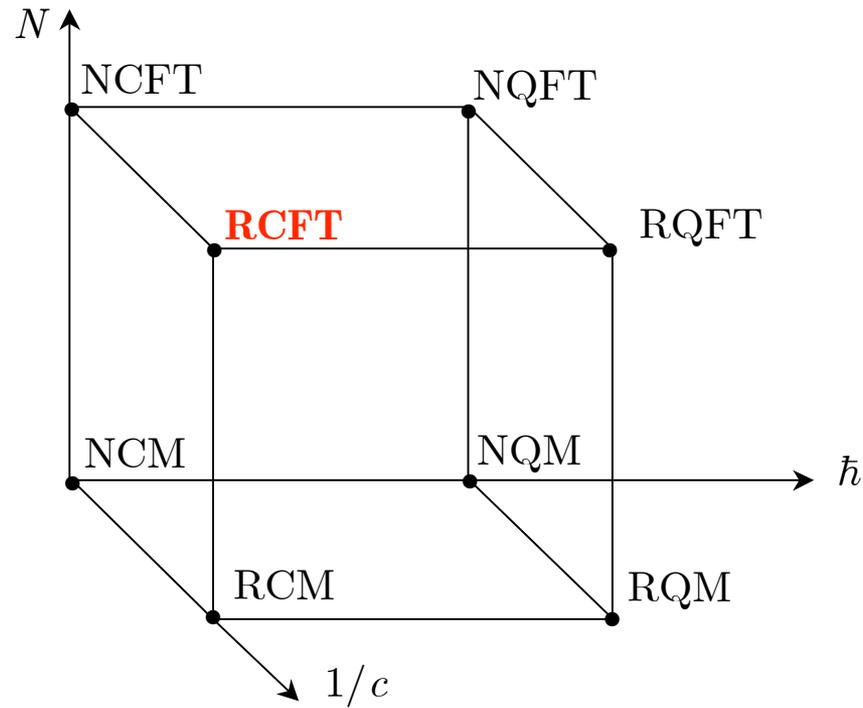
### 3. Intertheoretic Relations



#### Turning off $G$ in field theories:

- Non-relativistic classical field theory of gravity  $\rightarrow$  NCFT
- Asymptotically spatially flat NCG = "Island Universe" Neo-Newtonian NG
- $G \rightarrow 0$ : Galilei-invariant classical field theory in Neo-Newtonian spacetime

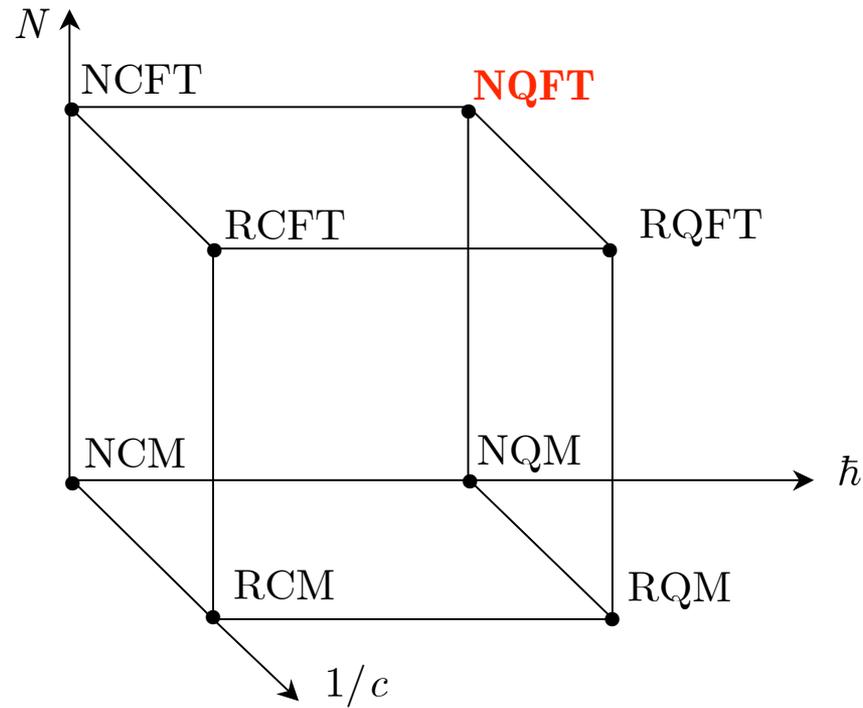
### 3. Intertheoretic Relations



#### Turning off $G$ in field theories:

- Relativistic classical field theory of gravity  $\rightarrow$  RCFT
- GR
- $G \rightarrow 0$ : Relativistic classical field theory in Ricci-flat Lorentzian spacetime

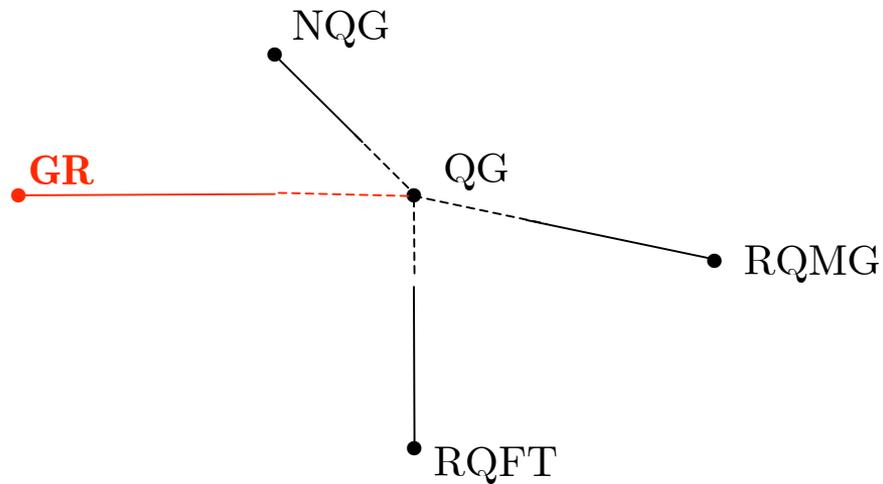
### 3. Intertheoretic Relations



#### Turning off $G$ in field theories:

- Non-relativistic quantum field theory of gravity  $\rightarrow$  **NQFT**
- **NQG**
- $G \rightarrow 0$ : **NQFT** in Ricci-flat classical spacetime

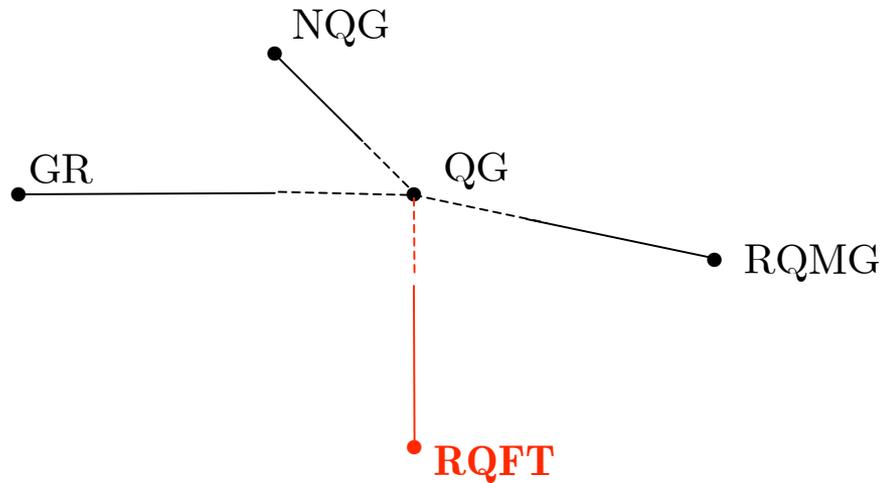
### 3. Intertheoretic Relations



Turning on quantum gravity:

- Quantizing GR.

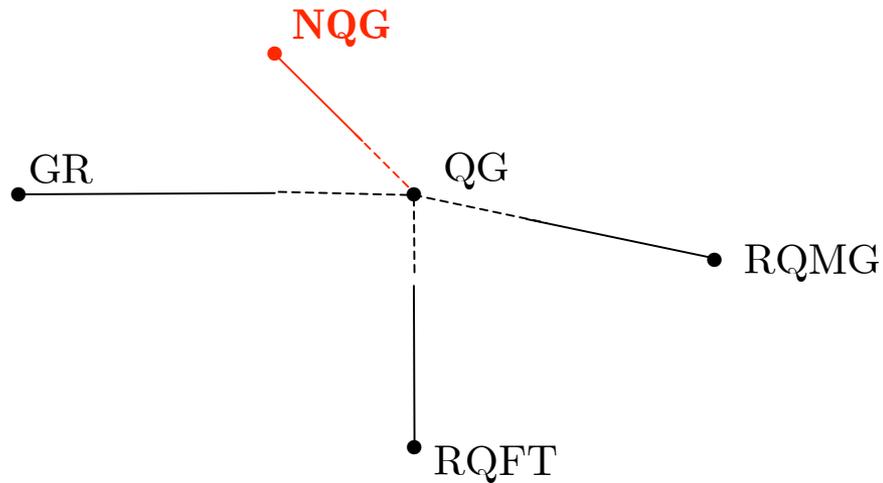
### 3. Intertheoretic Relations



#### Turning on quantum gravity:

- Quantizing GR.
- Turning on gravity in an RQFT.

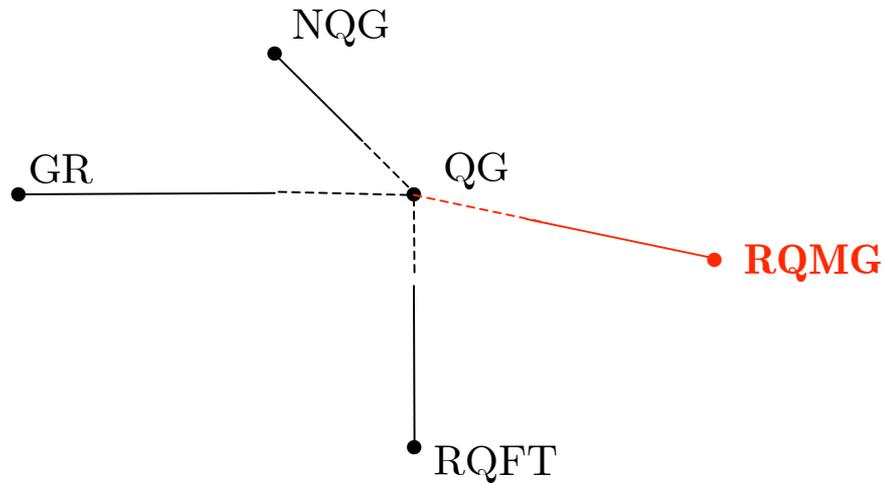
### 3. Intertheoretic Relations



#### Turning on quantum gravity:

- Quantizing GR.
- Turning on gravity in an RQFT.
- Relativizing NQG.

### 3. Intertheoretic Relations



#### Turning on quantum gravity:

- Quantizing GR.
- Turning on gravity in an RQFT.
- Relativizing NQG.
- Taking the "thermodynamic limit" of an RQMG.