

ASTR 105 The Solar System

Today: Group Lab at the end of class
Next THURSDAY 03/10: First Group Project

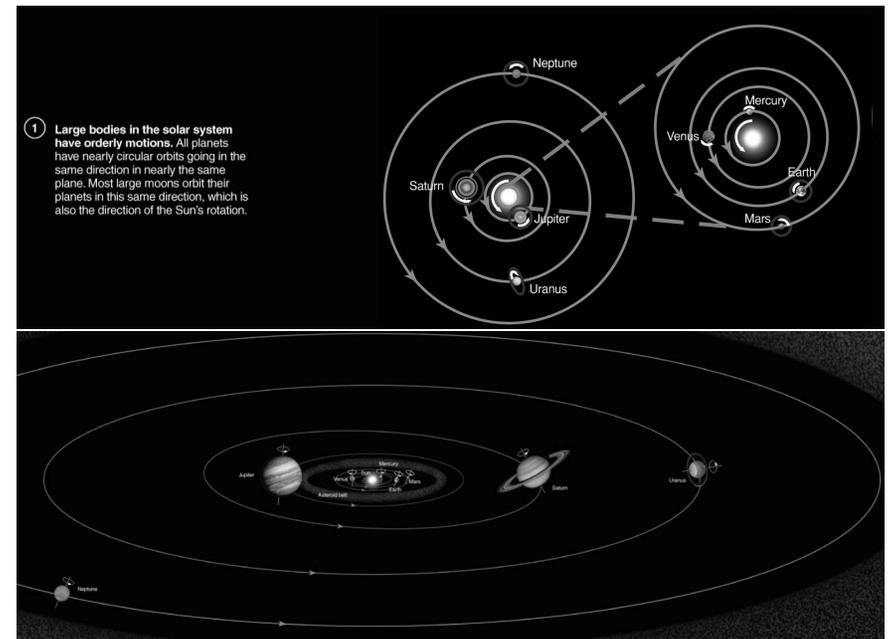
Patterns in the Solar System

- 1. Orderly motions***
- 2. Two kinds of planets***
- 3. Two kinds of small bodies***
- 4. Exceptions to the rules***

Patterns in the Solar System

Orderly Motions

- **Almost every body orbits & spins in the same direction.**
 - Counterclockwise from above
- **Planetary orbits nearly circular, lie in nearly the same plane.**
 - Large moons tend to exhibit the same properties.



Patterns in the Solar System

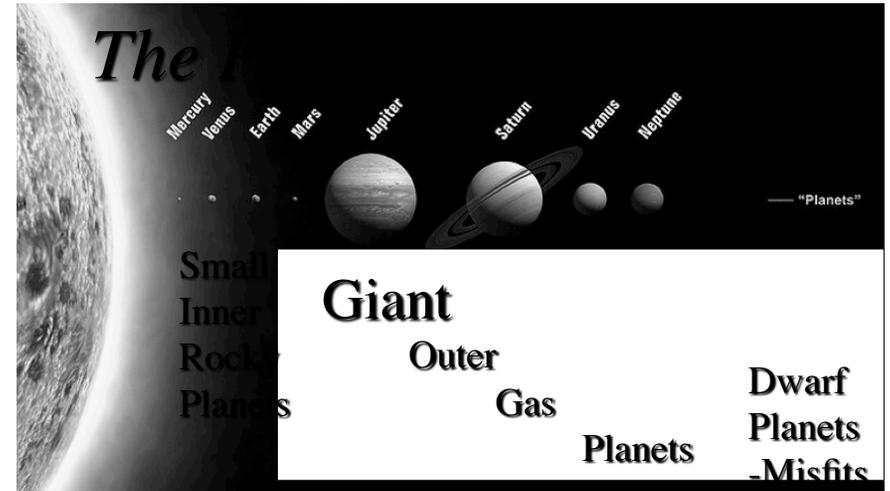
Two Types of Planets

Terrestrial Planets

- Smaller size and mass
- Mostly rock and metal
- High density
- Solid surface
- Few or no moons, no rings
- Closer to the Sun, closer together

Jovian Planets

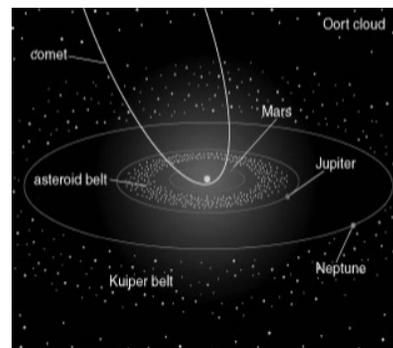
- Larger size and mass
- Mostly H, He, and H-compounds
- Low density
- No solid surface
- Rings and many moons
- Farther from the Sun, farther apart



Patterns in the Solar System

Two Types of Small Bodies

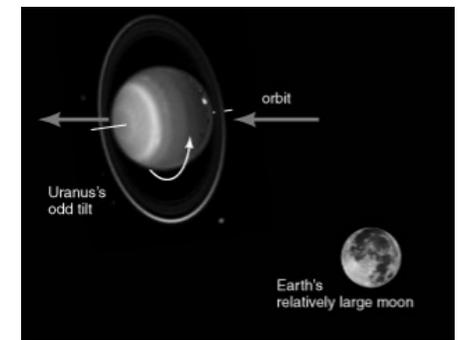
- **Rocky**
 - Asteroids
- **Ice and rock**
 - Comets
 - Kuiper belt objects
 - Oort cloud comets



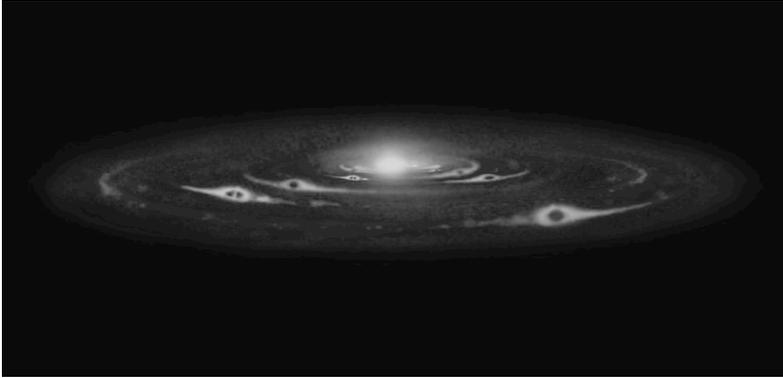
Patterns in the Solar System

Exceptions to the rules

- **Rotations**
 - Venus
 - Backwards
 - Uranus
 - On its side
- **Orbits**
 - Triton
 - Backwards
- **Moons**
 - Earth's Moon
 - Big (compared to Earth)
 - Most small moons
 - Screwy shapes & orbits



Next: Formation of the Solar System



Where did the elements in the solar nebula come from?

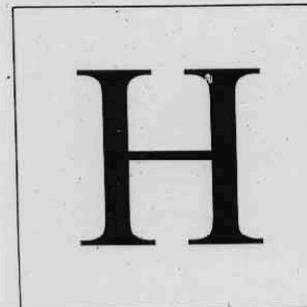


Periodic Table of the Elements

| | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
|--|----------|---|------------|------------|------------|------------|------------|------------|------------|-----------|-----------|-----------|-----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|----------|---------|----------|----------|----------|----------|----------|----------|----------|-----------|-----------|-----------|-----------|
| 1 H | | | | | | | | | | | | | | | | | 2 He | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 3 Li | 4 Be | <ul style="list-style-type: none"> ■ hydrogen ■ alkali metals ■ alkali earth metals ■ transition metals ■ poor metals ■ nonmetals ■ noble gases ■ rare earth metals | | | | | | | | | | 5 B | 6 C | 7 N | 8 O | 9 F | 10 Ne | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 11 Na | 12 Mg | | | | | | | | | | | 13 Al | 14 Si | 15 P | 16 S | 17 Cl | 18 Ar | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 19 K | 20 Ca | 21 Sc | 22 Ti | 23 V | 24 Cr | 25 Mn | 26 Fe | 27 Co | 28 Ni | 29 Cu | 30 Zn | 31 Ga | 32 Ge | 33 As | 34 Se | 35 Br | 36 Kr | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 37 Rb | 38 Sr | 39 Y | 40 Zr | 41 Nb | 42 Mo | 43 Tc | 44 Ru | 45 Rh | 46 Pd | 47 Ag | 48 Cd | 49 In | 50 Sn | 51 Sb | 52 Te | 53 I | 54 Xe | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 55 Cs | 56 Ba | 57 La | 72 Hf | 73 Ta | 74 W | 75 Re | 76 Os | 77 Ir | 78 Pt | 79 Au | 80 Hg | 81 Tl | 82 Pb | 83 Bi | 84 Po | 85 At | 86 Rn | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 87 Fr | 88 Ra | 89 Ac | 104 Unq | 105 Unp | 106 Unh | 107 Uns | 108 Uno | 109 Une | 110 Uun | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| <table border="1" style="width: 100%; border-collapse: collapse;"> <tr> <td>58 Ce</td><td>59 Pr</td><td>60 Nd</td><td>61 Pm</td><td>62 Sm</td><td>63 Eu</td><td>64 Gd</td><td>65 Tb</td><td>66 Dy</td><td>67 Ho</td><td>68 Er</td><td>69 Tm</td><td>70 Yb</td><td>71 Lu</td> </tr> <tr> <td>90 Th</td><td>91 Pa</td><td>92 U</td><td>93 Np</td><td>94 Pu</td><td>95 Am</td><td>96 Cm</td><td>97 Bk</td><td>98 Cf</td><td>99 Es</td><td>100 Fm</td><td>101 Md</td><td>102 No</td><td>103 Lr</td> </tr> </table> | | | | | | | | | | | | | | | | | | 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr |
| 58 Ce | 59 Pr | 60 Nd | 61 Pm | 62 Sm | 63 Eu | 64 Gd | 65 Tb | 66 Dy | 67 Ho | 68 Er | 69 Tm | 70 Yb | 71 Lu | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | |
| 90 Th | 91 Pa | 92 U | 93 Np | 94 Pu | 95 Am | 96 Cm | 97 Bk | 98 Cf | 99 Es | 100 Fm | 101 Md | 102 No | 103 Lr | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | | |

Right after the Big Bang

The Astronomer's Periodic Table
(Ben McCall)

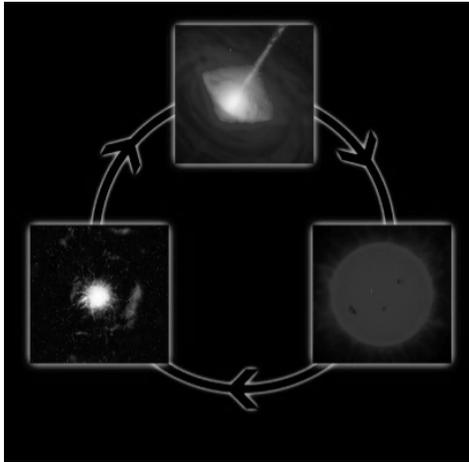


• **Mostly**
Hydrogen
(75%)

• **Some** Helium
(25%)

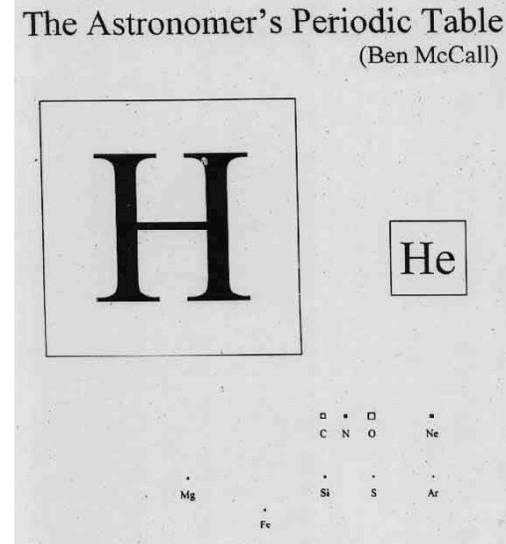
• Virtually nothing else

Galactic Recycling



- **H & He**
 - created in the Big Bang
- **Heavier elements**
 - made in stars and then recycled through interstellar space

Today



• **Almost entirely Hydrogen (70%)**

• **Some Helium (28%)**

• **Tiny amount of everything else (2%)**

Clicker Question

What would the "Astronomer's Periodic Table" look like if we looked at the universe in 13 billion years?

- It would look exactly the same
- The H would be a little smaller but not much, other elements a little bigger
- The H and He would be pretty much gone, all converted to heavier elements
- All the elements would be the same size (I.e. the same amount of everything)

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13 Billion Years From Today

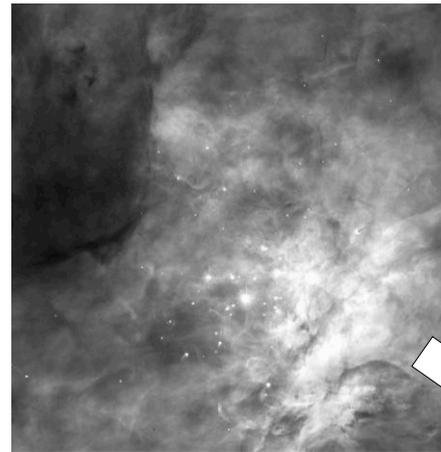
The Astronomer's Periodic Table
(Ben McCall)

| | | | |
|----|----|---|----|
| H | He | | |
| C | N | O | Ne |
| Mg | Si | S | Ar |
| Fe | | | |

- Still mostly Hydrogen (65%)

- A bit more Helium (31%)

- A little more of everything else (4%)



NEBULAR FORMATION
Theory
for the Solar System

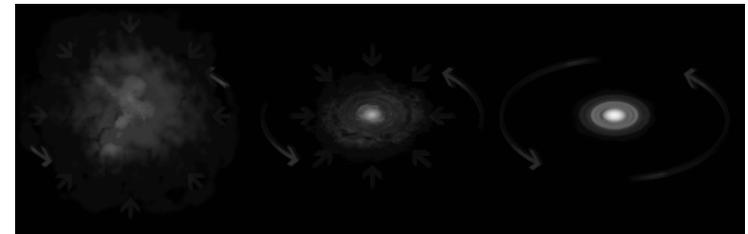
How does a solar system form from a cloud of gas?



Four Challenges for a Solar System Formation Theory

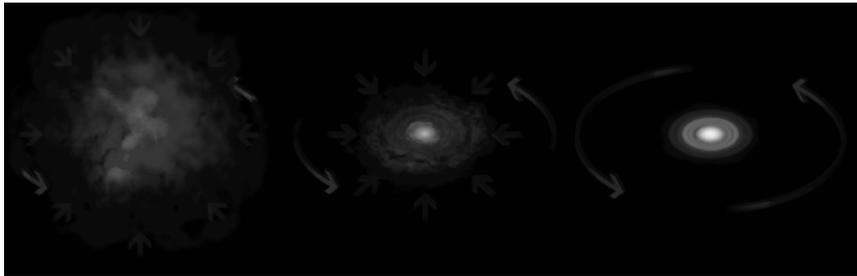
1. *Orderly motions*
2. *Two kinds of planets*
3. *Two kinds of small bodies*
4. *Exceptions to the rules*

Collapse of the Solar Nebula

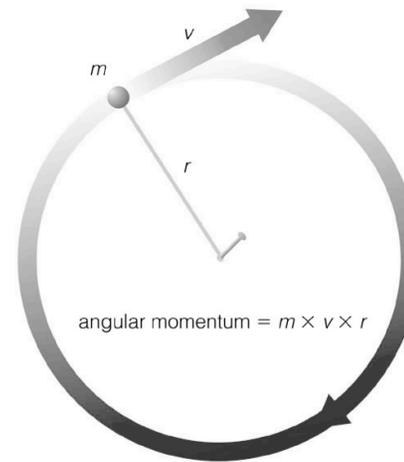


- As the solar nebula collapses it:
 - Spins faster
 - Heats up
 - Flattens out into a disk

Collapse of the Solar Nebula



Why does it spin faster?



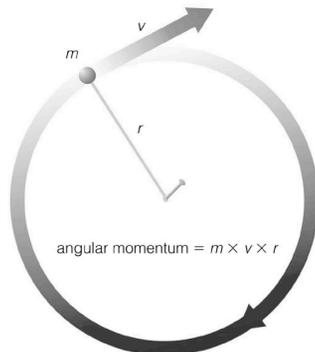
**Conservation of
Angular
Momentum**

$$M \times V \times R = \text{Constant}$$

Clicker Question

If angular momentum ($m \times v \times r$) is conserved (stays constant) what happens if r goes down (no change in m)?

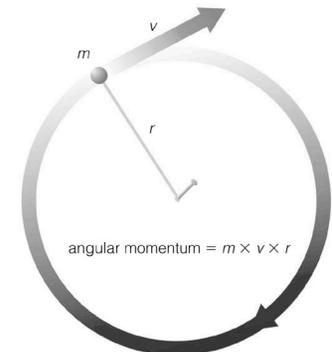
- A. v stays the same**
- B. v goes down**
- C. v goes up**

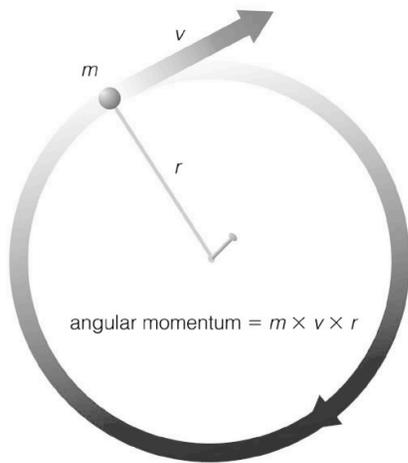


Clicker Question

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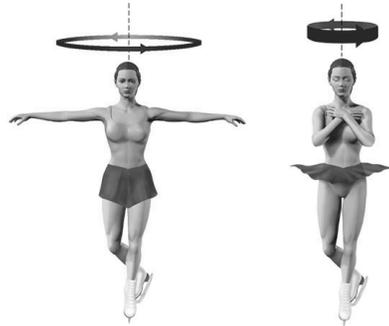
- A. v stays the same**
- B. v goes down**
- C. v goes up**





Conservation of Angular Momentum

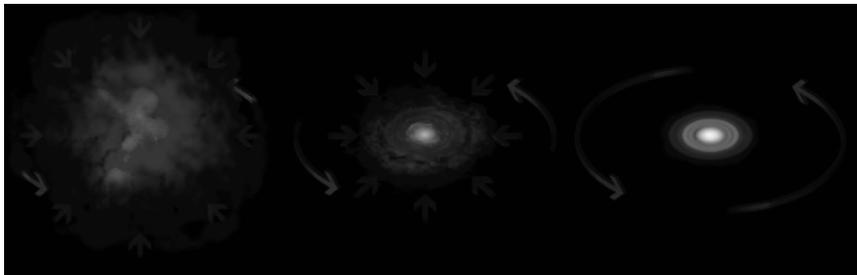
$$M \times V \times R = \text{Constant}$$



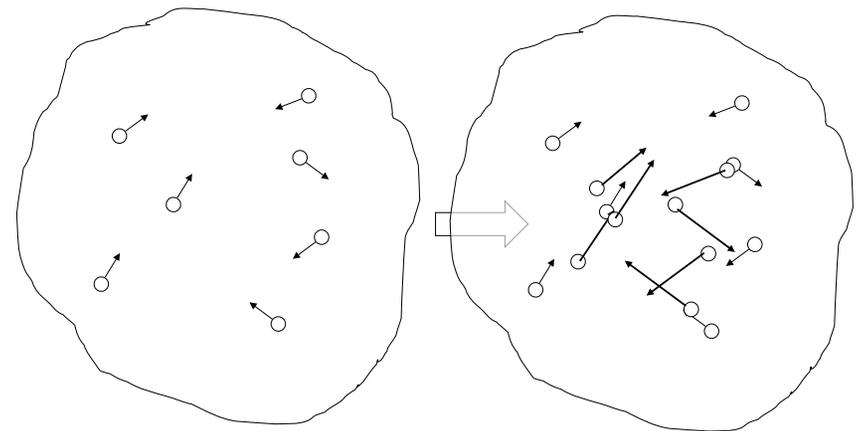
Spinning Demos

<https://www.youtube.com/watch?v=UZIW1a63KZs>

Collapse of the Solar Nebula

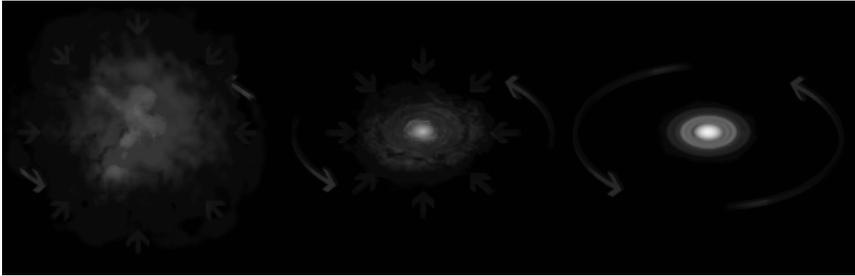


Why does it heat up?



Gravitational Energy \Rightarrow Kinetic Energy
Kinetic Energy = Thermal Energy

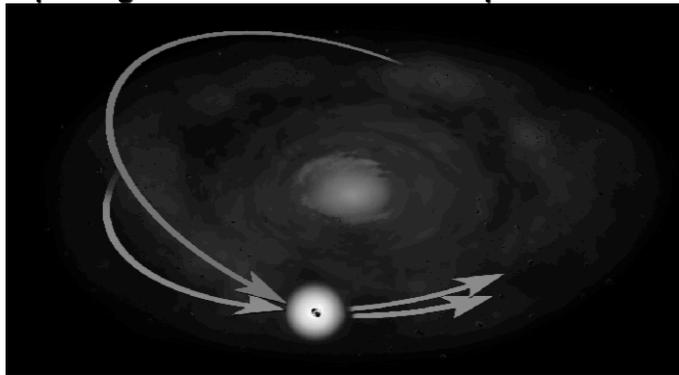
Collapse of the Solar Nebula



Why does it flatten into a disk?

Flattening of the Solar Nebula

- As the nebula collapses, clumps of gas collide & merge.
- Their random velocities average out into the nebula's direction of rotation.
- The spinning nebula assumes the shape of a disk.



Incorrect Reasons

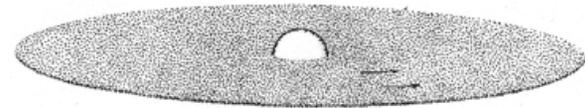
Gravity is "pulling the material into a disk"

- Gravity is pulling everything in (spherically, not into a flat disk)

OR

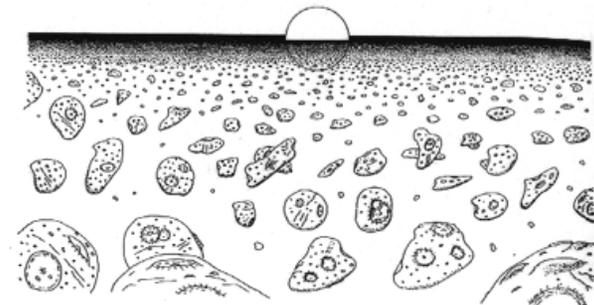
The disk is being "flung out into a disk"

- Individual gas particles are in simple orbits (orbits don't get flung out)



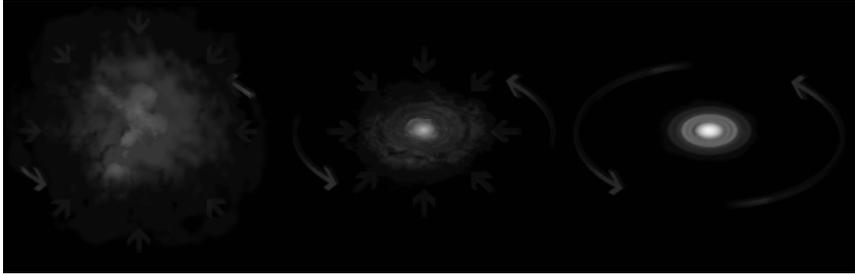
Extreme Conformism!

Go with the flow or crash to oblivion



Solar Nebula:

- SPINNING
 - Conservation of angular momentum
- HOT
 - Collapse \Rightarrow compression
- DISK
 - Collisions force common motions



- a) The orbital period of the Moon is 29.5 days, and its distance from Earth is 238,900 miles. What is the mass of the Earth? LAB
- b) The New Horizon spacecraft will take 9 years to travel to Pluto ($D=7.5 \times 10^9$ km). What is its average speed?
- c) Uranus's orbit lasts 84 years. If you live at its South pole, for roughly how long would you see continuous day light?
- d) A new planet is discovered, orbiting a star of mass $M=2 \times 10^{30}$ kg in 5.7 years. At what distance from the star is the planet orbiting?
- e) List at least 3 differences between Pluto and the other solar planets. What feature did astronomers agree upon as a way to differentiate a 'planet' from a 'dwarf planet'?