

Performance Modeling of the NIF Neutron Imaging System

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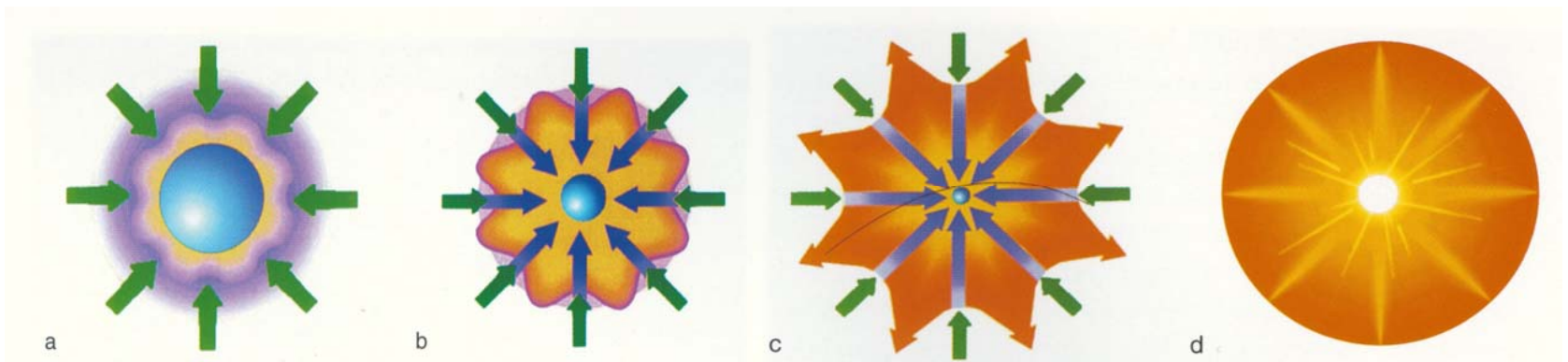
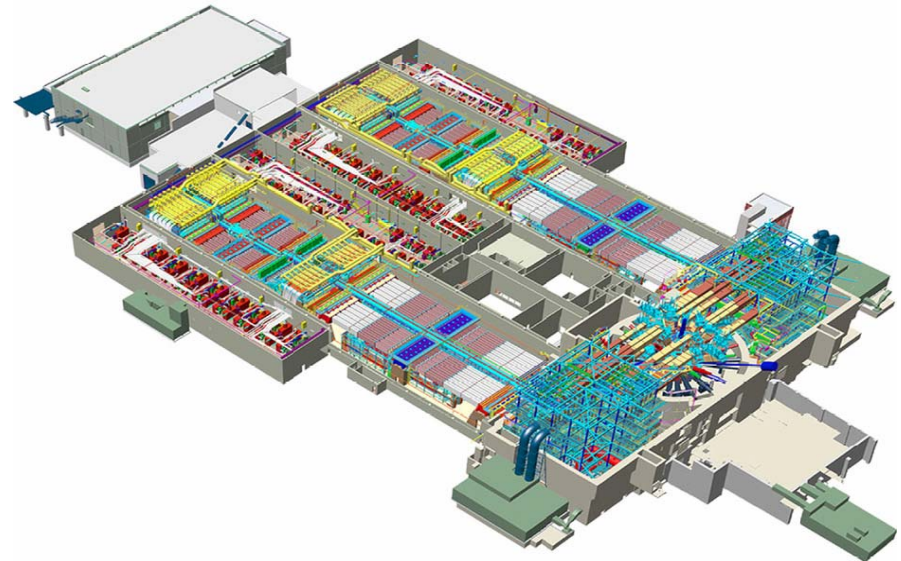
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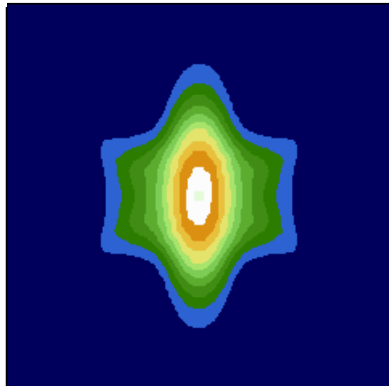
This work was performed under the auspices of the U.S. Department of Energy by the University of California, Lawrence Livermore National Laboratory under contract No. W-7405-Eng-48.

UCRL-PRES-226123

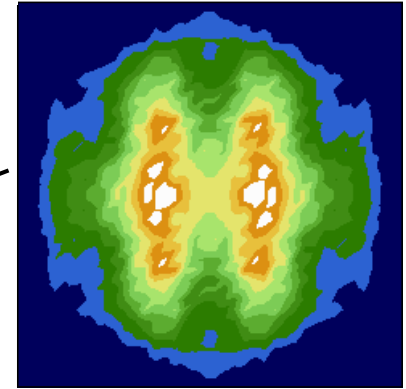
National Ignition Facility and ICF concept

- NIF is under construction
 - 85 % complete
 - Ignition Campaign in FY09
- Inertial Confinement Fusion
 - a. Laser/X-ray illumination
 - b. Capsule compression
 - c. Fusion ignition
 - d. Fusion burn

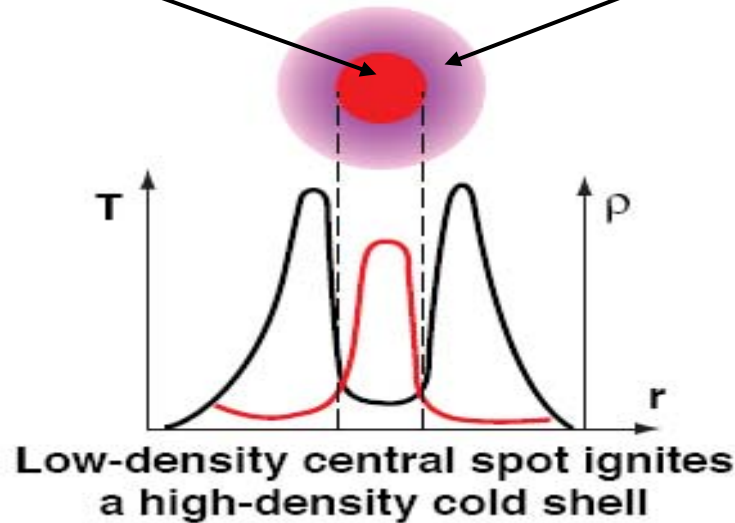




Primary*



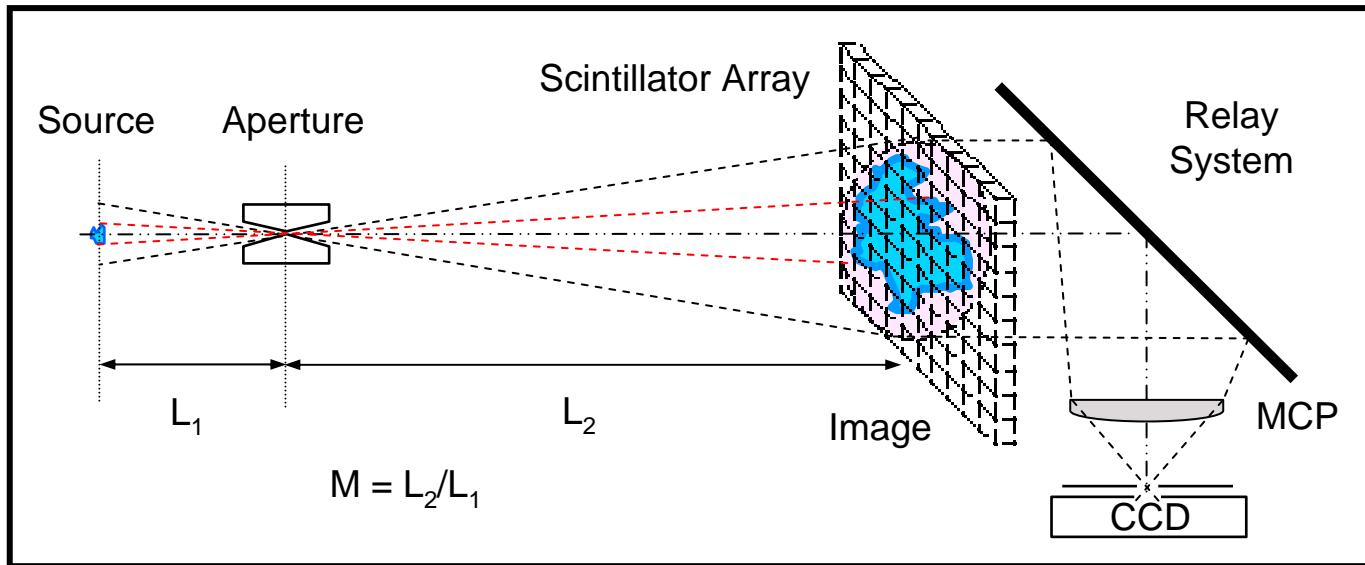
Downscattered*



- 14-MeV or primary neutrons are related to the central “Hot Spot”
- Downscattered neutrons (6 – 10 MeV) are related to the surrounding cold fuel
- Downscattered neutrons reveal the actual fuel distribution

* Simulated ICF images by S. Hatchett

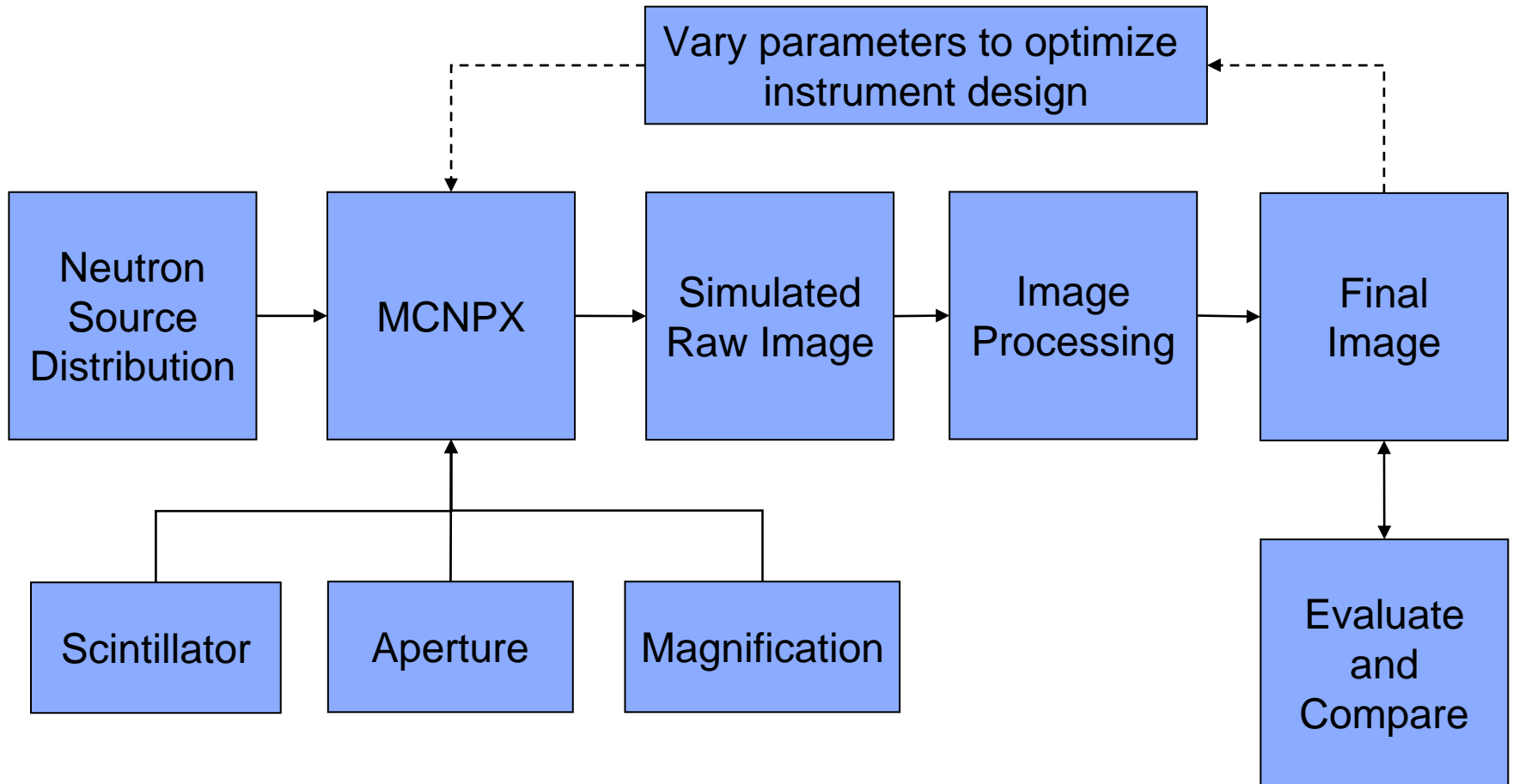
Neutron imaging concept: a typical line of sight



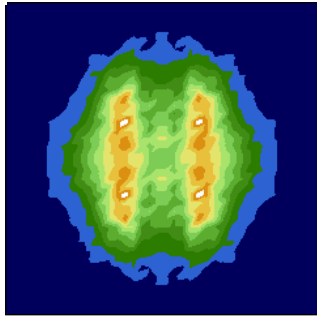
Typical NIF parameters:

- $L_1 = 40$ cm
- $L_2 = 40$ m
- $M = 100$
- Pinhole Material: Tungsten
- Pinhole Thickness = 13.7 cm
- Field of View (FoV) = 200 μm
- Resolution = 10 μm FWHM

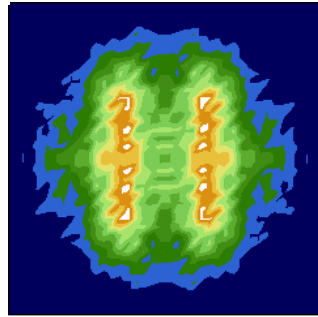
Overall system model being developed for neutron imaging design studies



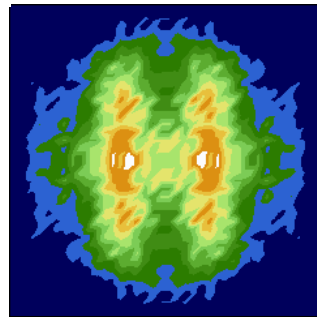
Starting point: Detailed source maps



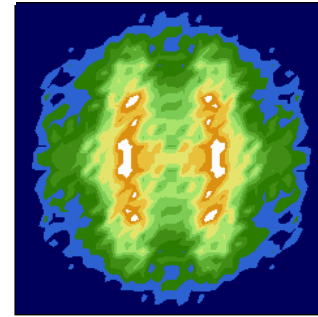
11 to 12 MeV



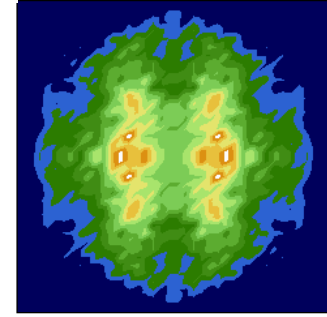
10 to 11 MeV



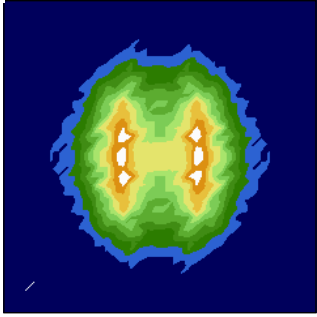
9 to 10 MeV



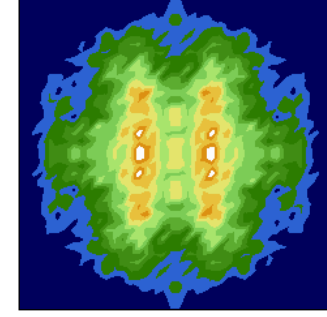
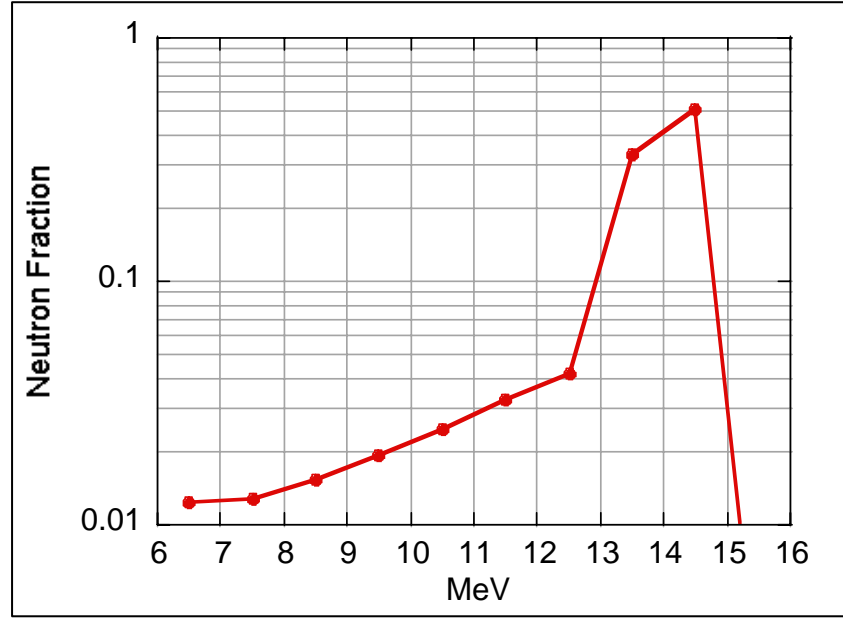
8 to 9 MeV



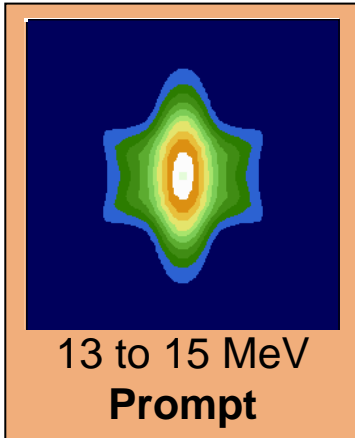
7 to 8 MeV



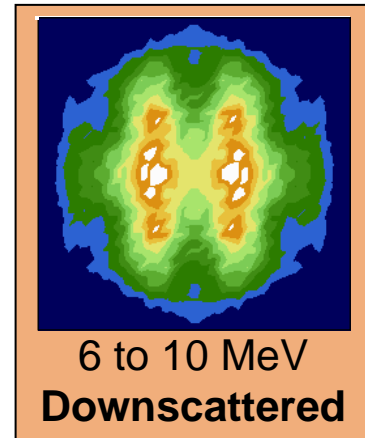
12 to 13 MeV



6 to 7 MeV



13 to 15 MeV
Prompt

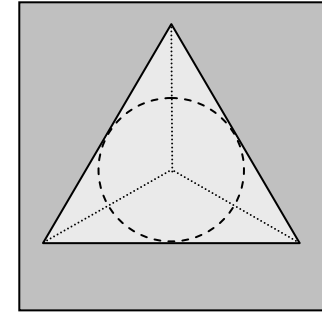
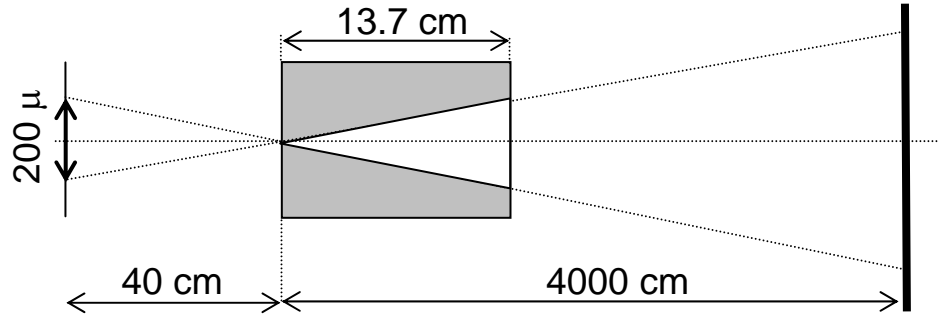


6 to 10 MeV
Downscattered

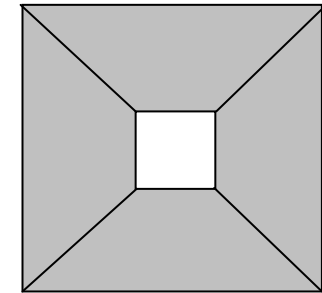
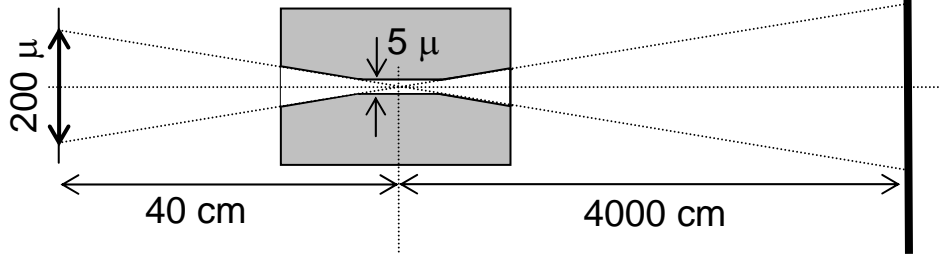
- S. Hatchett ICF calculation of a failed implosion
- 100 micron FoV maps with 2.5 micron pixels
- P6 drive perturbation
- Yield 16 kJ (7E15 13-15MeV, 5E14 6-10 MeV)

Aperture Geometry: 10 x 10 Arrays

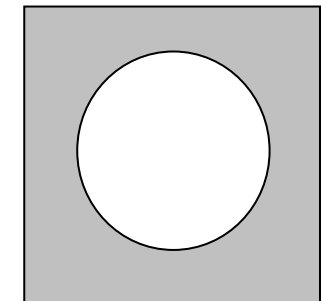
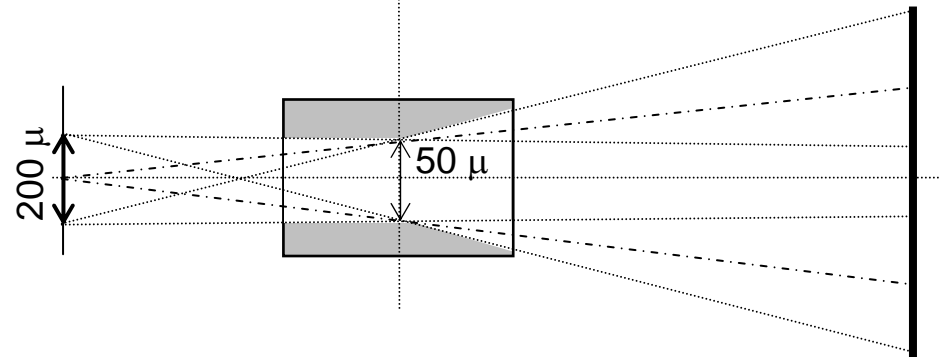
Triangular Wedge



Square Pinhole

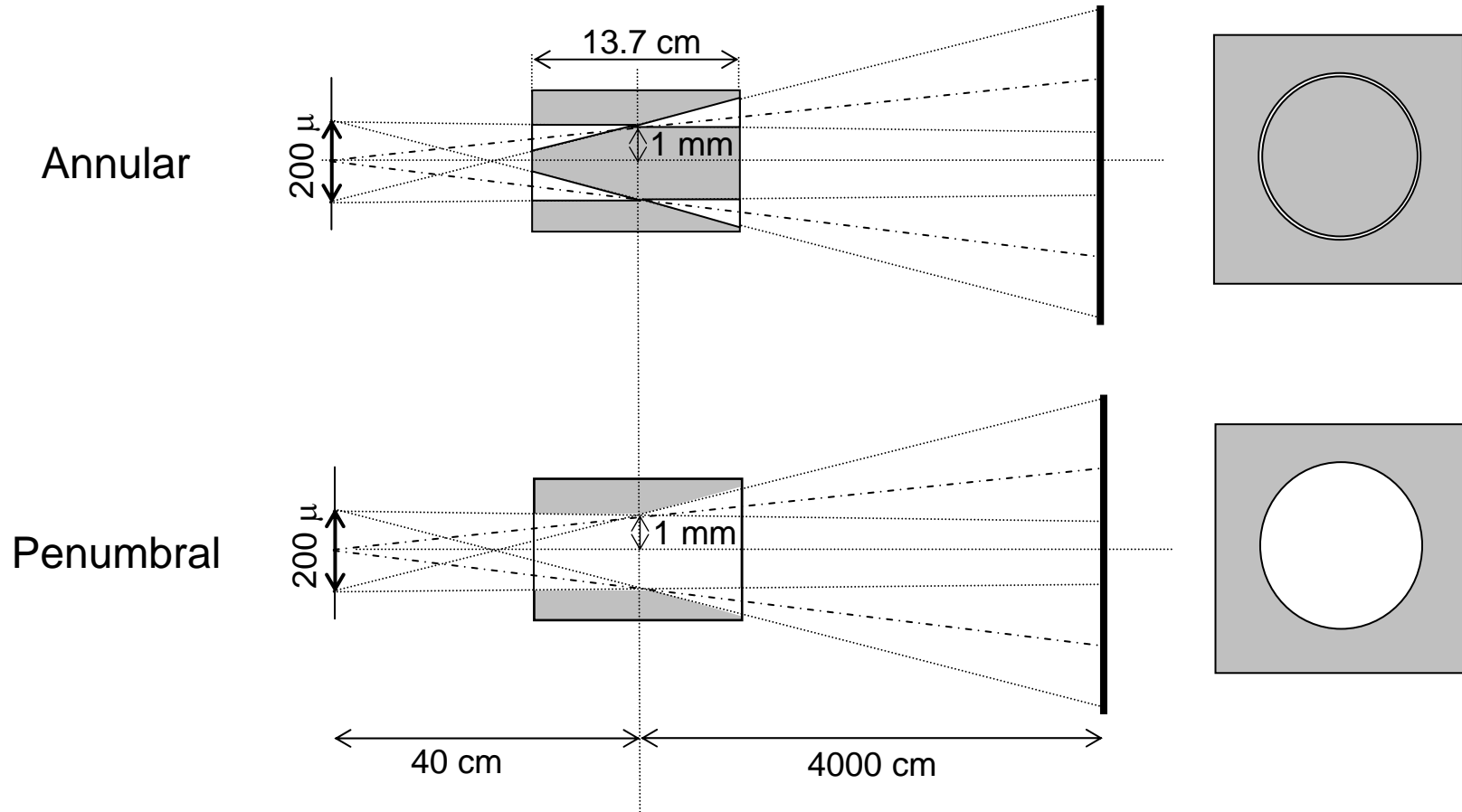


Mini Penumbra

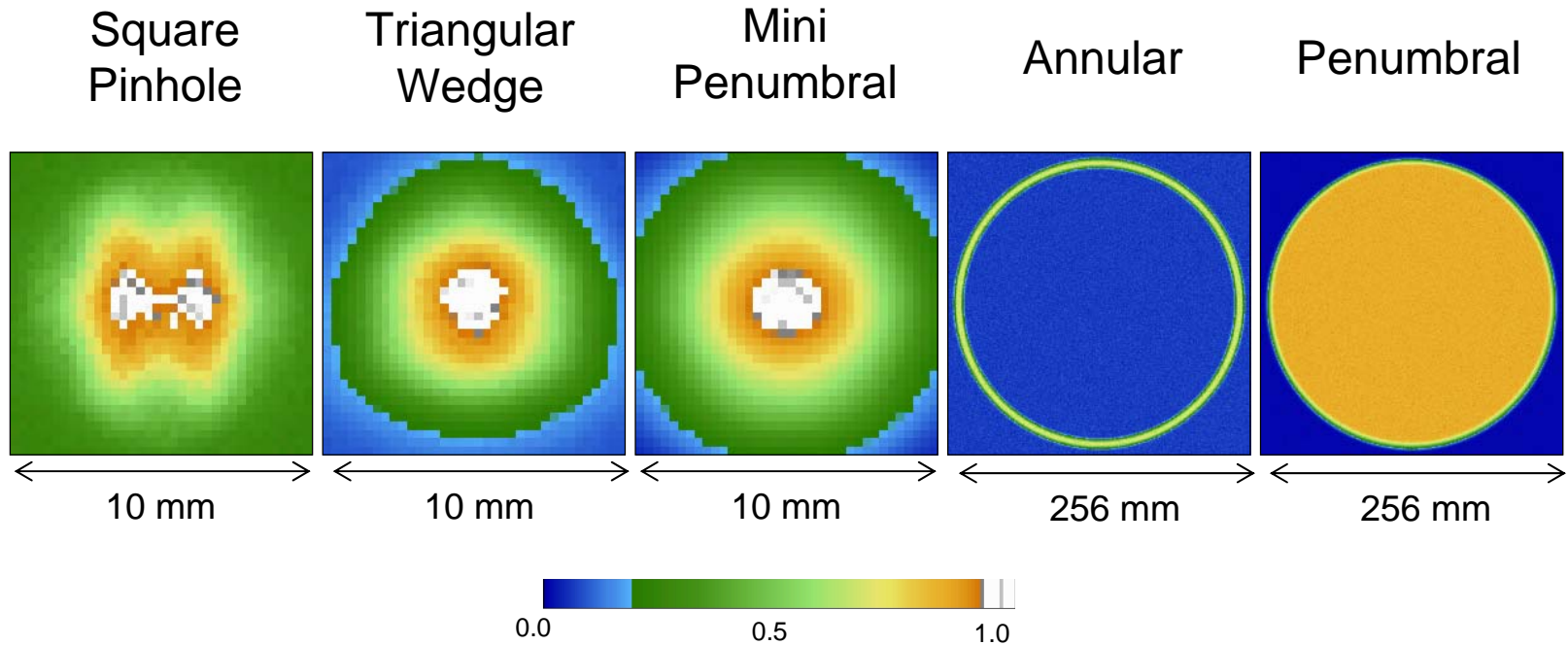


Drawings not to scale

Aperture Geometry: Single Aperture

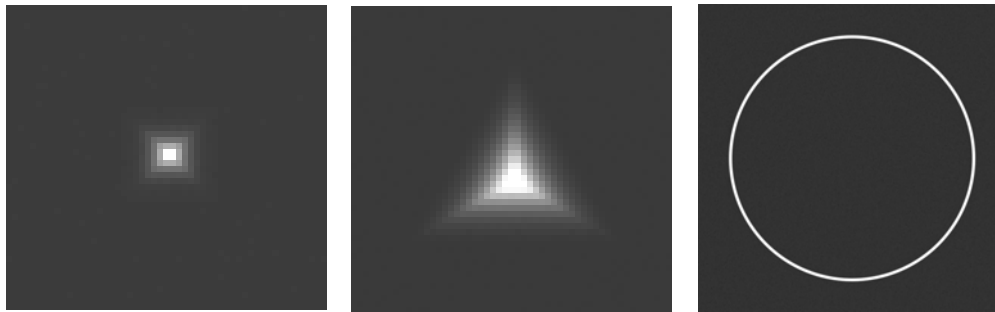


Drawings not to scale



- Detector Characteristics
 - BC422 Scintillator
 - Pixel geometry: 250 micron square, 4.6 cm long
- MCNPX calculation reflects aperture and scintillator Point Spread Functions (PSF)
 - No mathematical convolution
- Image arrays in proton energy deposited (MeV) per pixel
- Pinhole arrays are required due to single aperture small solid angle fraction

Accurate Aperture PSFs via MCNPX



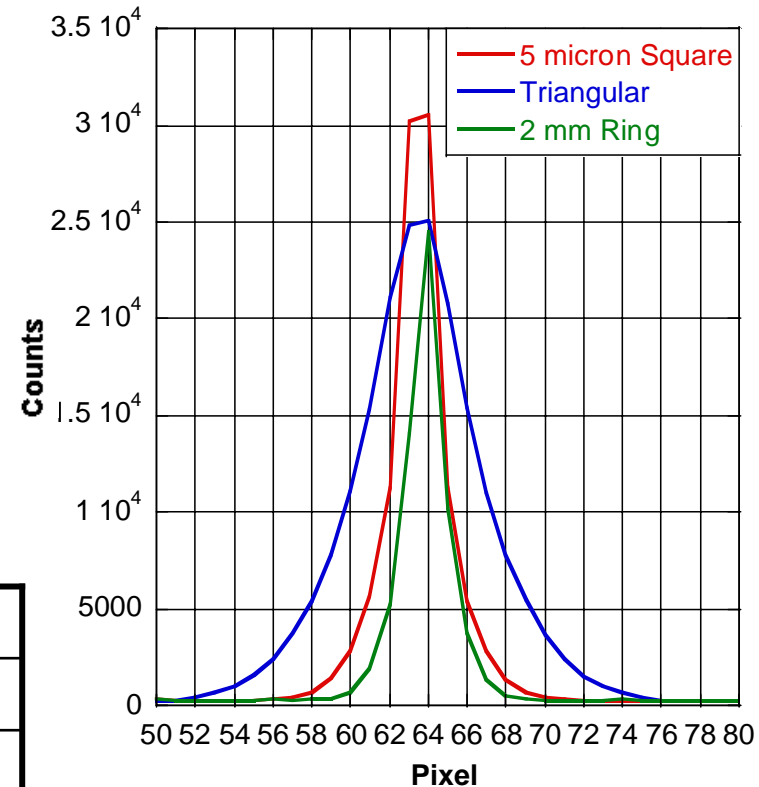
Square

Triangular

Ring

- 10^{16} neutrons equivalent source
- Same size as expected image
- Not opaque (direct transmission)
- Single on-axis (Isoplanatic approximation)

	Square	Triangular	Ring
FWHM (μ)	7.5	15	5
$\Delta\Omega_{\text{eff}}^*$	2.517×10^{-11}	1.771×10^{-10}	4.424×10^{-8}
$\phi_{\text{eff}} (\mu)^{**}$	8.02	24.91	336.55



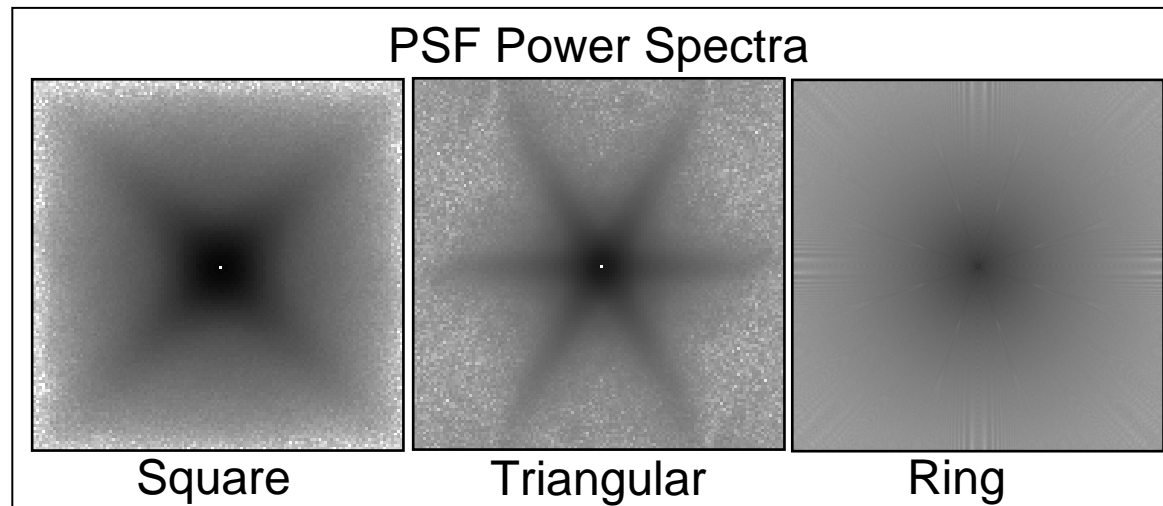
* Effective solid angle fraction: based on actual neutron flux inside bright PSF region

** Effective aperture diameter: corresponds to the circle located at the leverage point (40 cm, 46.81 cm for triangular) that has a $\Delta\Omega = \Delta\Omega_{\text{eff}}$. For the ring the effective slit width ΔR_{eff} is 14.26 microns

Inverse filtering in frequency space

- Uses normalized on-axis PSF
- Usually shows singularities due to zeros in PSF
- MCNPX PSF does not have zeroes
- Noise amplification due to small values in PSF

$$I(k_x, k_y) = O(k_x, k_y) \cdot \text{PSF}(k_x, k_y) \quad O(k_x, k_y) = \frac{I(k_x, k_y)}{\text{PSF}(k_x, k_y)}$$

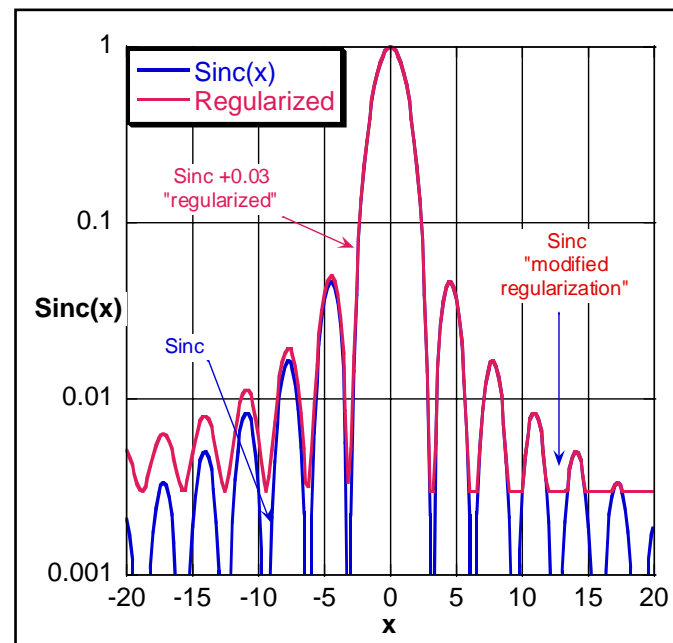


Regularization

- Shifts PSF by a constant

M-Regularization

- Replaces zeros and small values of PSF by a constant
- Regularization parameter is small
 - 0.001% - 0.1% of PSF maximum in frequency space
- Requires low-pass filtering
 - 7 - 15 micron FWHM Gaussian



Deconvolved source comparison

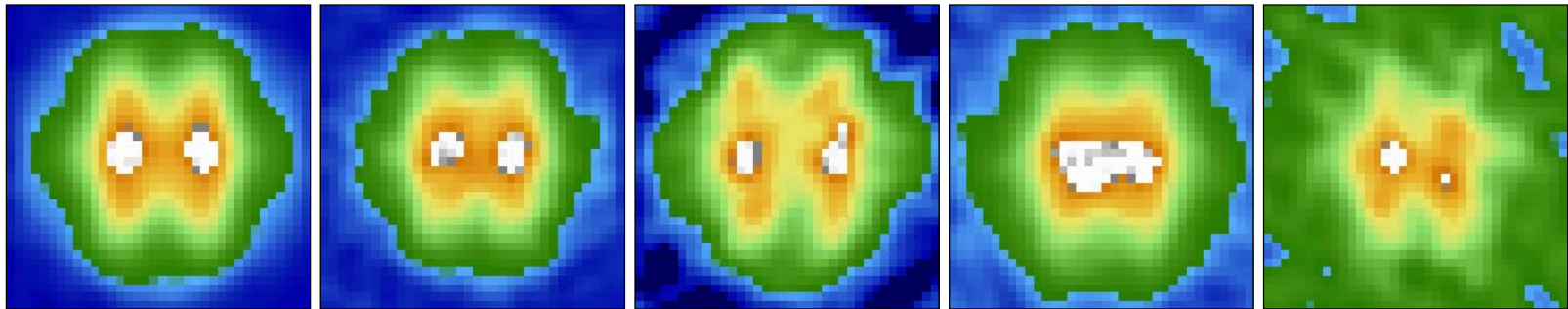
Square
Pinhole

Triangular
Wedge

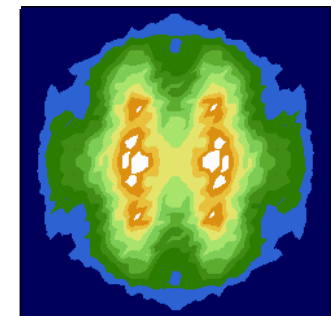
Mini
Penumbral

Annular

Penumbral



- Images are equivalent to 100x100 microns at the source
- 10 microns FWHM low-pass Gaussian filter
- Same regularization parameter
 - Should be optimized for annular and penumbral apertures
- Impossible to accurately determine resolution and signal-to-noise ratio from images



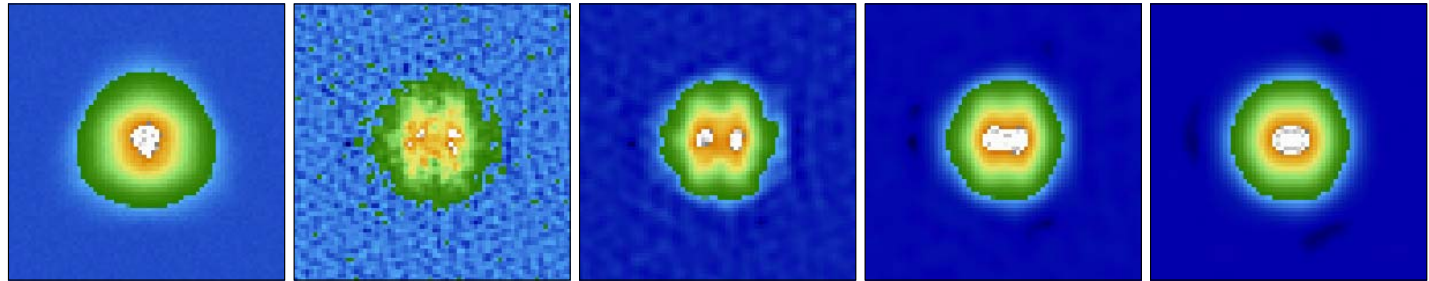
Original
Downscattered
Source

A large uniform surrogate source is used to estimate SNR for the images at the 20% contour

Raw

Deconvolved and Filtered

Triangular
wedge array
6 - 10 MeV



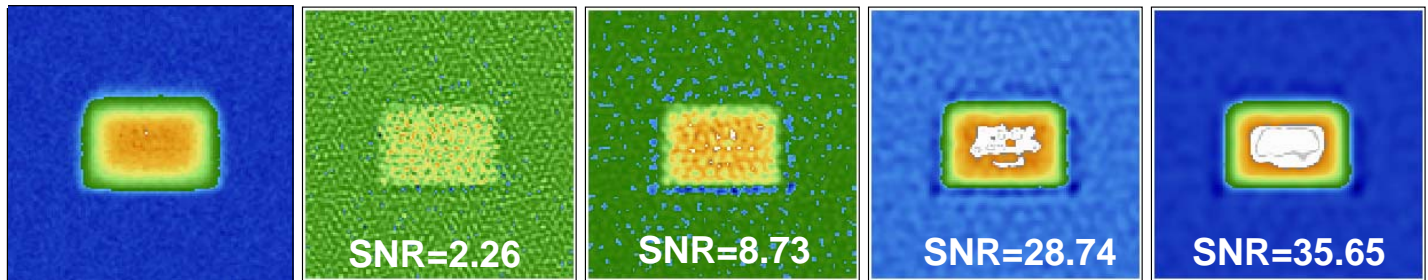
7 μ

10 μ

15 μ

20 μ

Flat Source
150 μ x 100 μ
250 μ pixels



SNR=2.26

SNR=8.73

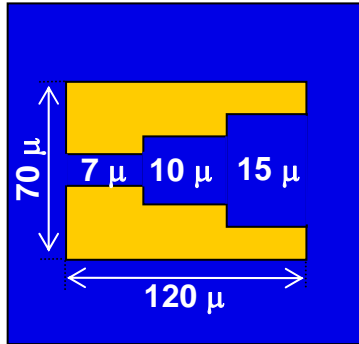
SNR=28.74

SNR=35.65

- Collisions per pixel follow Poisson statistics
- Energy deposited and photo-electrons per pixel are non-Poisson convolutions
 - Very close to Gaussian
- Artificial slab source can be used to “calibrate” 20% contour values after deconvolution
- NIS Requirements:
 - Resolution = 10 microns
 - SNR 20% contour = 10

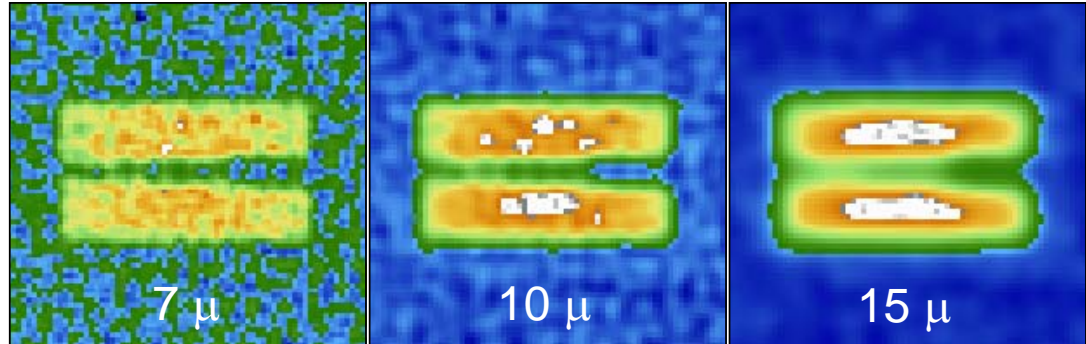
A different surrogate source provides an estimate of resolution

MCNPX Source Geometry



Not to scale

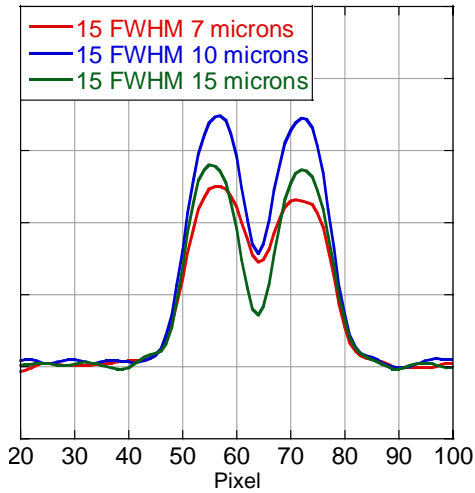
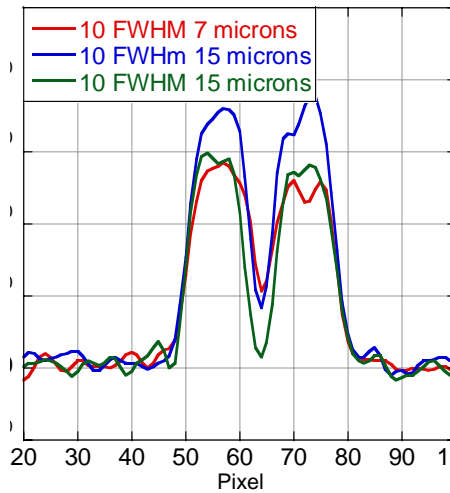
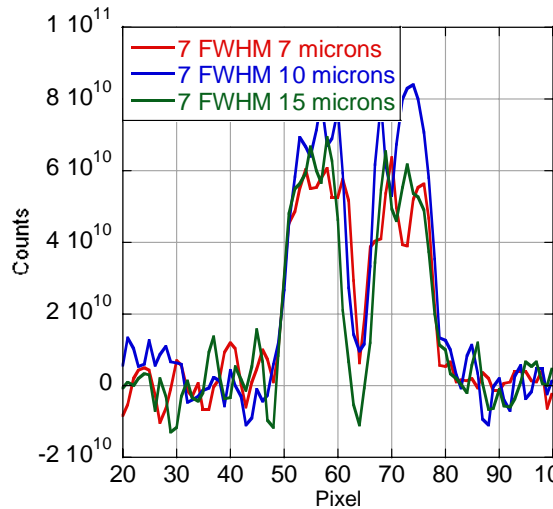
Square Pinhole Array Deconvolved Sources



SNR=8

SNR=15

SNR>15



- Aperture satisfies the “Raleigh criterion” for resolution
- NIS Requirements:
Resolution = 10 microns

SNR 20% contour = 10

- Developed imaging system end-to-end model
- MCNPX used to simulate various aperture geometries and neutron sources
- Surrogate sources show that some configurations satisfy the NIS resolution and signal-to-noise requirements
 - Final decision will be based on aperture fabrication
 - Optics / recording system must be included for final conclusions
- Simple image processing (modified regularization) is adequate for system studies

- Study sensitivities to system imperfections
- Aperture and detector options must be compared for optimization
- Refine calculations of aperture PSFs
- Refine calculations of resolution vs. SNR