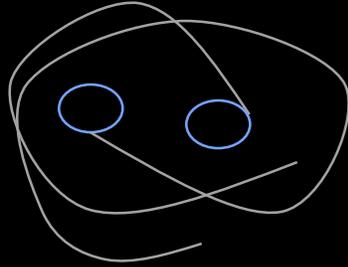


# A Periodic Table for Black Hole Orbits



Ultimate observation

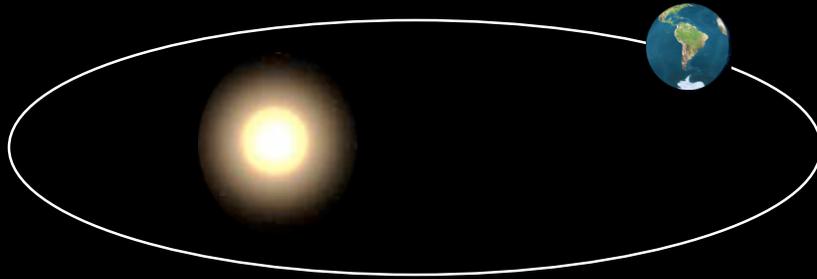
Gabe Perez-Giz  
NSF

JL

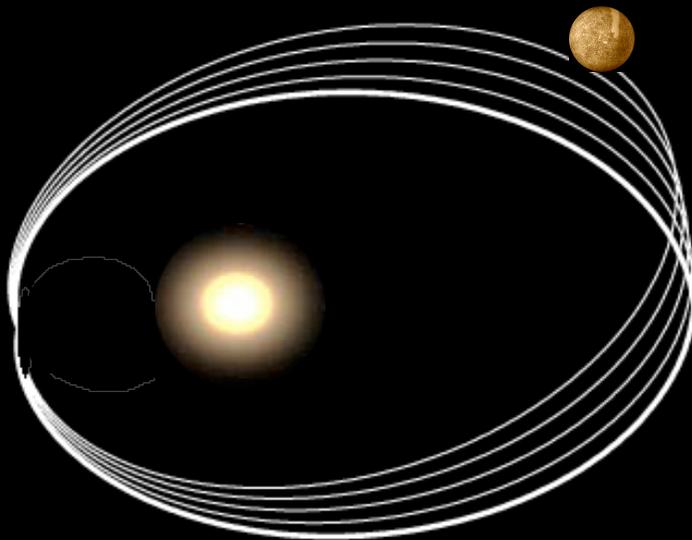


Janna Levin

# 2body problem is insoluble



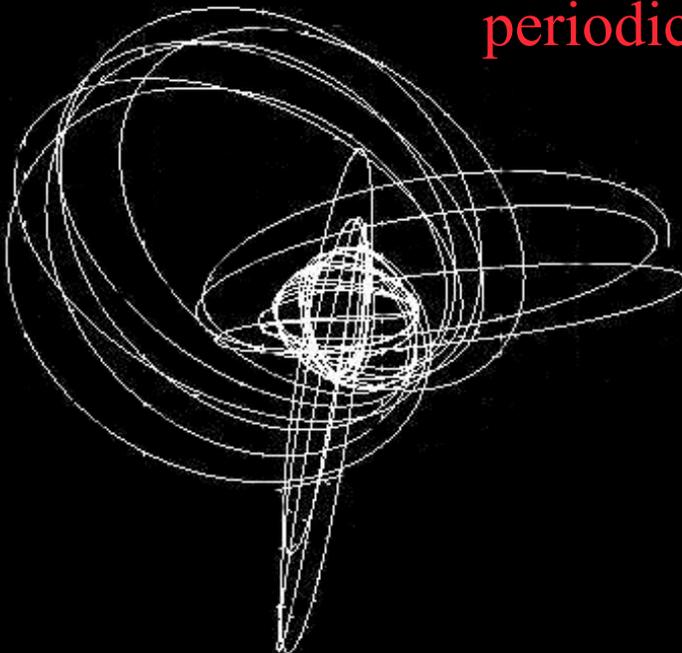
Newtonian  
elliptical  
orbits  
regular



Relativistic  
precession  
test particle  
regular



In the strong-field regime  
periodic table of BH orbits



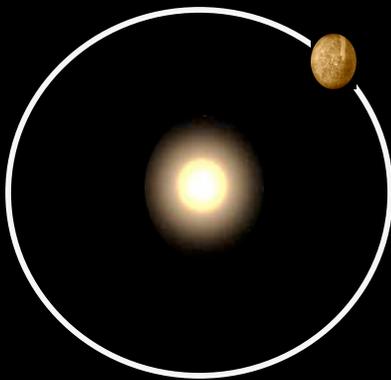
Relativistic  
2body with  
spins -  
chaotic

# Periodic Orbits are Special

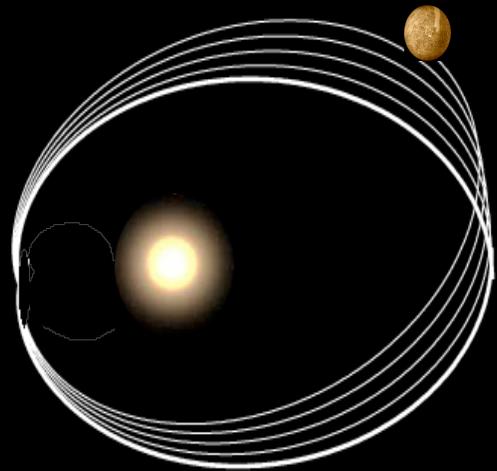
Some orbits are more equal than others

For example, **Circular Orbits** are special

1. Easy to handle
2. Some orbits look like small perturbations to circular ones



Circular orbit



Nearby low-eccentricity  
precessing orbit

## Periodic Orbits are Special

Some orbits are more equal than others

For example, **Circular Orbits** are special

1. Easy to handle
2. Some orbits look like small perturbations to circular ones

Yet, most orbits are not near circular orbits  
circulars cannot encode entire BH dynamics

**Periodic Orbits** are even more special

1. Easy to handle
2. **All** orbits look like small perturbations to periodic ones

**all** orbits are near periodic orbits  
periodics **can** encode entire BH dynamics

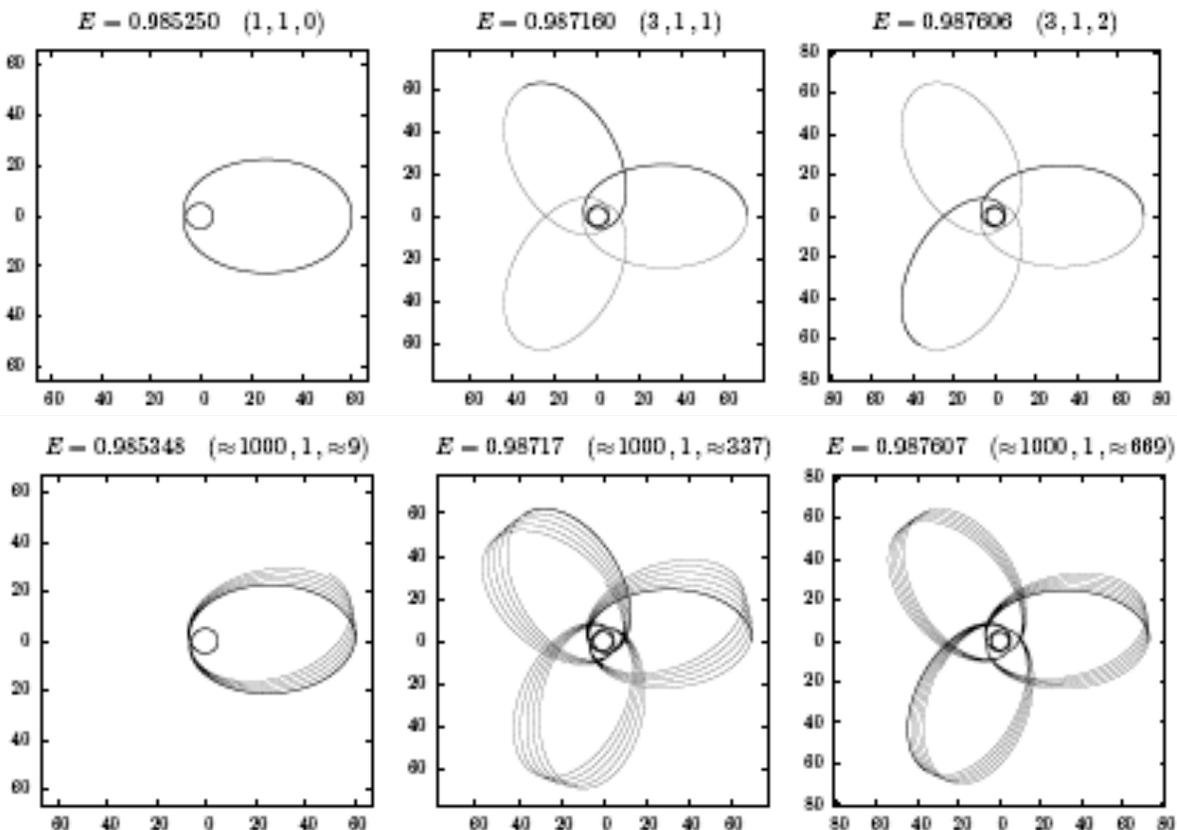
# Periodic Orbits Encode Entire Black Hole dynamics

Periodic Orbits are even more special

1. Easy to handle

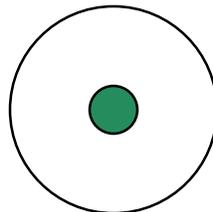
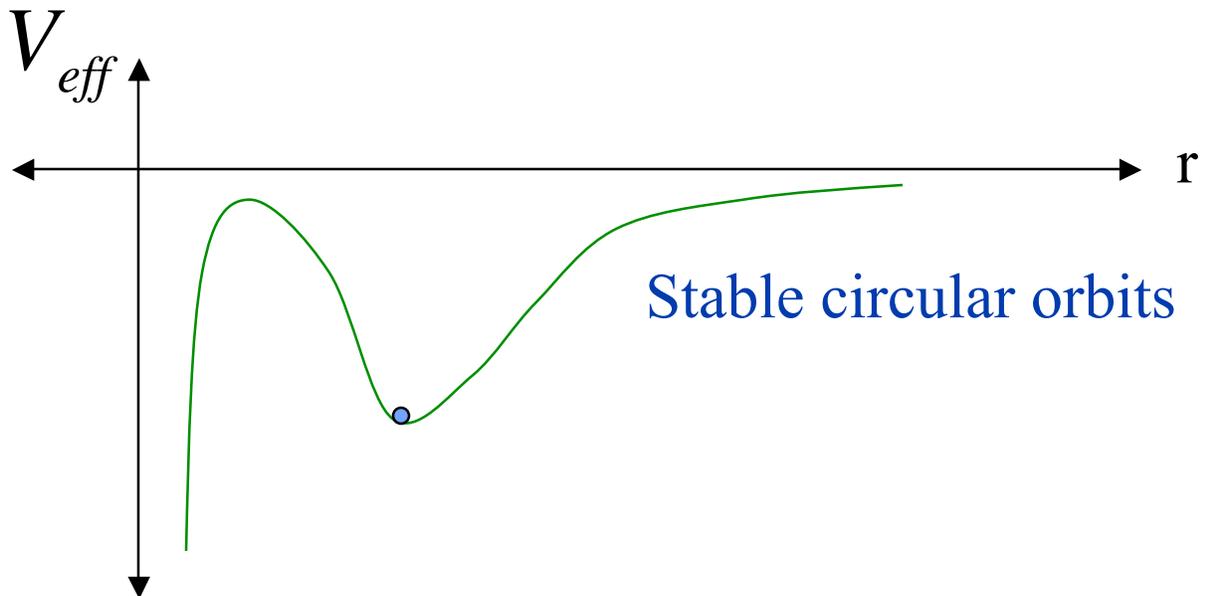
2. *All* orbits look like small perturbations to periodic ones

*all* orbits are near periodic orbits  
periodics *can* encode entire BH dynamics



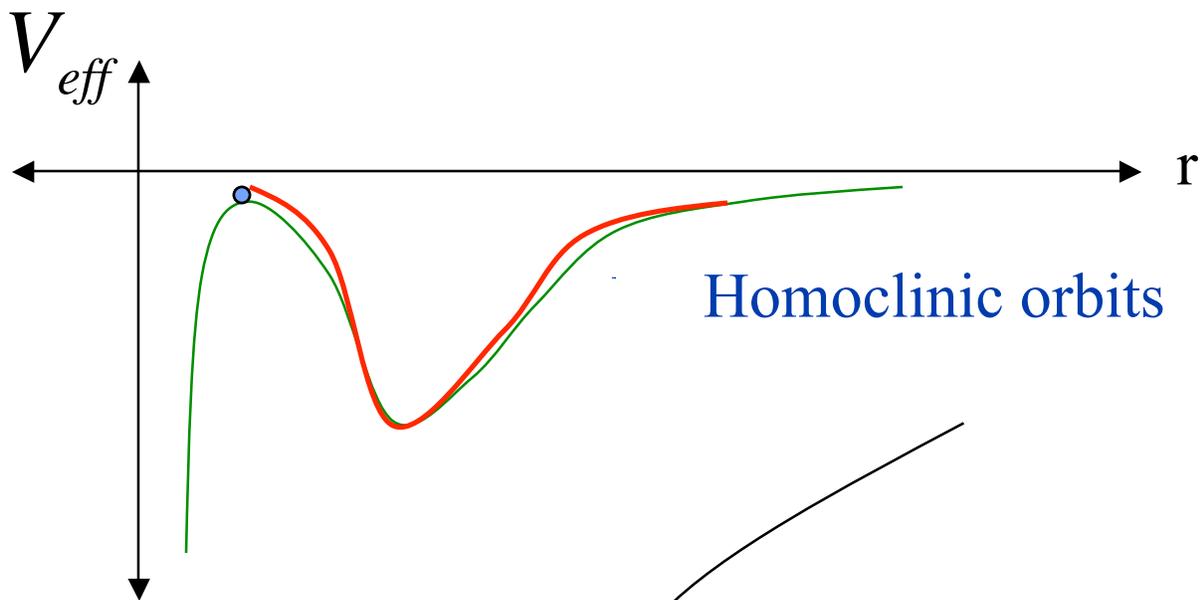
# Schwarzschild reduces to motion in one dimension

$$\frac{1}{2}\dot{r}^2 + \underbrace{\frac{1}{2}\left(1 - \frac{2M}{r}\right)\left(\frac{L^2}{r^2} + 1\right)}_{V_{eff}} = \frac{E^2}{2}$$

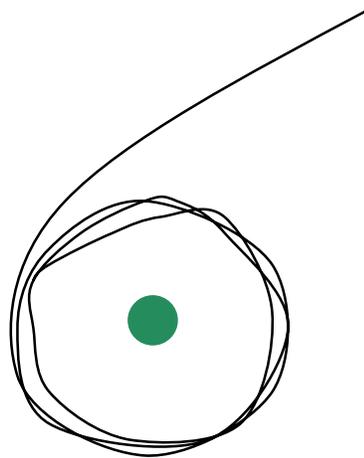


# Schwarzschild reduces to motion in one dimension

$$\frac{1}{2} \dot{r}^2 + \underbrace{\frac{1}{2} \left( 1 - \frac{2M}{r} \right) \left( \frac{L^2}{r^2} + 1 \right)}_{V_{eff}} = \frac{E^2}{2}$$

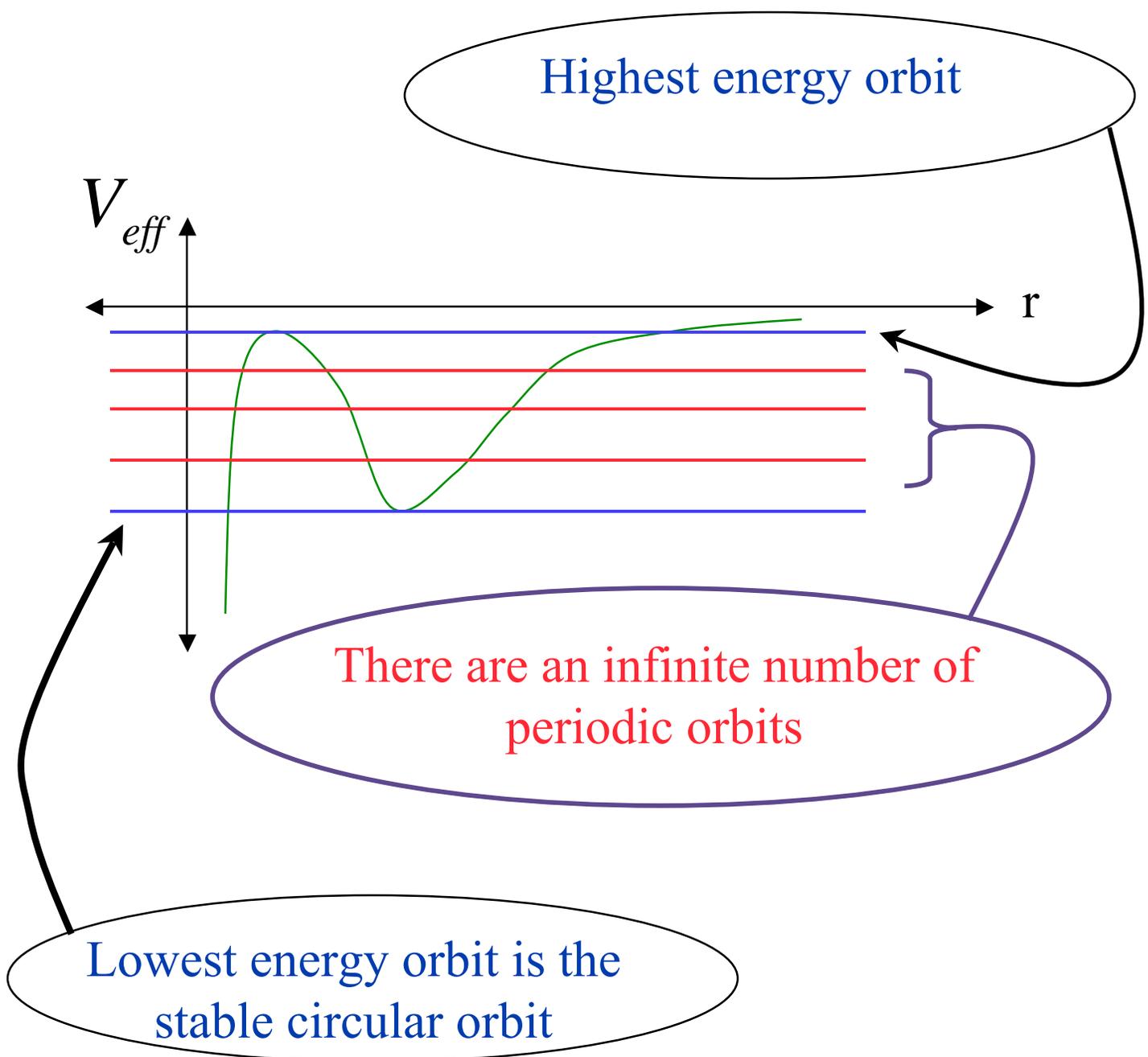


Homoclinic orbits approach the same fixed point in the infinite past & the infinite future: mark intersection of stable and unstable manifolds



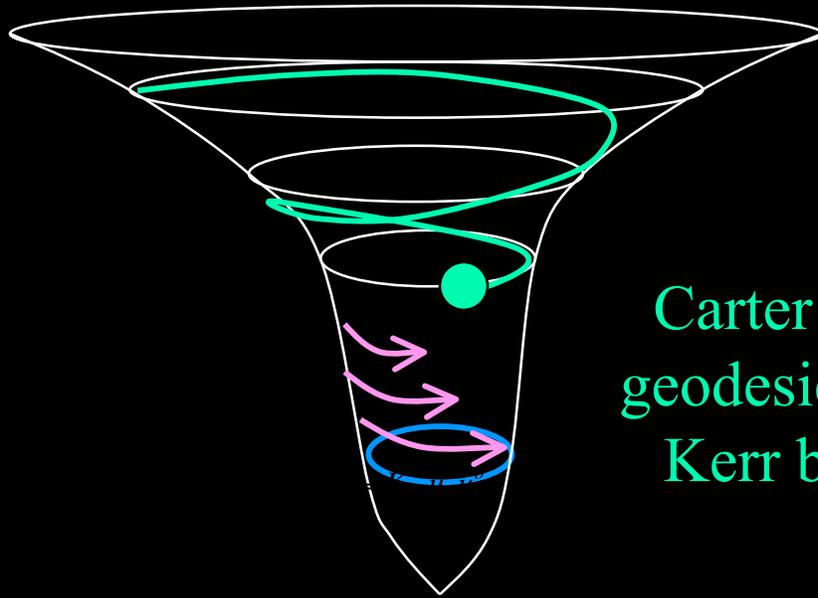
# A Periodic Table of Black Hole Orbits

$$\frac{1}{2} \dot{r}^2 + V_{eff} = \frac{E^2}{2}$$



# Begin Our Story Here

We know orbits around a spinning black hole



Carter found the geodesics around a Kerr black hole

Timelike Killing field, an axial Killing field, and orbits still timelike

$$E, L, \mu, Q$$

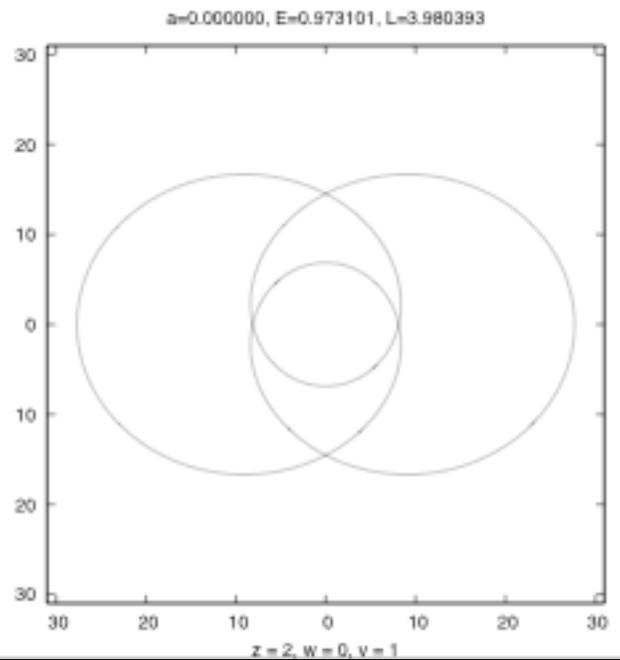
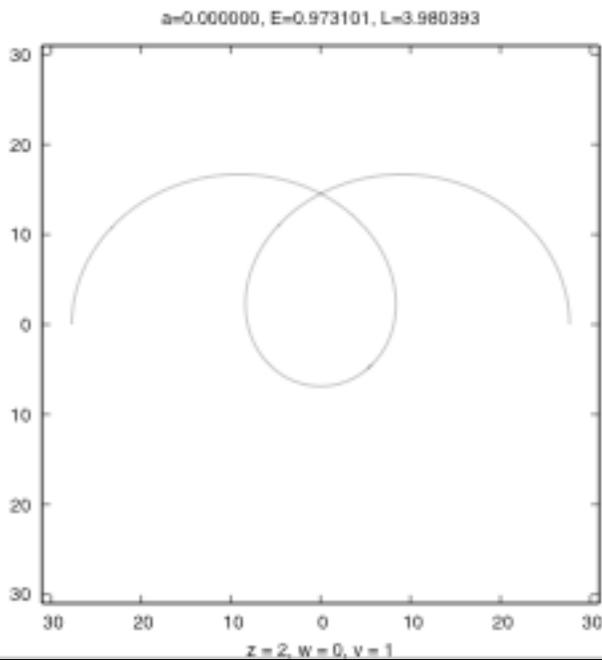
Where  $Q$  is the Carter constant generated by a Killing tensor

Therefore motion lies on tori and integrable

# Taxonomy of Periodic Orbits

$z = \text{zooms}$   
(leaves)

Kerr  
Not Chaotic  
Equatorial  
Bound, Non-plunging

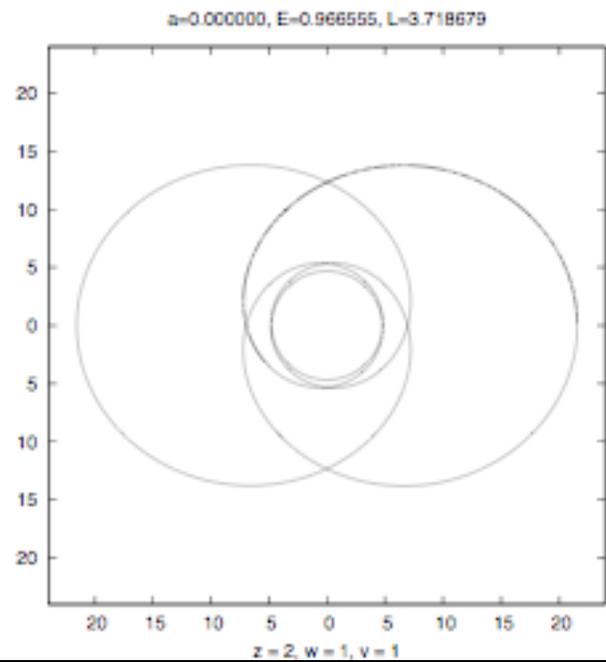
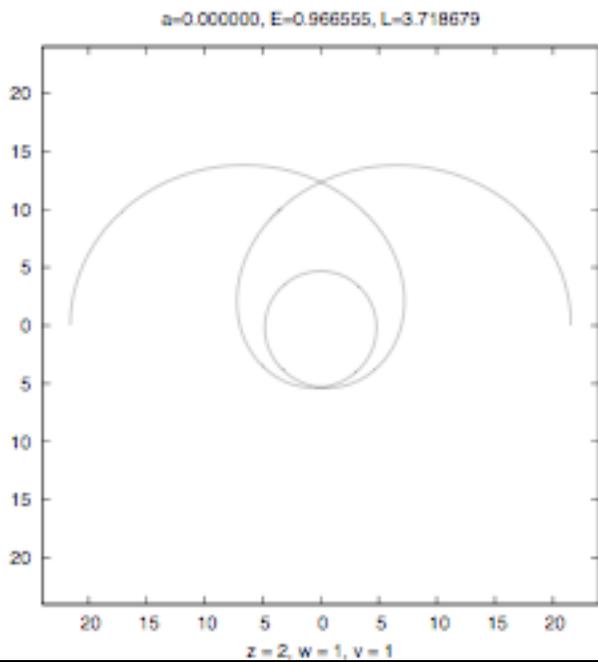


$$\Delta\phi_r = 3\pi$$

$$\Delta\phi = 3\pi z = 6\pi$$

# Taxonomy of Periodic Orbits

$$z = 2$$
$$w = \text{whirls}$$

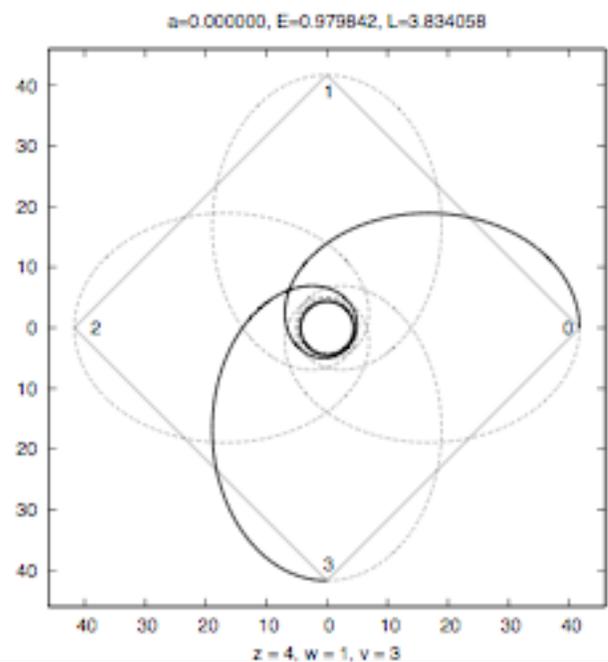
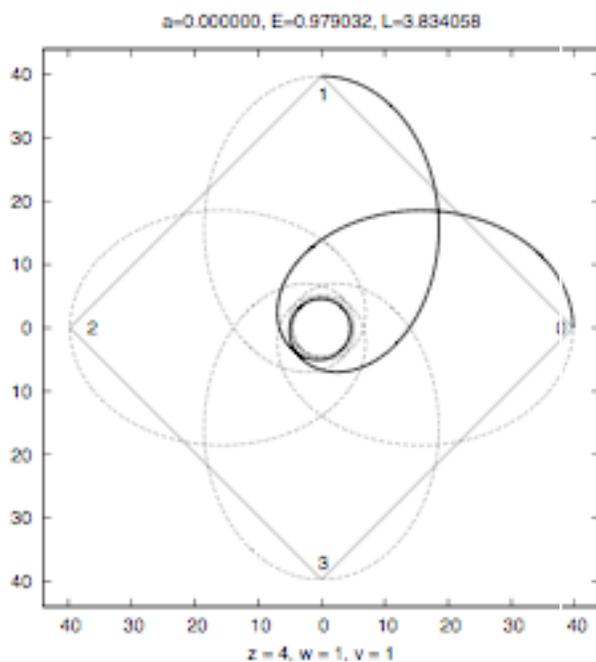


$$\Delta\phi_r = 5\pi$$

$$\Delta\phi = 5\pi z = 10\pi$$

# Taxonomy of Periodic Orbits

$$z = 4$$
$$w = 1$$
$$v = \text{vertices}$$



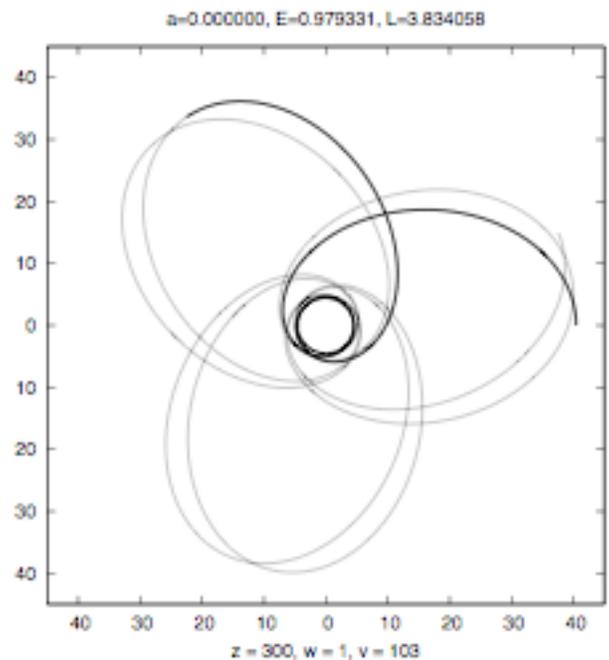
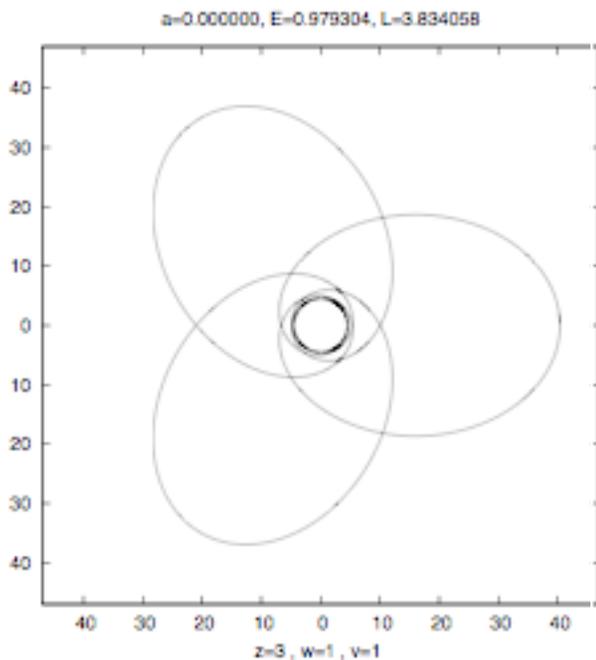
$$\Delta\phi_r = 2\pi \left( 1 + w + \frac{v}{z} \right)$$

$$4\pi + \pi/2$$

$$4\pi + 3\pi/2$$

# High $z$ orbits as precession of low $z$ orbits

$(z = 3, w = 1, v = 1)$



$(z = 300, w = 1, v = 103)$

## Finding periodic orbits with rational numbers

$$(z, w, v)$$

$$\Delta\varphi_r = 2\pi \left(1 + w + \frac{v}{z}\right)$$

Writing this another way, every periodic orbit is associated with a rational number:

$$q = w + \frac{v}{z} = \frac{\Delta\varphi_r}{2\pi} - 1$$

We can find all of the periodic orbits by integrating the Kerr equations subject to the condition that the accumulated angle is fixed by the rational number  $q$

$$\Delta\varphi_r = 2 \int_{t(r_p)}^{t(r_a)} \frac{d\varphi}{dt} dt$$

## Another derivation of $q$

Every equatorial orbit has two fundamental frequencies

Radial frequency;  $T_r$  is  
the radial period

$$\omega_r = \frac{2\pi}{T_r}$$



Angular frequency;  
average of  
 $\frac{d\varphi}{dt}$  over  $T_r$

$$\omega_\varphi = \frac{1}{T_r} \int \frac{d\varphi}{dt} dt = \frac{\Delta\phi_r}{T_r}$$

For periodic orbits, their ratio is rationally related

$$\frac{\omega_\varphi}{\omega_r} = \frac{\Delta\varphi_r}{2\pi} = 1 + w + \frac{v}{z} = 1 + q$$

$\omega_r$  and  $\omega_\varphi$  are the same frequencies derived from  
an action-angle formulation of the dynamics

# Circulars

Can still define this ratio for stable circulars

$$\omega_{\varphi} = \frac{d\varphi}{dt}$$

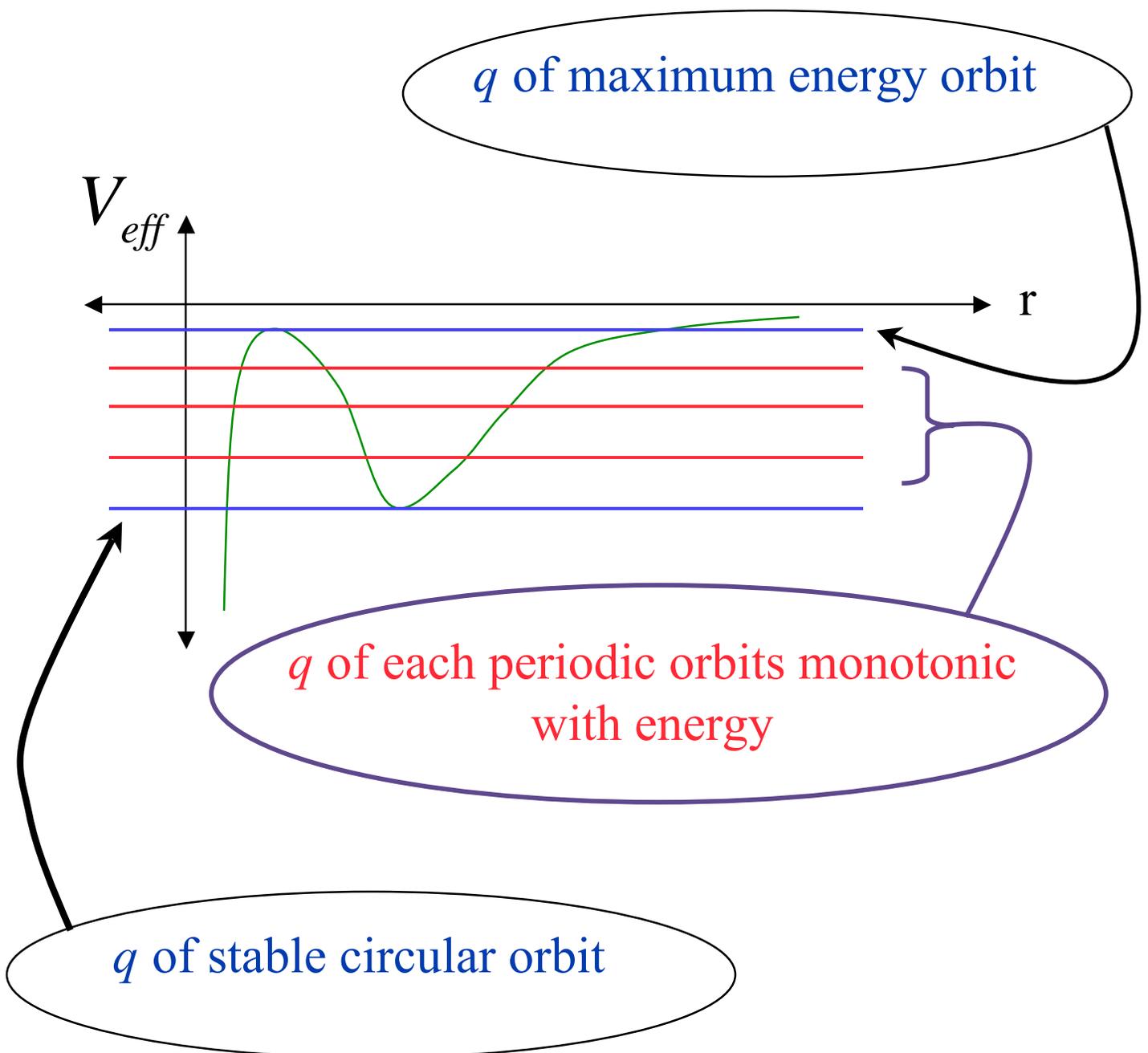
$\omega_r \rightarrow$  Frequency of radial oscillations about the circular

Can also define as the limit as the eccentricity  $e$  goes to zero

$$q_c = \lim_{e \rightarrow 0} \left( \frac{\omega_{\varphi}}{\omega_r} - 1 \right)$$

This allows us to build a periodic diagram

$$q_c < q < q_{max}$$

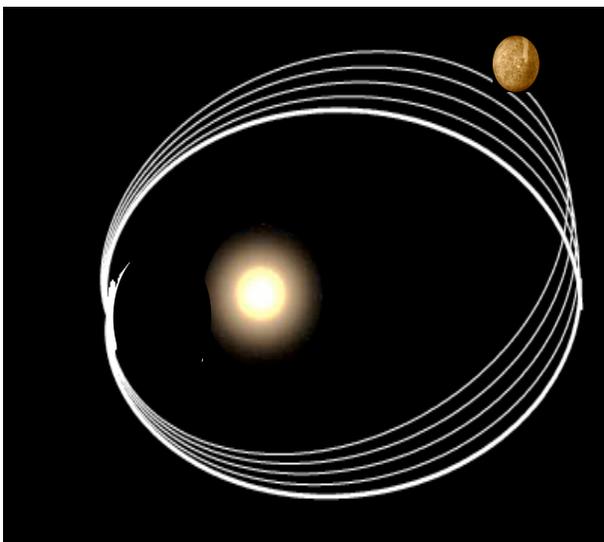
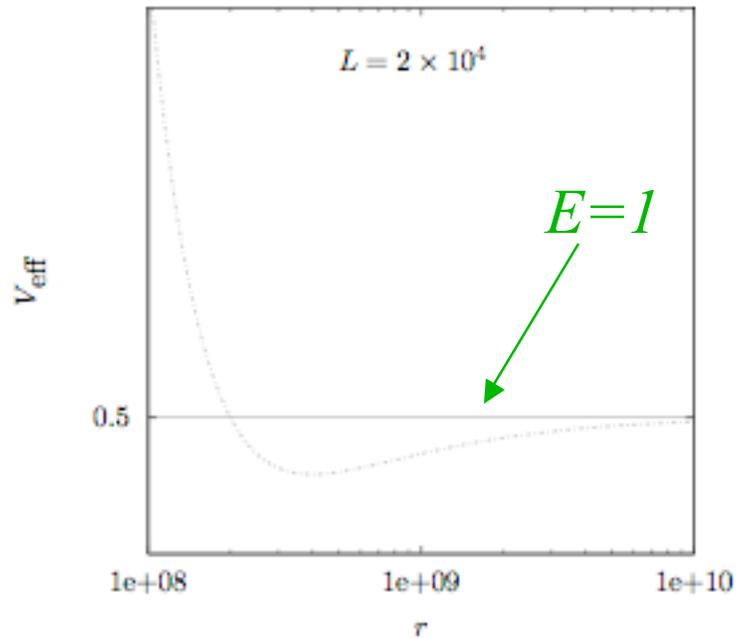


$$q_c < q < q_{max}$$

The weak-field regime

Effective potential for  
solar system values

$$q_{max} = q(E=1)$$



$$q_c < q_{max} \ll 1$$

$$w + \frac{v}{z} \ll 1$$

$$w = 0$$

$$z \rightarrow \infty$$

$$q_c < q < q_{max}$$

The weak-field regime

Mercury, for instance, precesses 43"/century. With an orbital period of 88 days this comes to  $\sim 0.1''/\text{orbit}$  or  $360^\circ/(3600^2)$

$$\begin{aligned} q_c < q_{max} &\ll 1 \\ \omega + \frac{v}{z} &\ll 1 \\ \omega &= 0 \\ z &\rightarrow \infty \end{aligned}$$

Mercury is close to an orbit with

$$z \approx (3600^2) = 1.296 \times 10^7$$

leaves and no whirls

$$(1.296 \times 10^7, 0, 1)$$

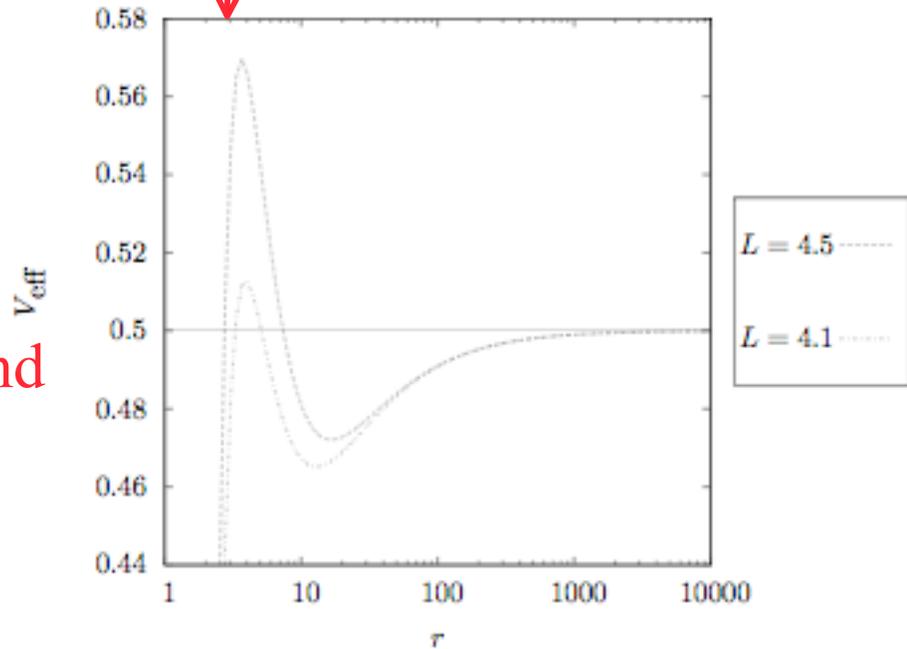
In the weak-field limit, orbits are Mercury-type precessing ellipses

$$q_c < q < \infty_{max}$$

The very strong-field regime

$E =$   
 IBCO

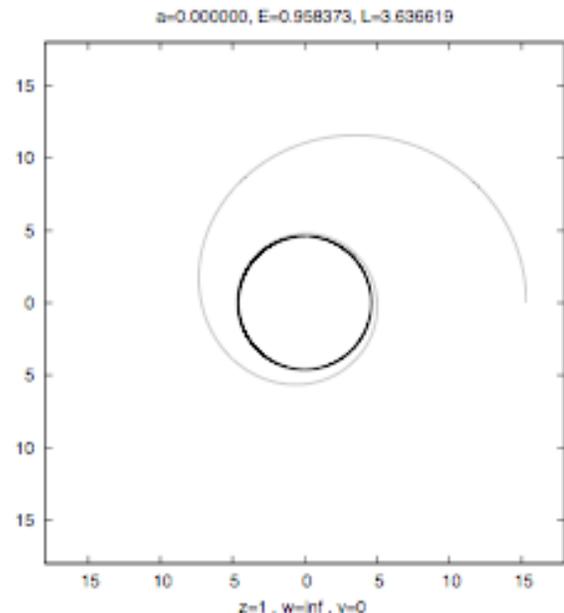
Innermost bound  
 circular orbit



Every unstable bound circular orbit has a  
 homoclinic orbit

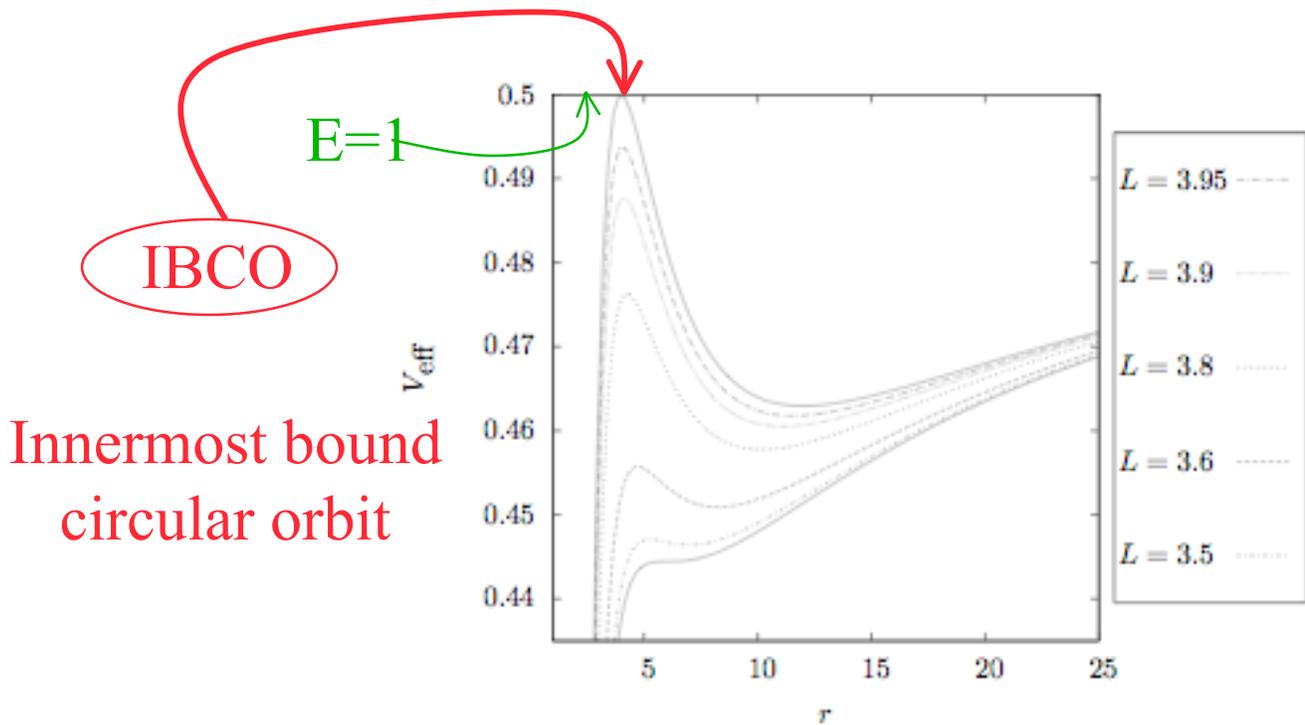
Homoclinic orbits whirl an  
 infinite number of times

$$q_{max} = w + \frac{v}{z} = \infty$$



$$q_c < q < \infty$$

The very strong-field regime



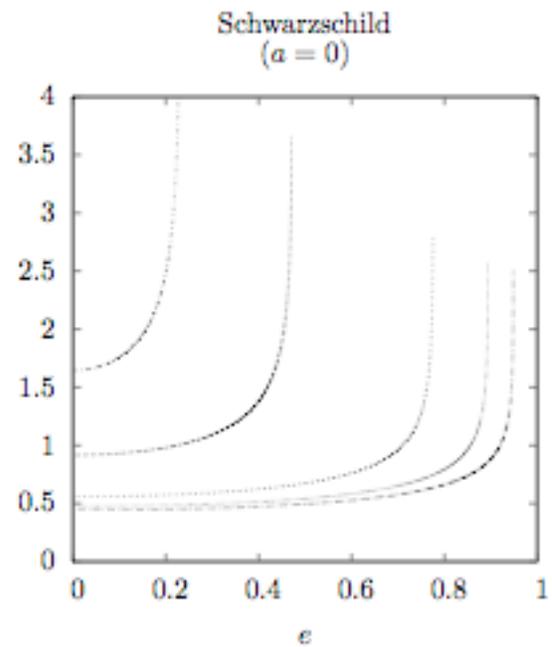
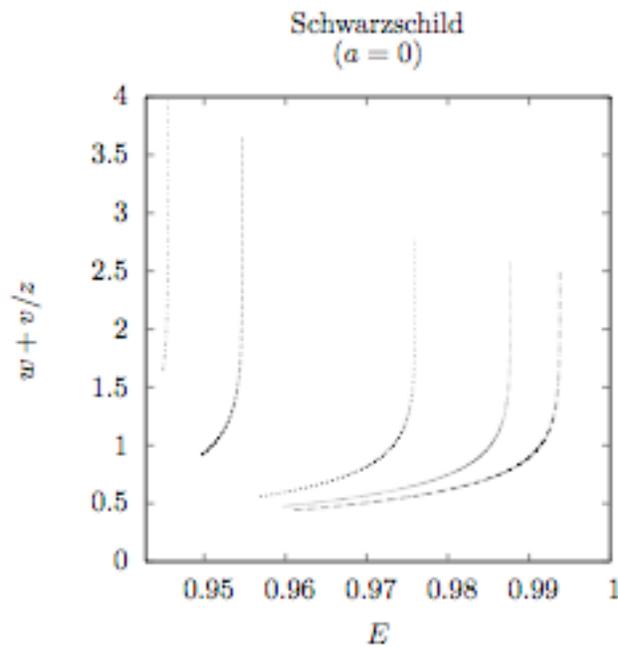
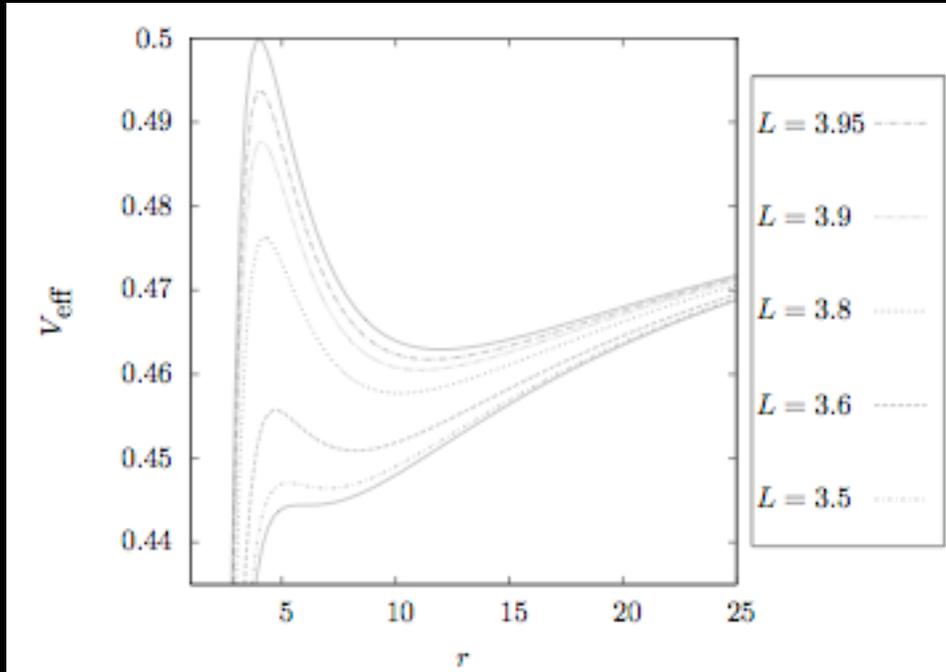
What is  $q_c$  ?

Is  $q_c = 0$  ?

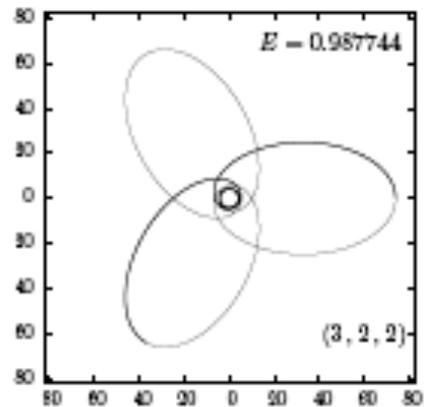
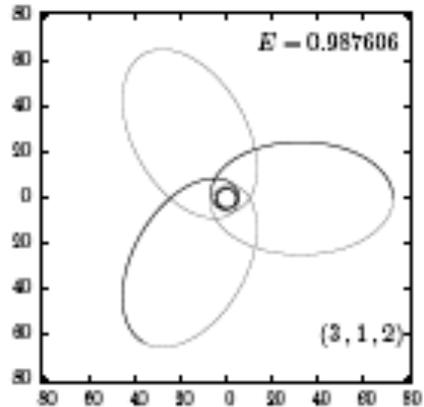
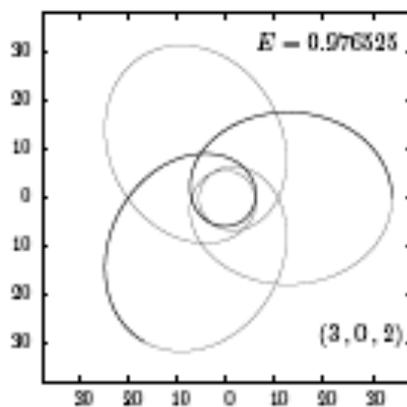
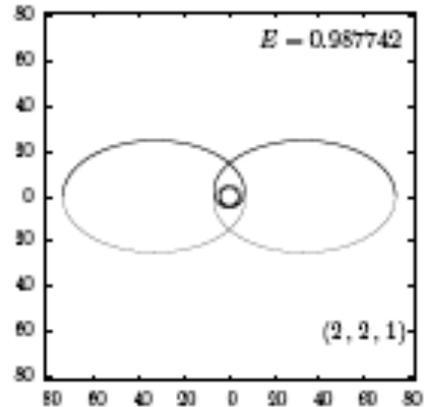
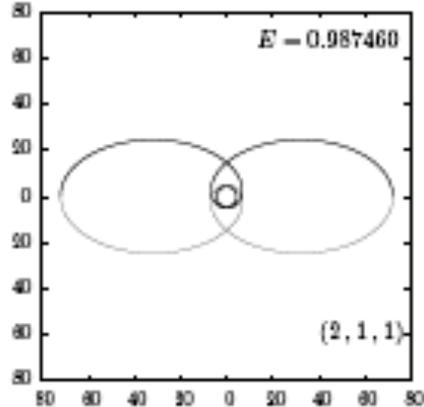
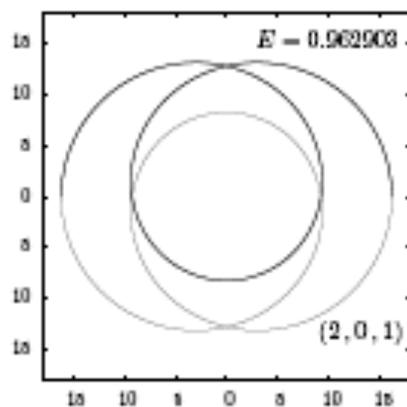
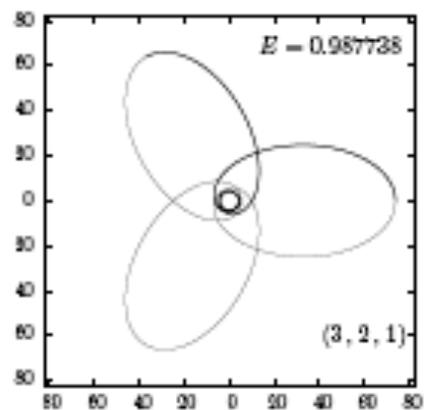
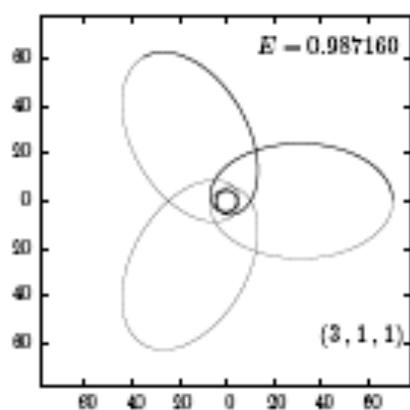
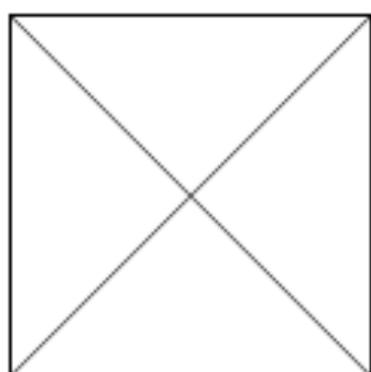
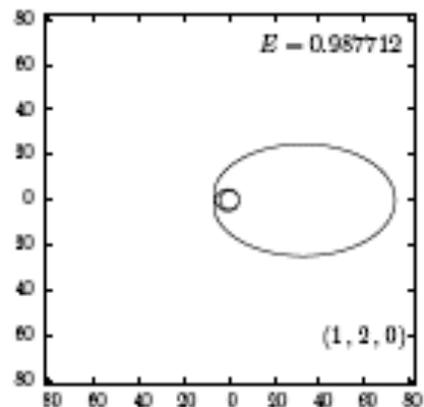
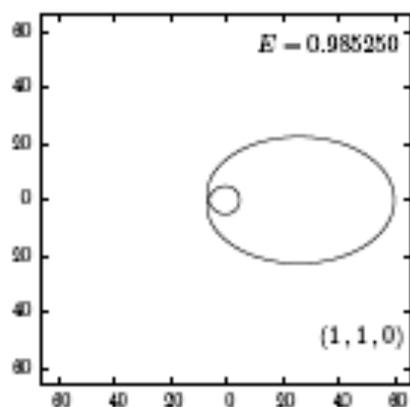
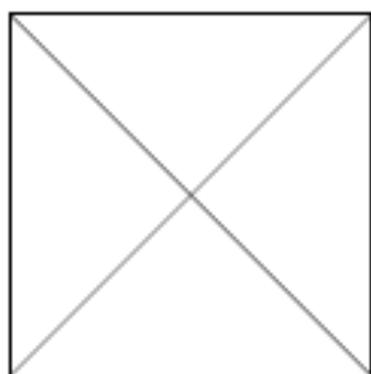
In other words, are all  $q$ , and therefore all periodic orbits, allowed?

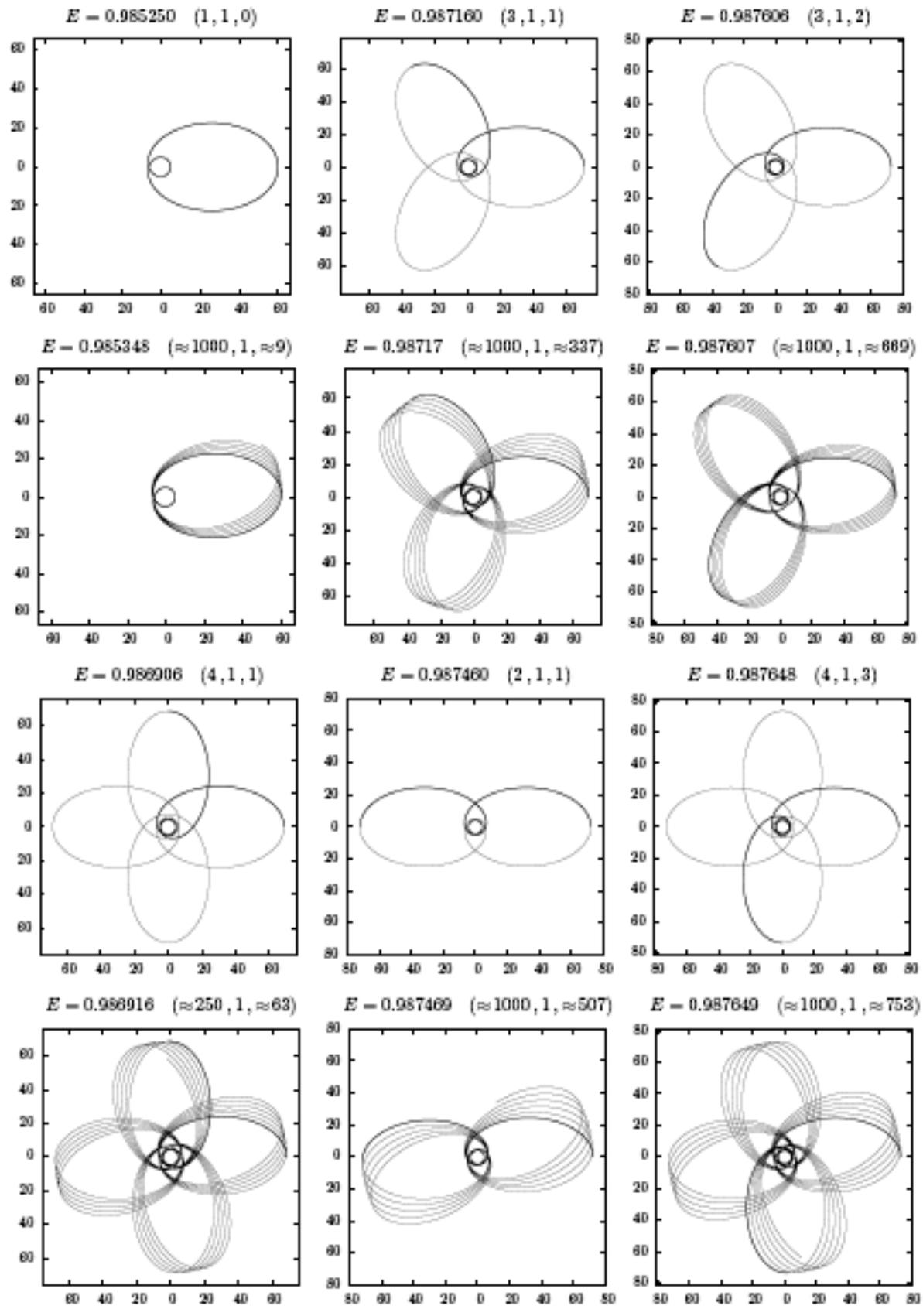
No!

# Range of Periodic Orbits

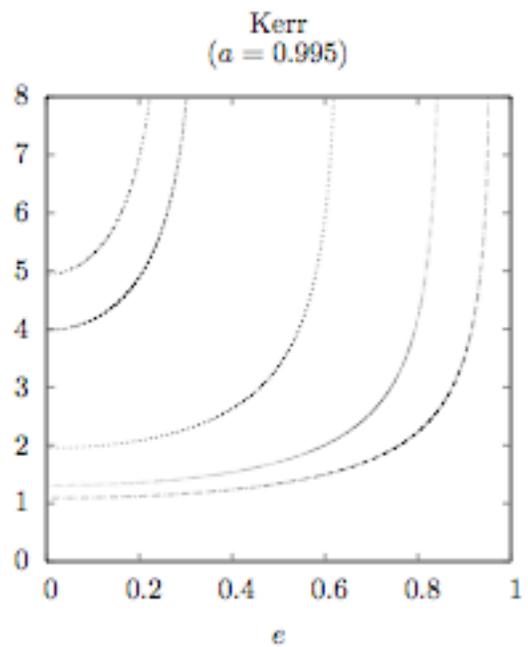
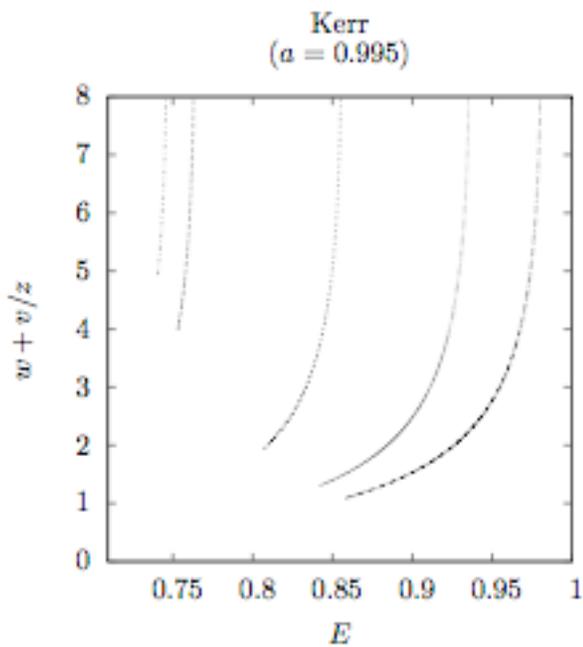


$$\left(w + \frac{v}{z}\right)_{\text{stable circ}} \leq \left(w + \frac{v}{z}\right) \leq \left(w + \frac{v}{z}\right)_{\text{max}}$$



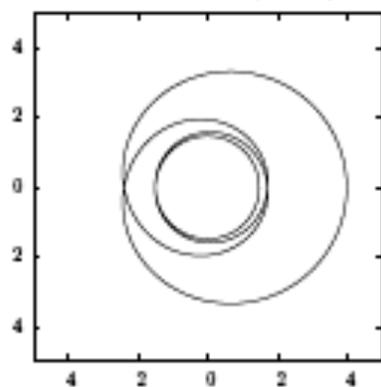


# Range of Periodic Orbits ( $a=0.995$ )

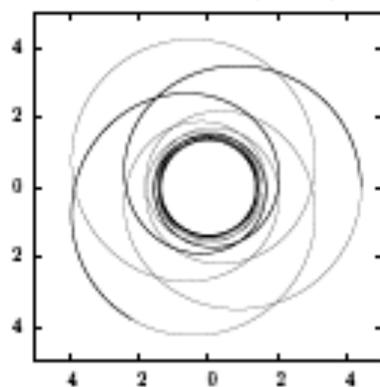


$$\left(w + \frac{v}{z}\right)_{stable\ circ} \leq \left(w + \frac{v}{z}\right) \leq \left(w + \frac{v}{z}\right)_{max}$$

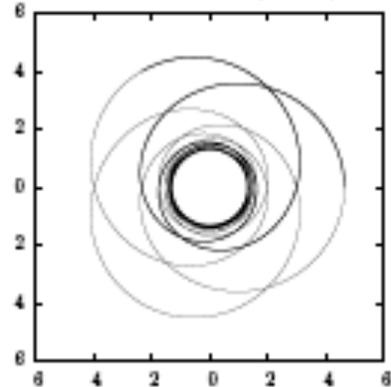
$E = 0.833302$  (1,3,0)



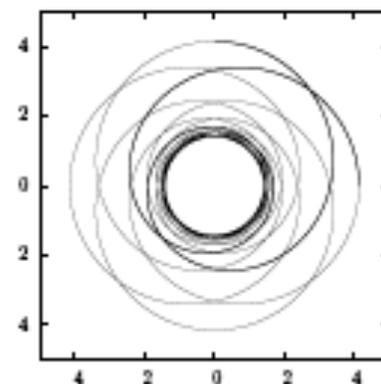
$E = 0.841423$  (3,3,2)



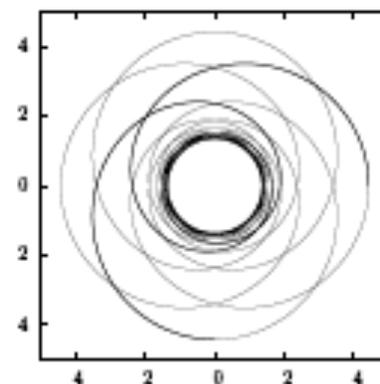
$E = 0.846486$  (3,4,1)



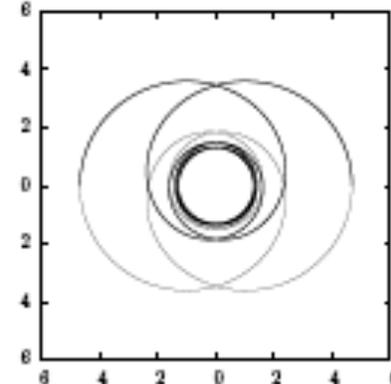
$E = 0.836824$  (4,3,1)



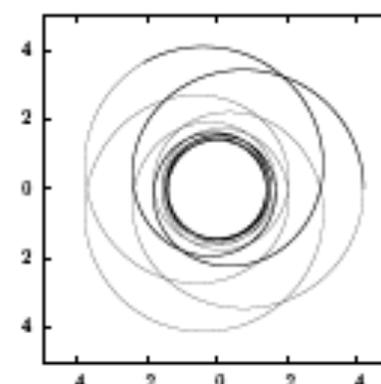
$E = 0.842189$  (4,3,3)



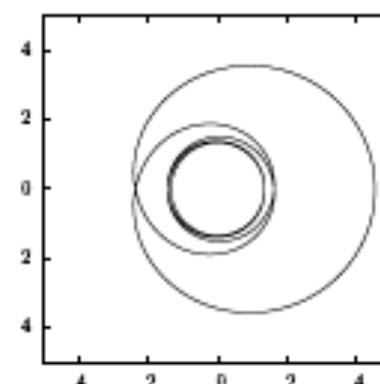
$E = 0.847436$  (2,4,1)



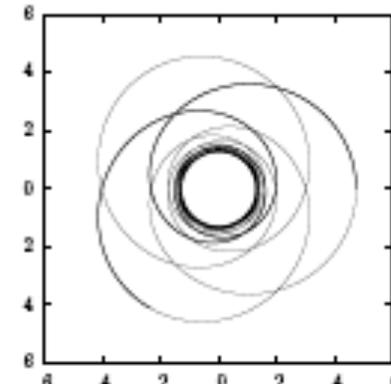
$E = 0.837858$  (3,3,1)



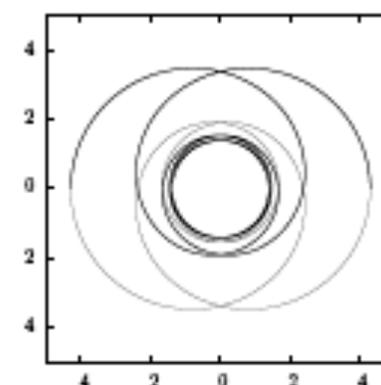
$E = 0.844241$  (1,4,0)



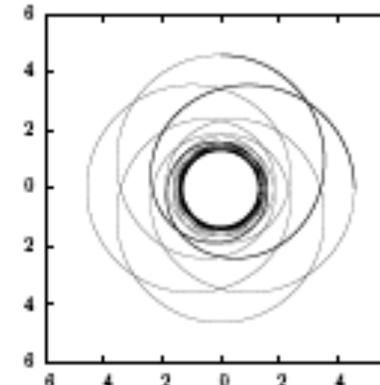
$E = 0.848288$  (3,4,2)



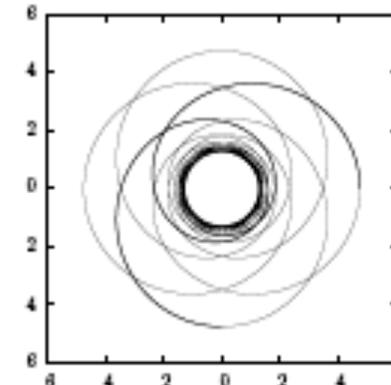
$E = 0.839747$  (2,3,1)



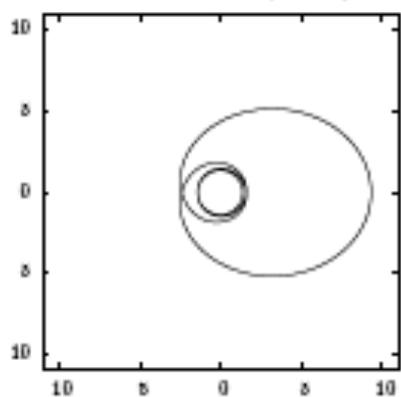
$E = 0.845971$  (4,4,1)



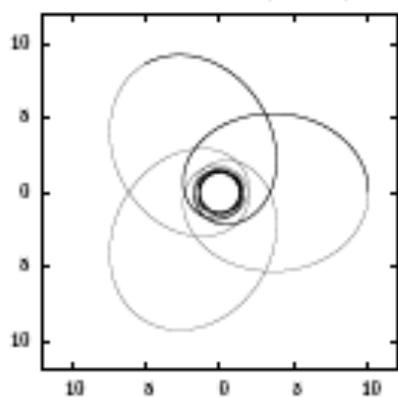
$E = 0.848681$  (4,4,3)



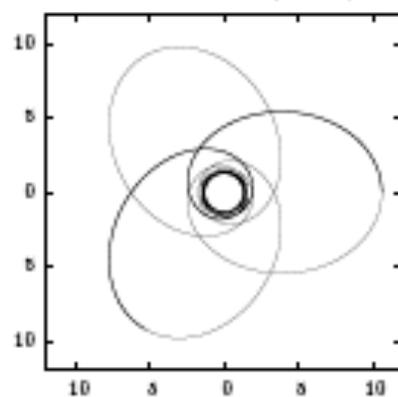
$E = 0.911154$  (1, 3, 0)



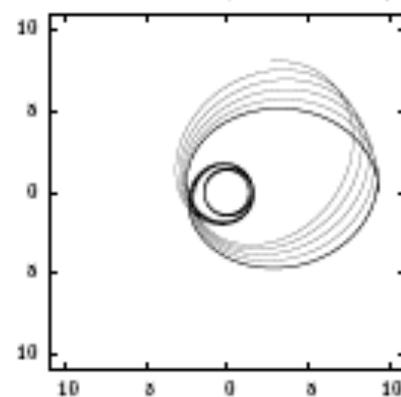
$E = 0.916235$  (3, 3, 1)



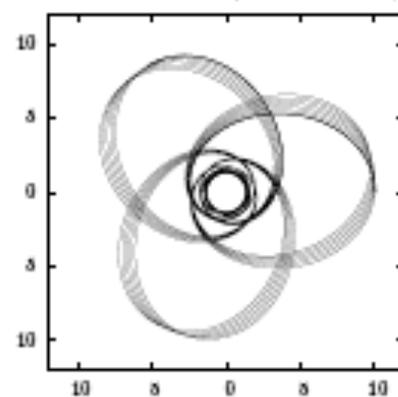
$E = 0.920204$  (3, 3, 2)



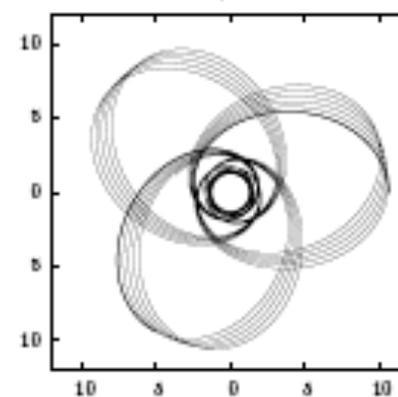
$E = 0.911161$  ( $\approx 500, 3, \approx 13$ )



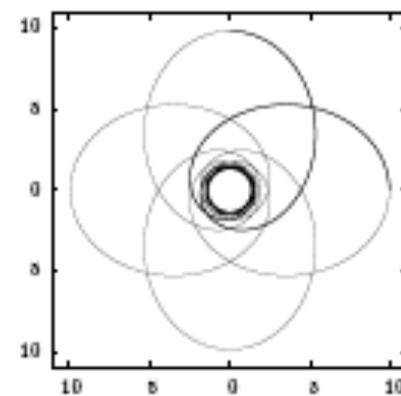
$E = 0.916281$  ( $\approx 125, 3, \approx 42$ )



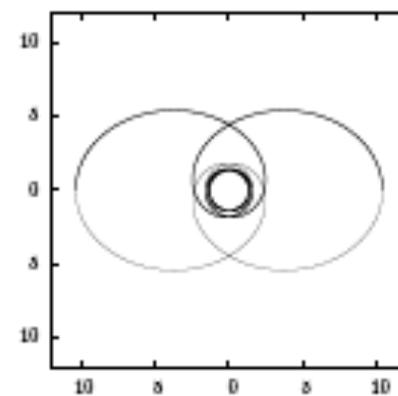
$E = 0.92025$  ( $\approx 1000, 3, \approx 671$ )



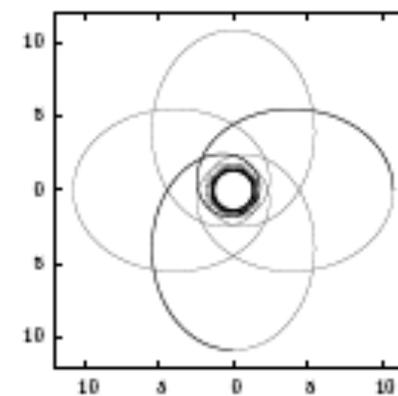
$E = 0.915082$  (4, 3, 1)



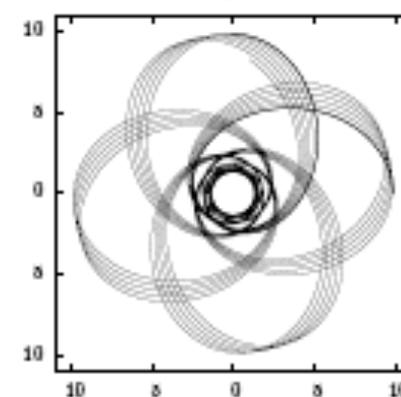
$E = 0.918339$  (2, 3, 1)



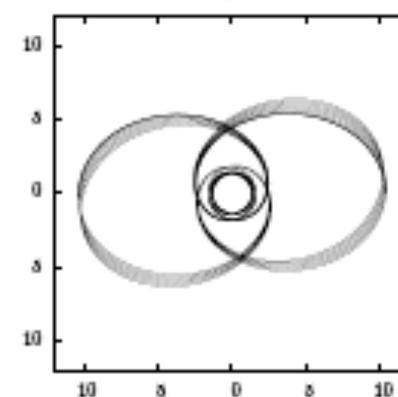
$E = 0.921057$  (4, 3, 3)



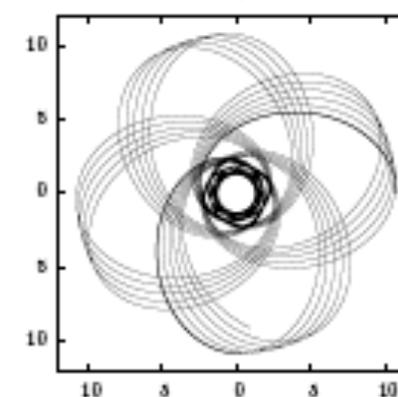
$E = 0.915128$  ( $\approx 1000, 3, \approx 253$ )



$E = 0.918385$  ( $\approx 1000, 3, \approx 503$ )



$E = 0.921103$  ( $\approx 500, 3, \approx 377$ )



# Recap of Dynamical Results

- For a given  $L$ ,

$$q_c < q < q_{max}$$

- In the Newtonian limit,  $q_c \sim q_{max} \sim 0$  ;  
Keplerian orbits are ellipses

- No zoom-whirl behavior in the weak-field regime

- At the ISCO,  $q_c \sim q_{max} \sim \infty$

- In the very strong-field regime (  $L < L_{IBCO}$  )

$$q_c < q < \infty$$

- In the strong-field regime, the simple precessing ellipse familiar from planetary orbits is forbidden

- In the strong-field, all aperiodic eccentric orbits are precessions of low-leaf clovers

- As the ISCO is approached, all orbits zoom and whirl for any  $a$

# Utility of Taxonomy

## Gravitational wave astronomy

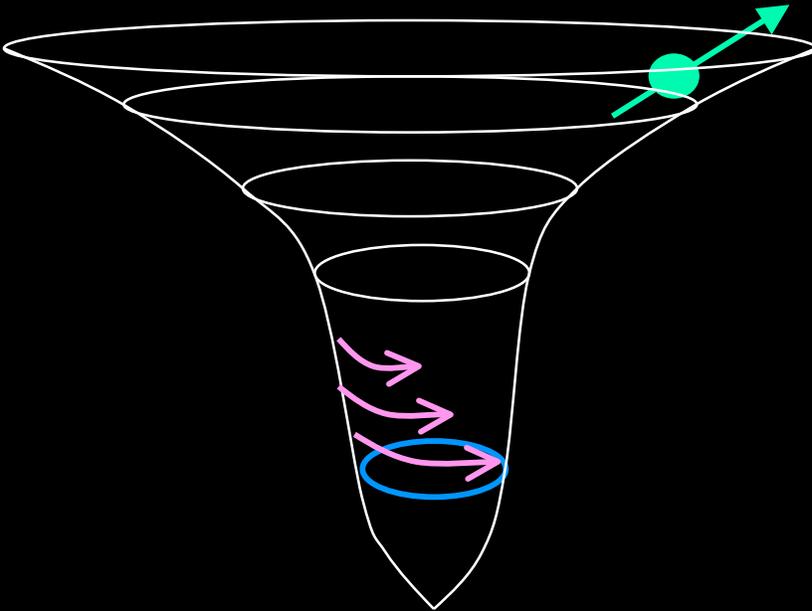
- EMRI's on eccentric inner orbits
- All eccentric orbits will show zooms and all high spin show whirls
- Data analysis

## Computation of Inspiral

- Inspiral through periodic orbits
- Fourier decomposition in 1 frequency instead of 3

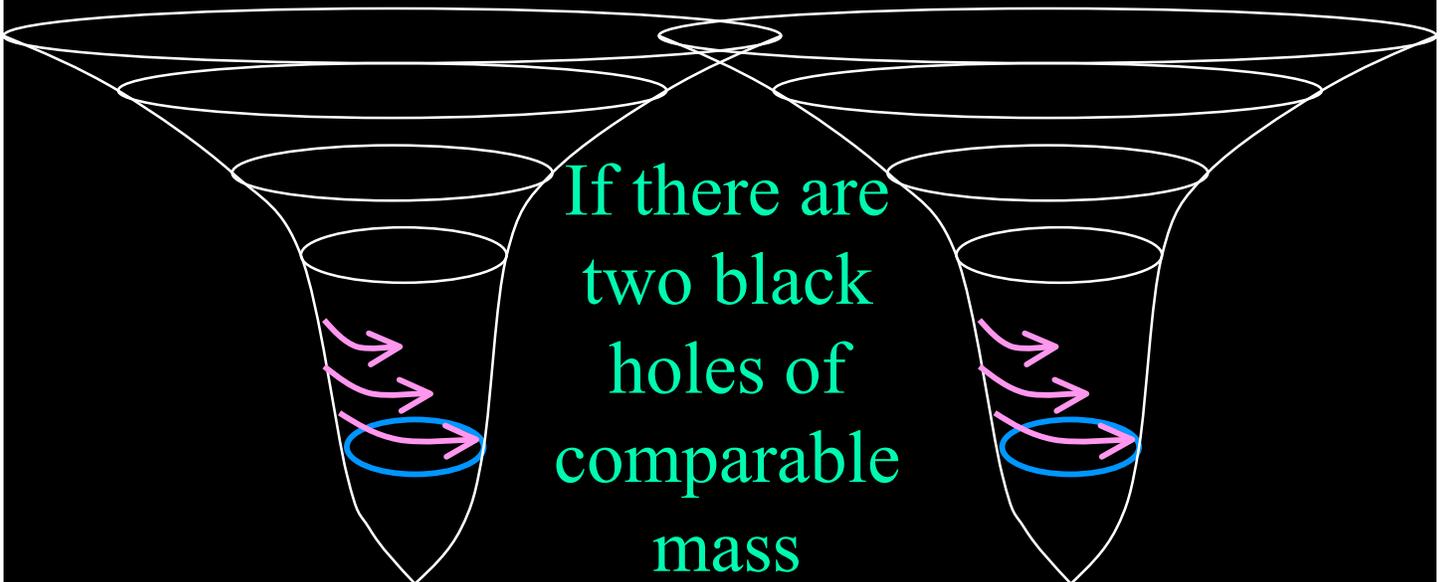
## Spin including transition to chaos

# What don't we Know



If the  
companion  
spins -  
orbits are  
strongly  
altered

There are no known exact solutions to the  
orbits of these BH binaries

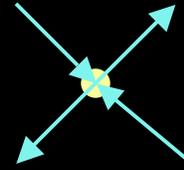


If there are  
two black  
holes of  
comparable  
mass

# The periodic orbits



Elliptical/stable

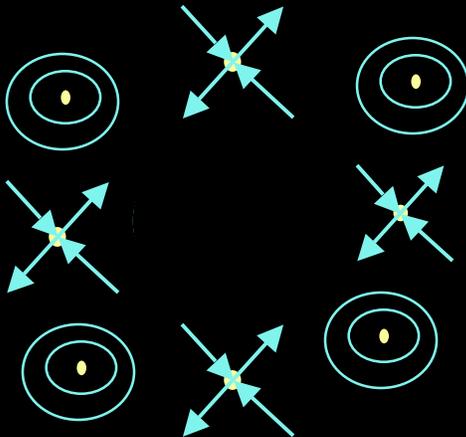


Hyperbolic/unstable

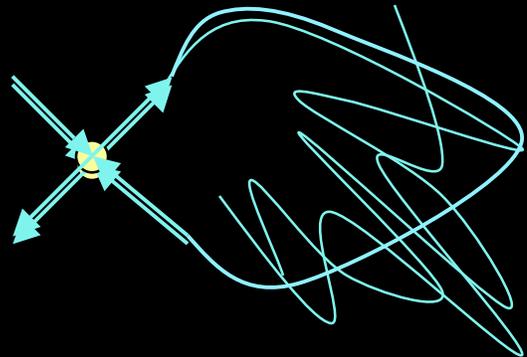
Periodic orbits define the phase space

KAM theorem

Chaos develops around certain unstable periodic orbits



Proliferate into pairs of elliptical and hyperbolic



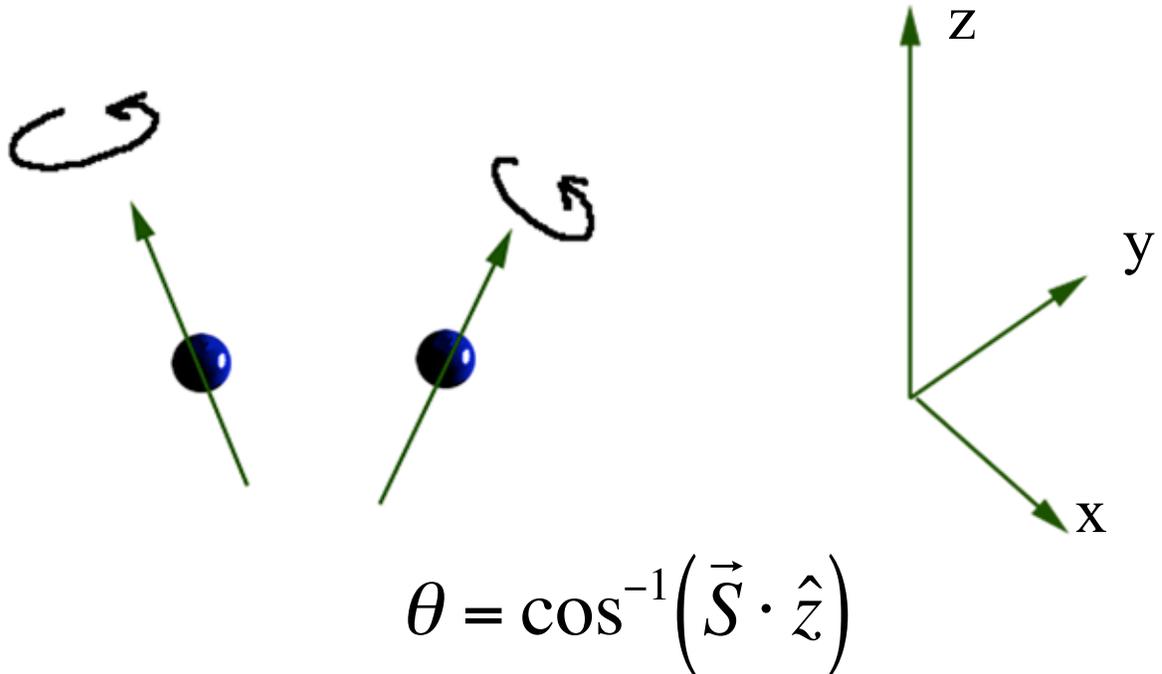
Homoclinic tangle

To pack a proliferating infinity of orbits  
in a finite phase space form a

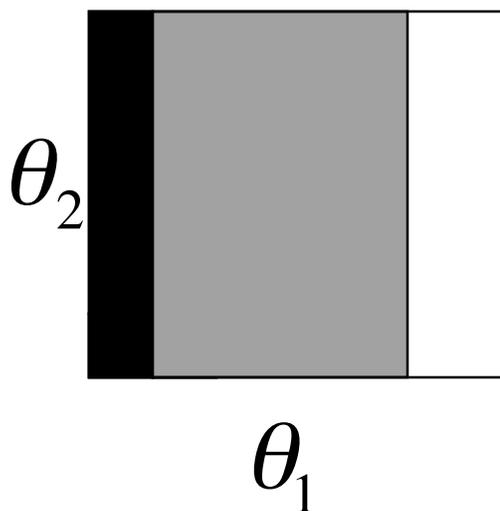
Fractal

# Fractal Basin Boundaries

Post-Newtonian approximation to the 2body problem

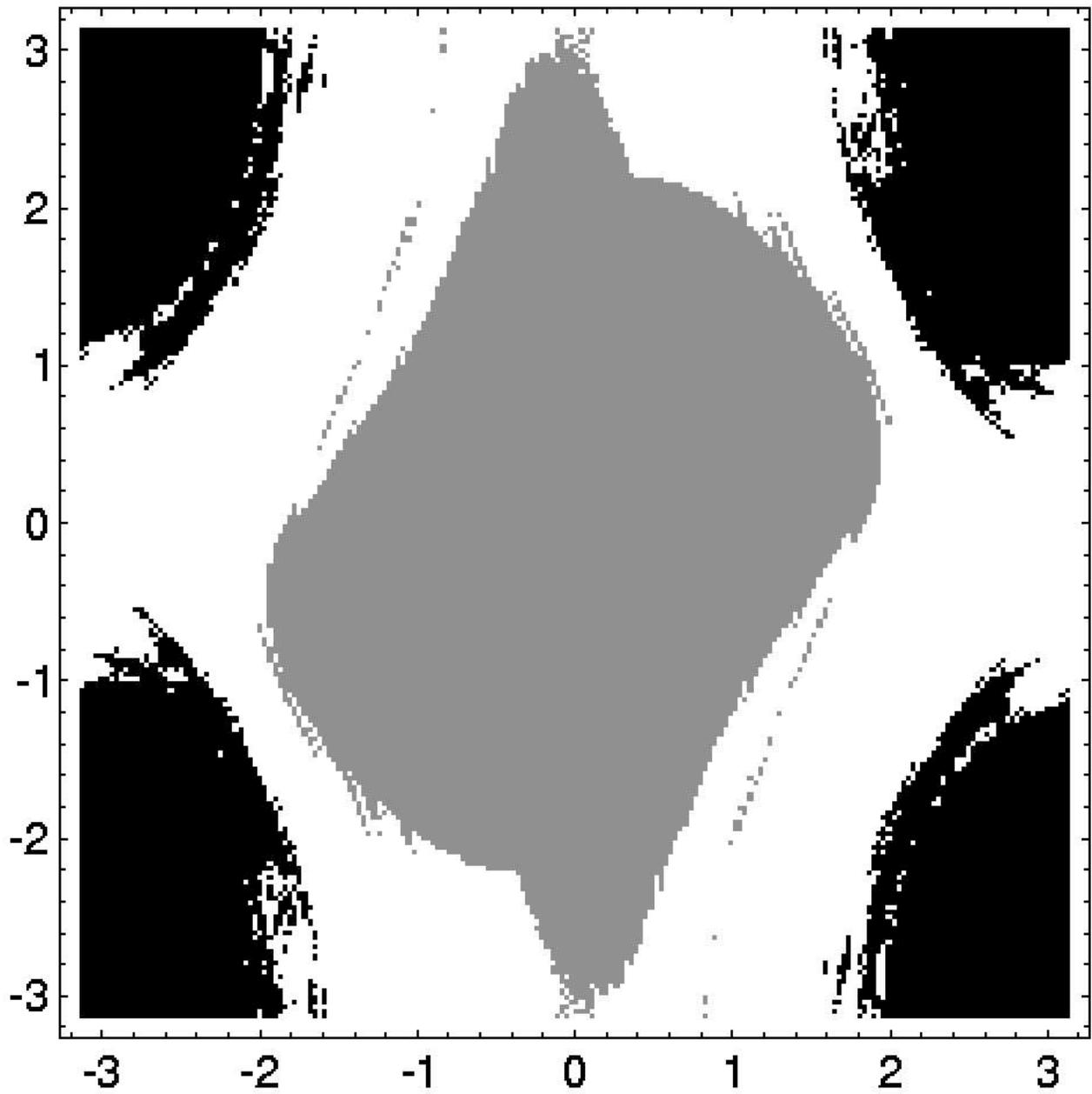


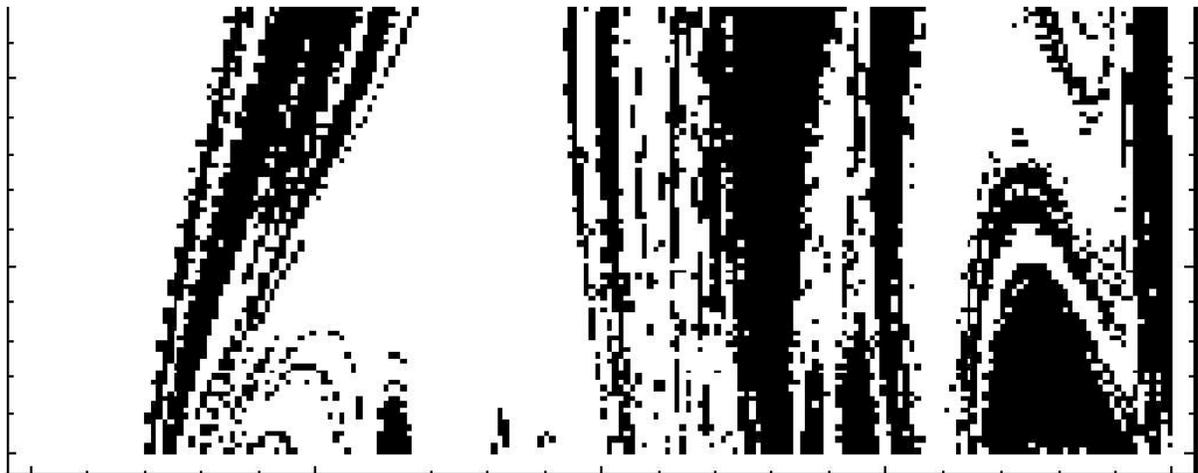
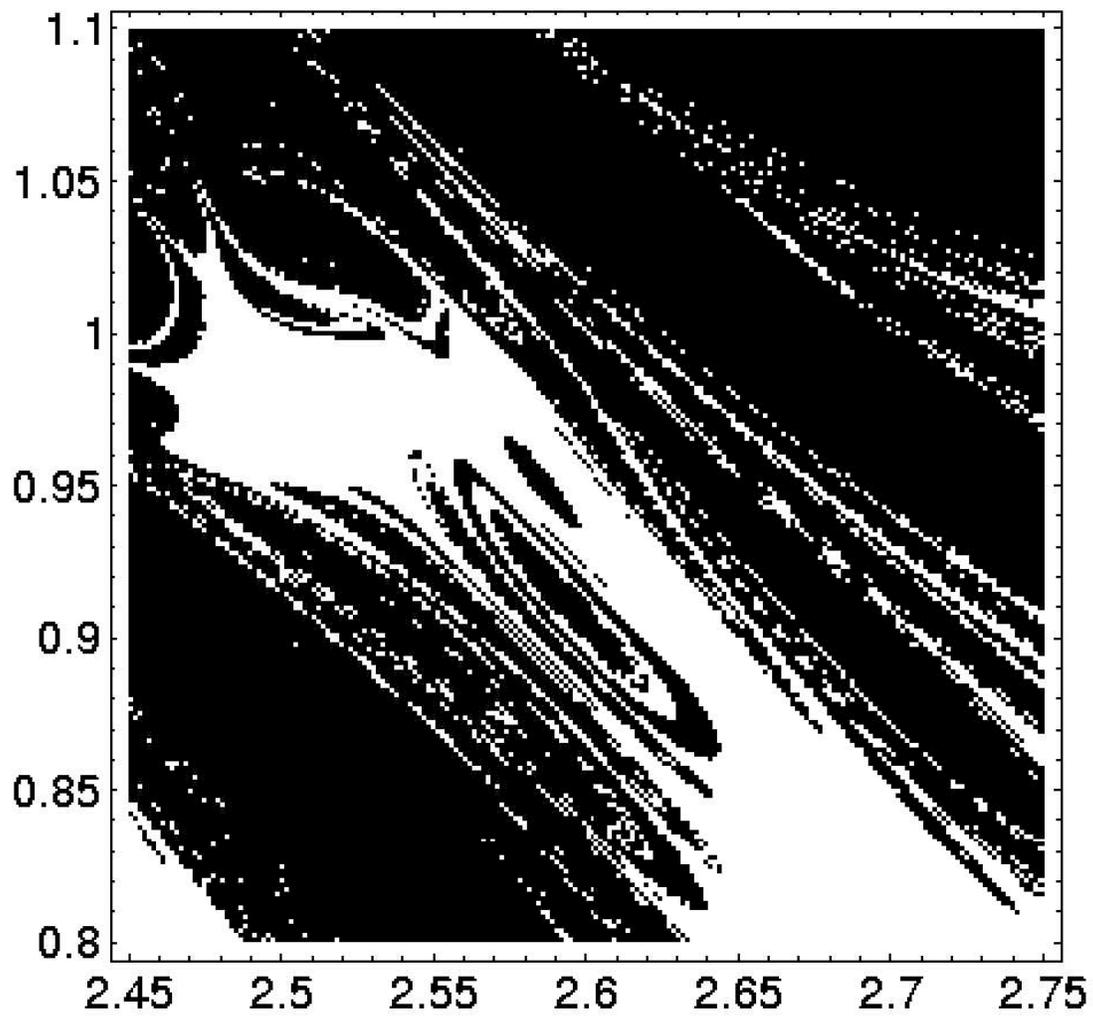
look at initial basins as the spin angles are varied  
find they're fractal



- Black  $\Rightarrow$  merger
- White  $\Rightarrow$  stable
- Grey  $\Rightarrow$  escape

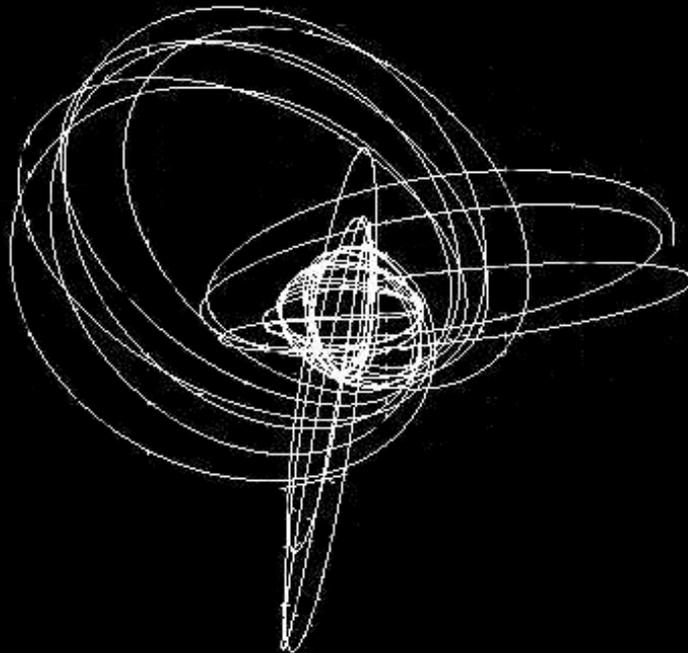
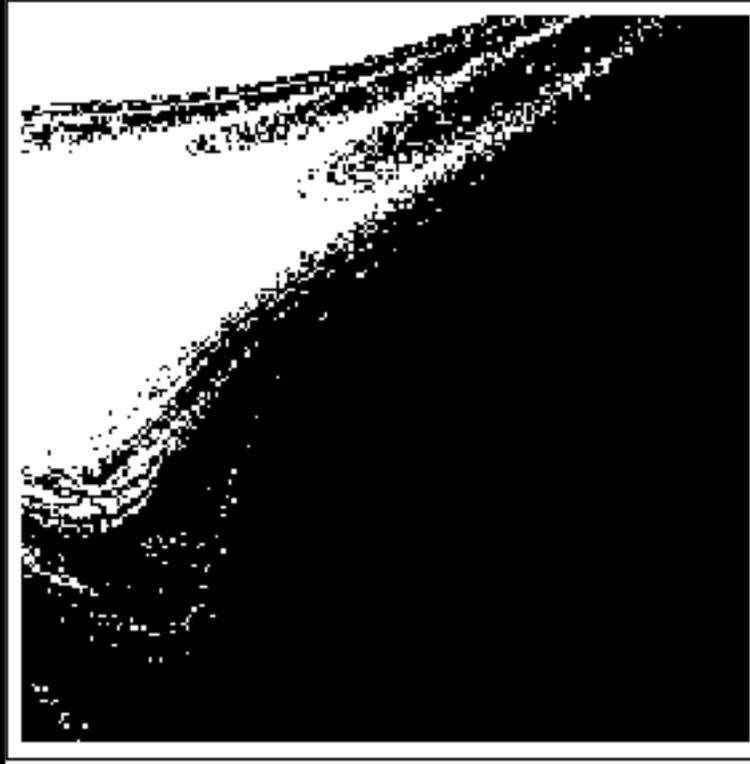
# Fractal Basin Boundary





Chaotic scattering off an a fractal set of  
periodic orbits

$$\frac{m_1}{m_2} = 3, S_1 = 1, S_2 = 0, \theta_1 = 95^\circ, r = 5M$$



# Enough Chaos for Now

## Difficult to Proceed:

1. Chaos in Conservative Dynamics
  2. Damped by Gravitational Waves
  3. Whether completely damped or observable in grey area
- 
- 2PN not relativistic enough
- Kerr with spinning test particle doesn't include mass of companion

## In Brief

First test of GR the precession of the  
perihelion of mercury

Almost a century later, looking for precession  
of multi-leaf orbits

Zoom-whirls

Might even observe transition to  
chaos