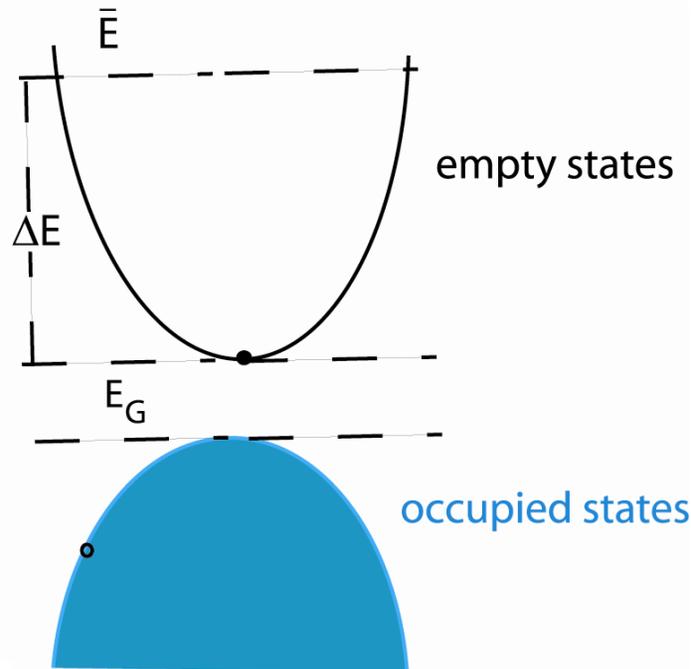


# Photovoltaic effect for narrow-gap Mott Insulators

Efstratios Manousakis  
Florida State University

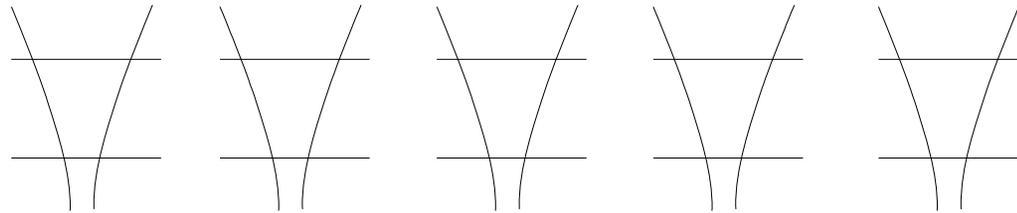
# What is wrong with a band semiconductor



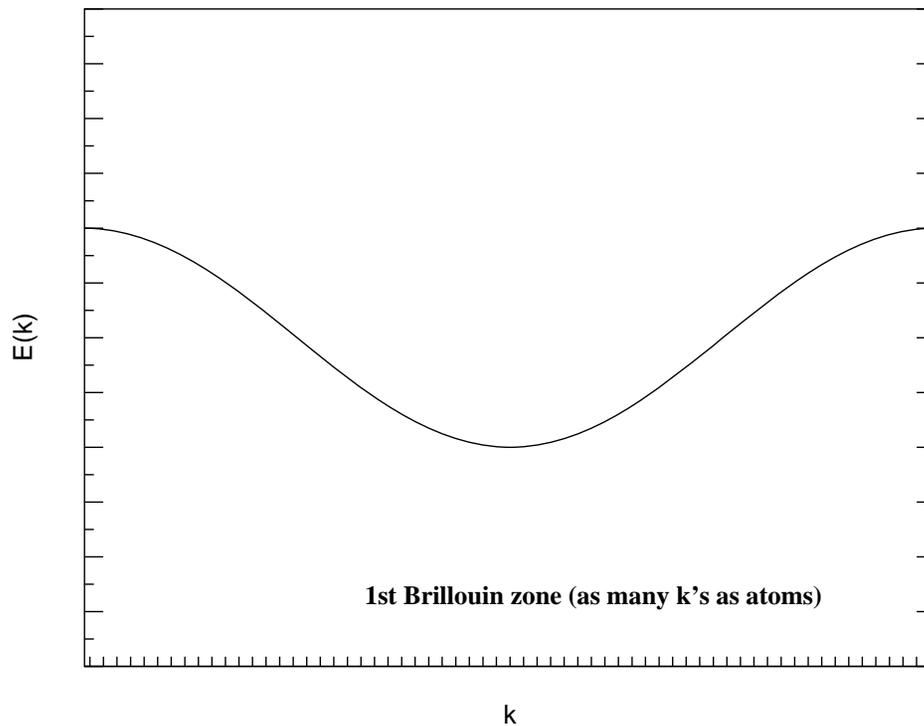
- The average energy range of the incident photon is 0.5 eV-3.5 eV.
- If the gap is chosen to be small, the excess energy  $\Delta E$  beyond the gap of the higher energy photons is wasted into heat via phonon emission or electron-phonon scattering and within  $10^{-12}$  secs the created electrons or holes relax in their band edges.
- If the gap is chosen to be large, it cuts-off most of the solar photons.
- A stack of cascaded multiple p-n junction with various gaps maybe an expensive way to increase efficiency.

# What is a band insulator versus a Mott insulator

N atoms. Each having 2 energy levels.



From each atomic level we have N atomic orbitals giving rise to N Bloch states

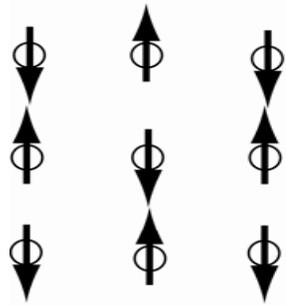


If we consider spin, there are  $2N$  states.

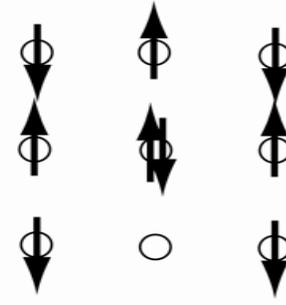
- a) one electron per atomic orbital, leads to a half-filled band (metal).
- b) two electrons per atomic orbital, lead to a filled band (insulator).

In a **Mott insulator**, one electron per atomic orbital leads to an insulator.

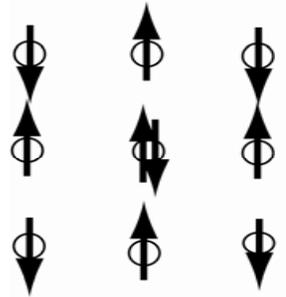
# The Mott insulator



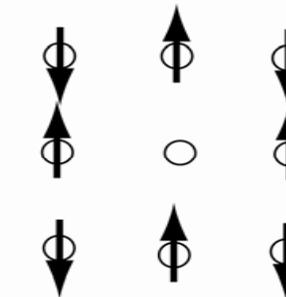
(a)



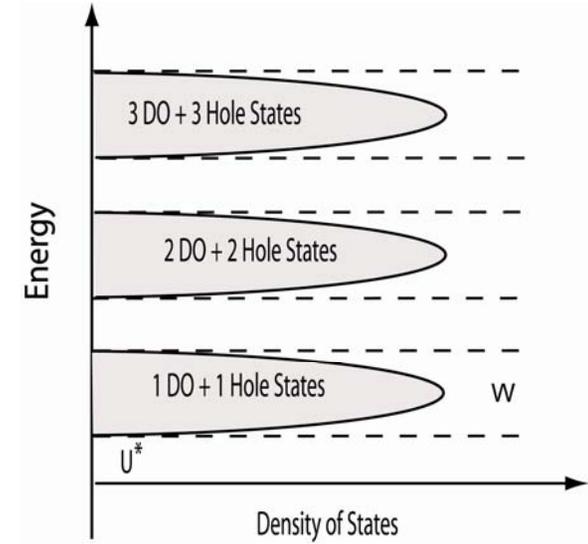
(b)



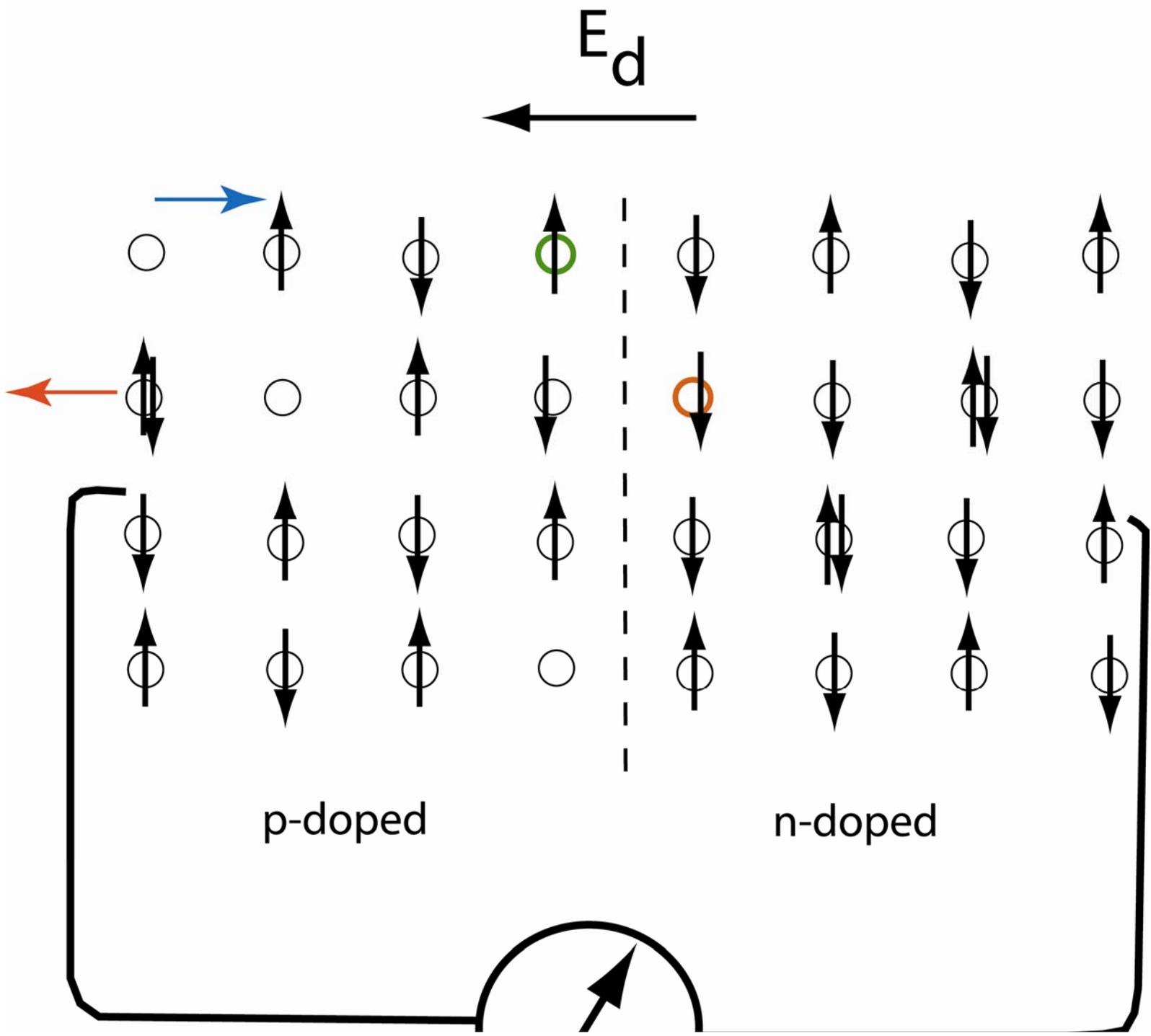
(c)



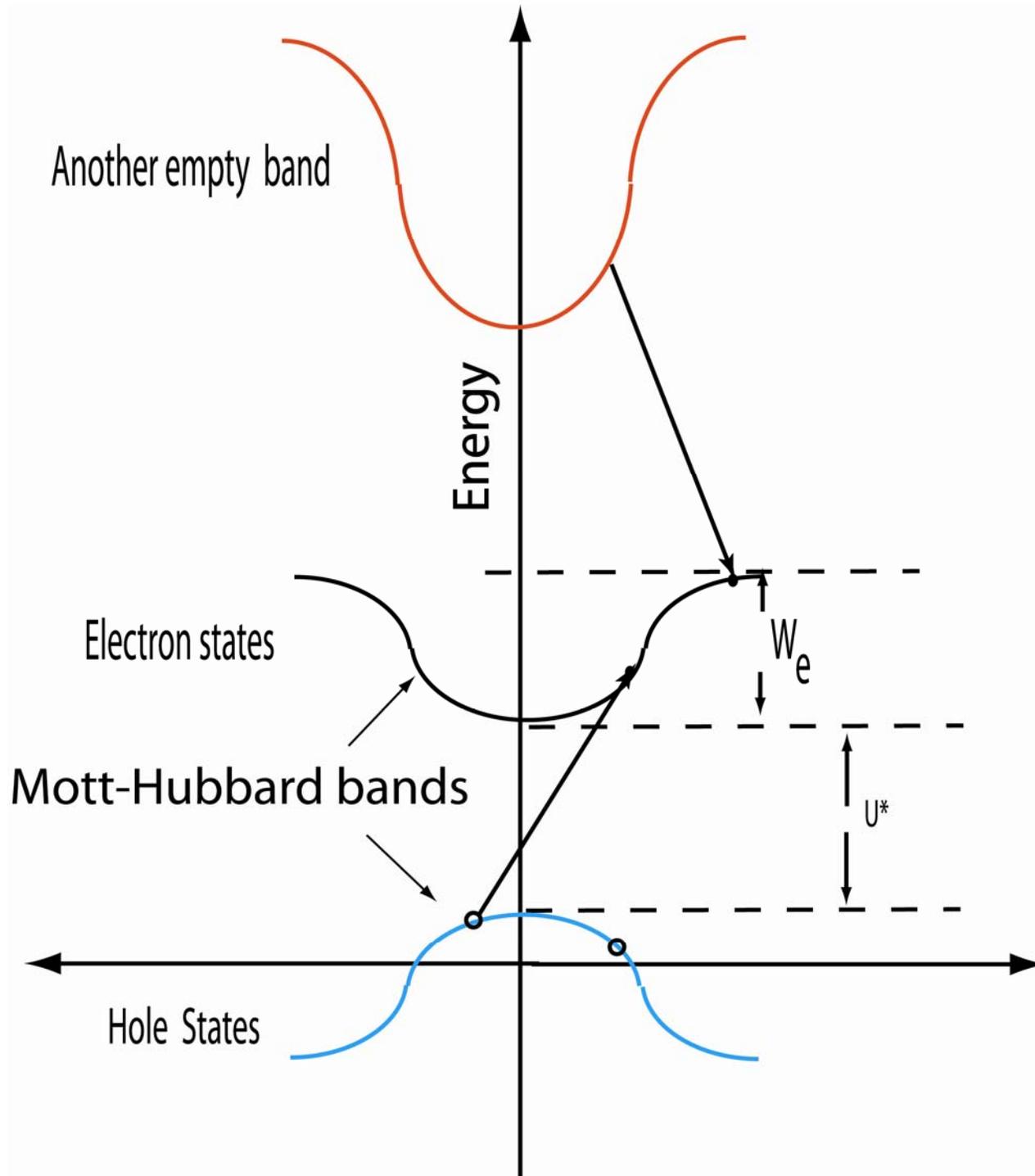
(d)

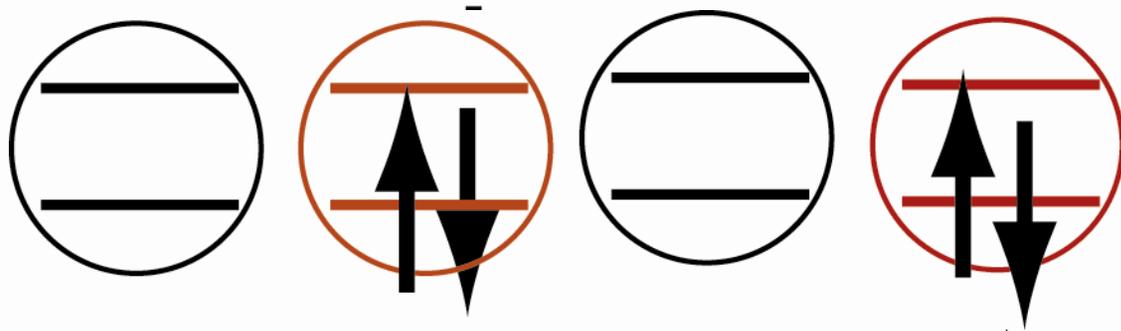


$$\hat{H} = -t \sum_{\langle ij \rangle \sigma} c_{i\sigma}^\dagger c_{j\sigma} + U \sum_i \hat{n}_{i\uparrow} \hat{n}_{i\downarrow}$$

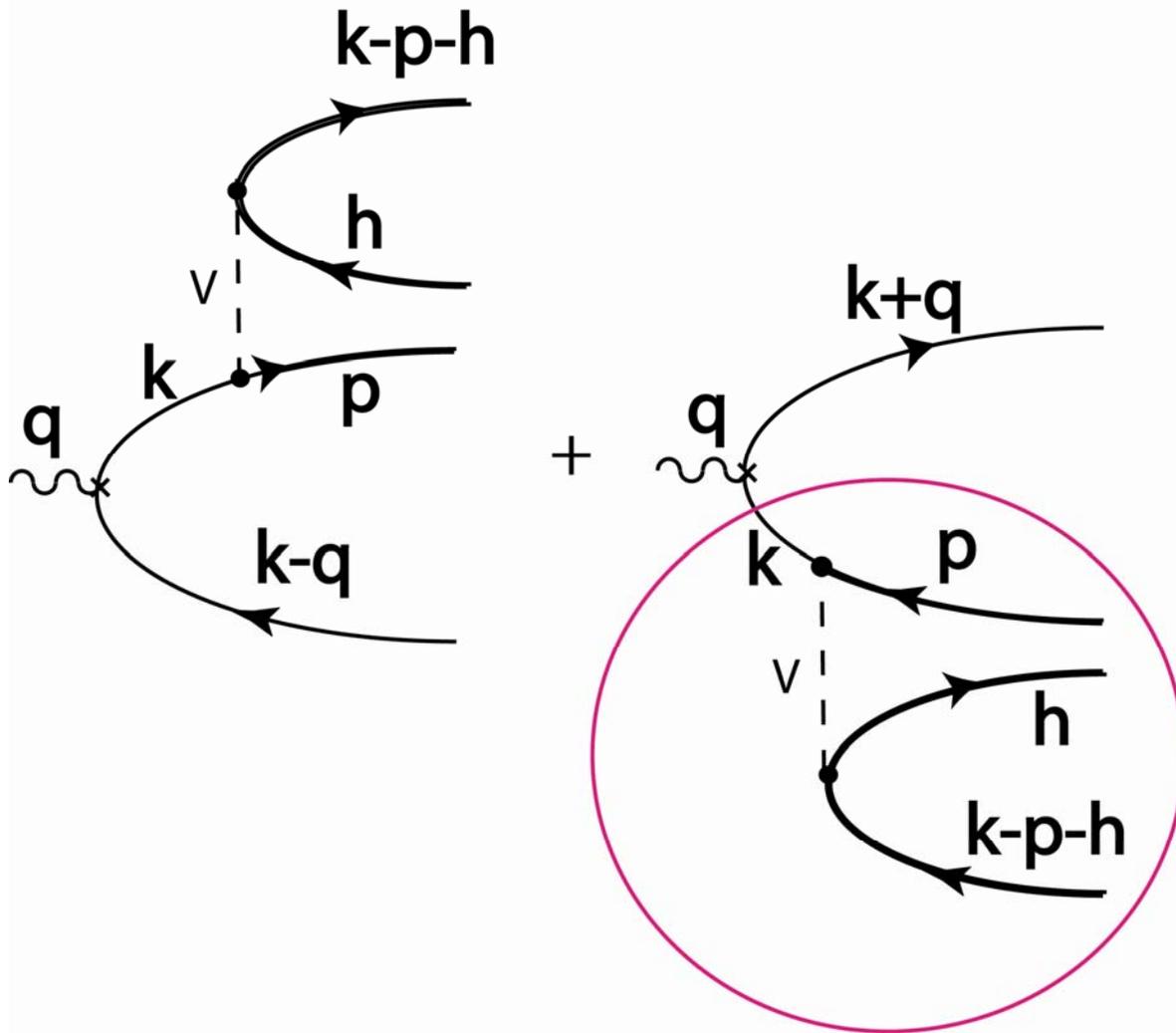


Photovoltaic effect in **narrow-band**  
narrow-gap Mott Insulator





- As a perturbative process



Fermi's golden rule

$$\Gamma_{\mathbf{k}}(\epsilon) = \frac{2\pi}{\hbar} \sum_{\mathbf{p},\mathbf{h}} |M|^2 \delta(\epsilon - E_{\mathbf{k}-\mathbf{p}+\mathbf{h}} - E_{\mathbf{p}} + E_{\mathbf{h}}),$$

$$\epsilon \equiv E - U^*.$$

$M$  is the coupling of the initially excited state  $|\mathbf{k} \nu\rangle$  with the 2qp-1qh state  $\alpha_{\mathbf{h}} \alpha_{\mathbf{p}}^\dagger \alpha_{\mathbf{k}-\mathbf{p}-\mathbf{h}}^\dagger |0\rangle$ ,

$$M = \frac{V}{N} Z_{\mathbf{h}} Z_{\mathbf{p}} Z_{\mathbf{k}-\mathbf{p}+\mathbf{h}},$$

$$V \equiv \langle \mathbf{k} \nu, \mathbf{h} | \hat{V} | \mathbf{p}, \mathbf{k} - \mathbf{p} + \mathbf{h} \rangle,$$

$$Z_p = \langle 0 | \alpha_p | \mathbf{p} \mu \rangle.$$

Here the state  $|\mathbf{k} \nu\rangle$  is the high energy band where the electron or hole is initially excited and  $|\mathbf{p}\rangle$  is the band which becomes the Mott band after including the effects of the Hubbard interaction.

It will be further approximated as

$$\Gamma_{\mathbf{k}}(\epsilon) \simeq \frac{2\pi}{\hbar} V^2 \bar{Z}^6 D(\mathbf{k}, \epsilon),$$

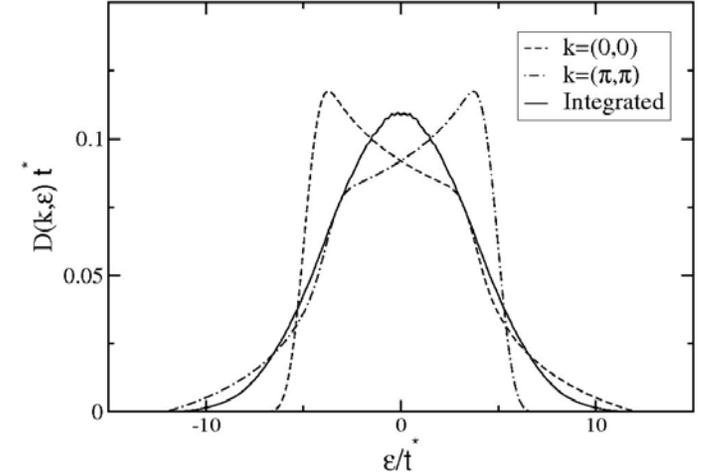
$$D(\mathbf{k}, \epsilon) = \frac{1}{N^2} \sum_{\mathbf{p},\mathbf{h}} \delta(\epsilon - E_{\mathbf{k}-\mathbf{p}-\mathbf{h}} - E_{\mathbf{p}} + E_{\mathbf{h}}).$$

$\bar{Z}$  is the  $Z$  averaged over the Brillouin zone. Similarly

$$\bar{\Gamma}(\epsilon) \simeq \frac{2\pi}{\hbar} V^2 \bar{Z}^6 \Delta(\epsilon),$$

$$\Delta(\epsilon) = \frac{1}{N} \sum_{\mathbf{k}} D(\mathbf{k}, \epsilon).$$

$$E_{\mathbf{k}} = -2t^* \sum_{\mu=1}^d \cos(k_{\mu} a_{\mu})$$



$$\Delta \square 1/W$$

The value of  $Z$  is calculated using the strong  $U$  limit of Hubbard i.e, the t-J model

Take  $V=U=1$  eV,  $t=1/4$

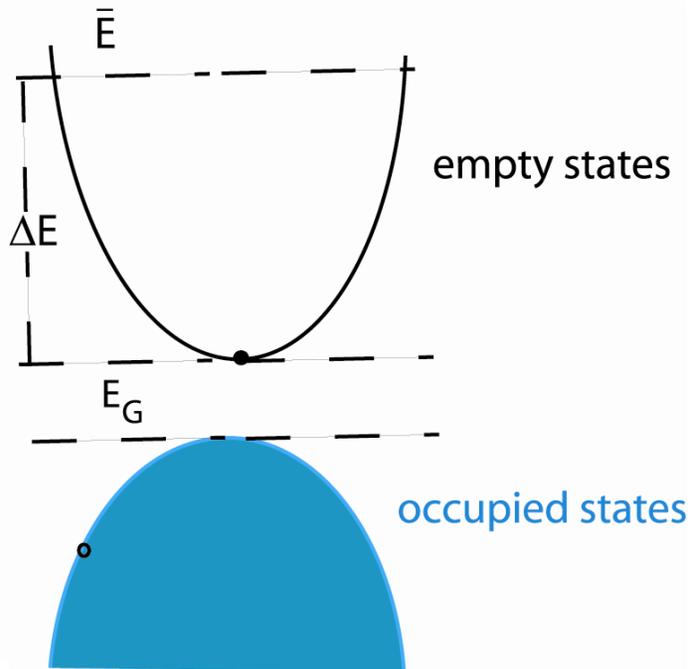
For these value

$J/t=4t/U=1$  and  $W=0.96 t$ ,  $Z=0.53$

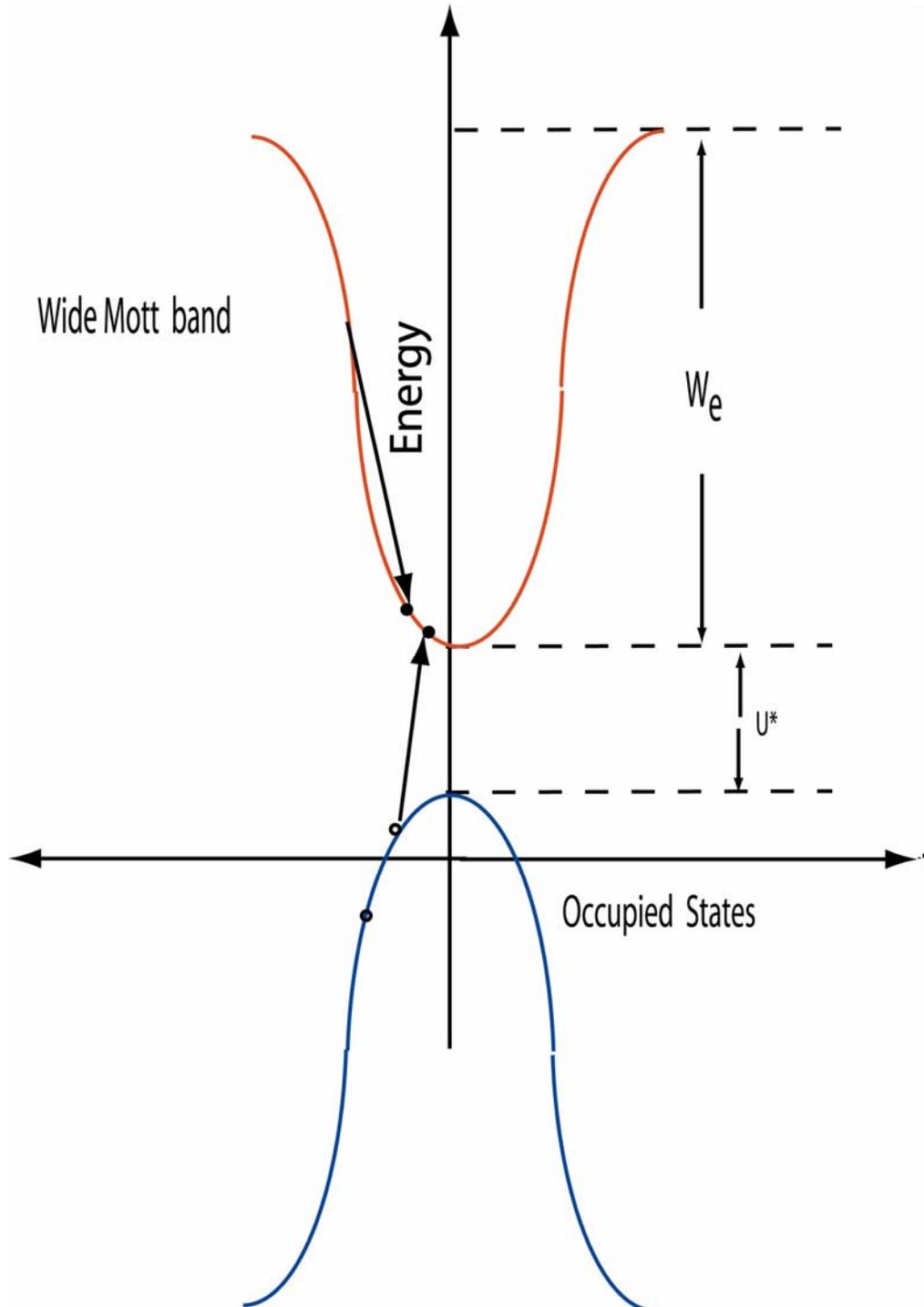
i.e.  $W=0.24$  eV  $\Gamma=0.9 \times 10^{15} \text{ sec}^{-1}$

which is much larger than decay via phonon emission or phonon scattering.

# What is wrong with a band semiconductor



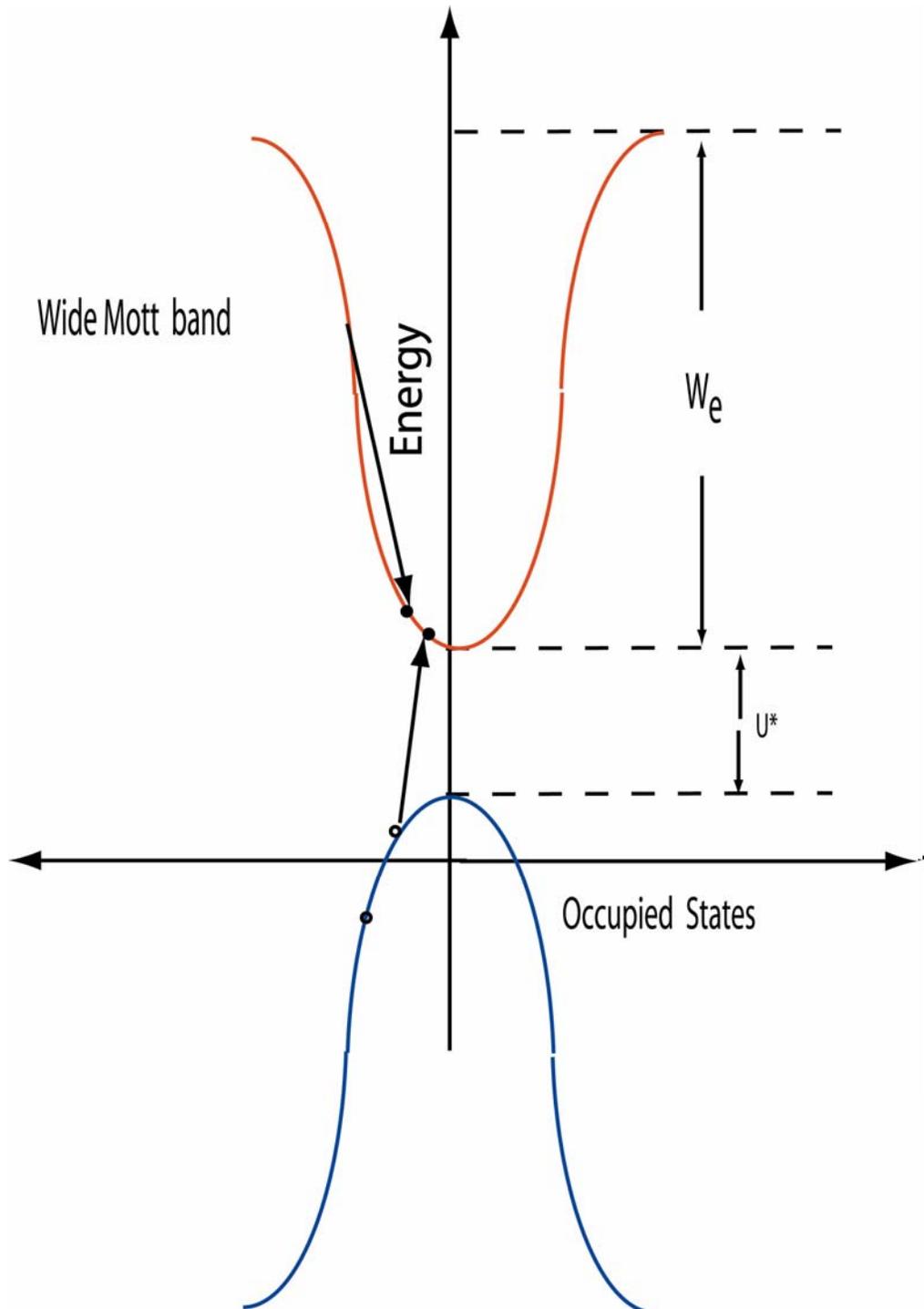
The excess energy  $\Delta E$  beyond the gap of the higher energy photons is wasted into heat via phonon emission or electron-phonon scattering and within  $10^{-12}$  secs the created electrons or holes relax in their band edges.



This process takes only  $10^{-14}$ - $10^{-15}$  secs.

So, before the phonons find enough time to take away the excess energy, it is converted to multiple doublon/hole pairs

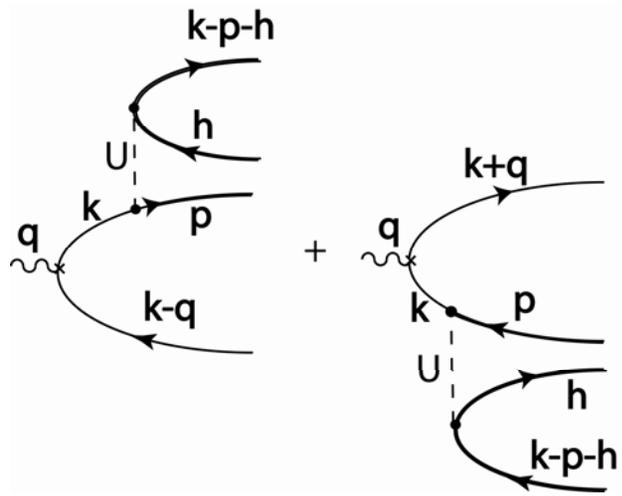
# Photovoltaic effect in **wide-band** narrow-gap Mott Insulators



Narrow-gap Wide-band  
Mott insulators.

When the energy  $E$  of the photo-excited electron is greater than  $nU^*$ , the electron can decay into  $n+1$  electrons plus  $n$  holes by means of the on-site Coulomb interaction  $U$ .

## Narrow-gap Wide-band Mott insulators.



*Narrow-gap wide-band case:  $W \gg U^*$*

For any value of  $W$  we can tune  $U$  to get a small gap  $U^*$  and  $U > W/2$ .

For  $W = 8$  eV and  $U^* = 0.5$  eV we need a  $U > 4$  eV.

The decay rate can be approximated by

$$\hbar\Gamma \sim 2\pi \frac{U^2}{W} \bar{Z}^6$$

For  $W = 8$  eV and  $U > 4$  eV, we obtain (using  $\bar{Z} \sim 0.5$ )

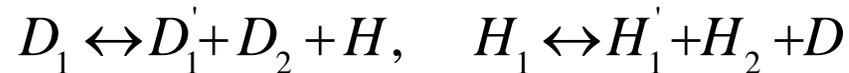
$$\Gamma \sim 0.5 \times 10^{15} \text{sec}^{-1}$$

## Issues regarding efficiency

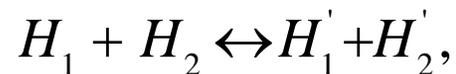
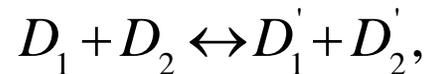
$$J_{absorbed} = \frac{\Omega_s}{4\pi^3 \hbar^3 c^2} \int_{U^*}^{\infty} d\epsilon \frac{\epsilon^3}{\exp(\frac{\epsilon}{k_B T}) - 1}$$

$T=5760$  K the average temperature on Sun's surface,  $\Omega_s$ =Solid angle by which we see the Sun.

### Impact Ionization and Auger recombination



### Carrier-carrier scattering



### Photon absorption and recombination (photo-luminescence)



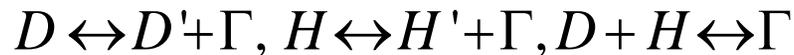
## Equilibrium state

$$J_{emitted} = J_{absorbed},$$

$$J_{emitted} = \frac{\Omega_e}{4\pi^3 \hbar^3 c^2} \int_{U^*}^{\infty} d\epsilon \frac{\epsilon^3}{\exp\left(\frac{\epsilon - \mu_\gamma}{k_B T_e}\right) - 1}.$$

$T_e$  is the cell's temperature,  $\mu_\gamma = 0$  is the chemical potential of the D-H system in equilibrium with photons

Since all the processes involving only electronic degrees of freedom are much faster than the relaxation processes involving **phonons**



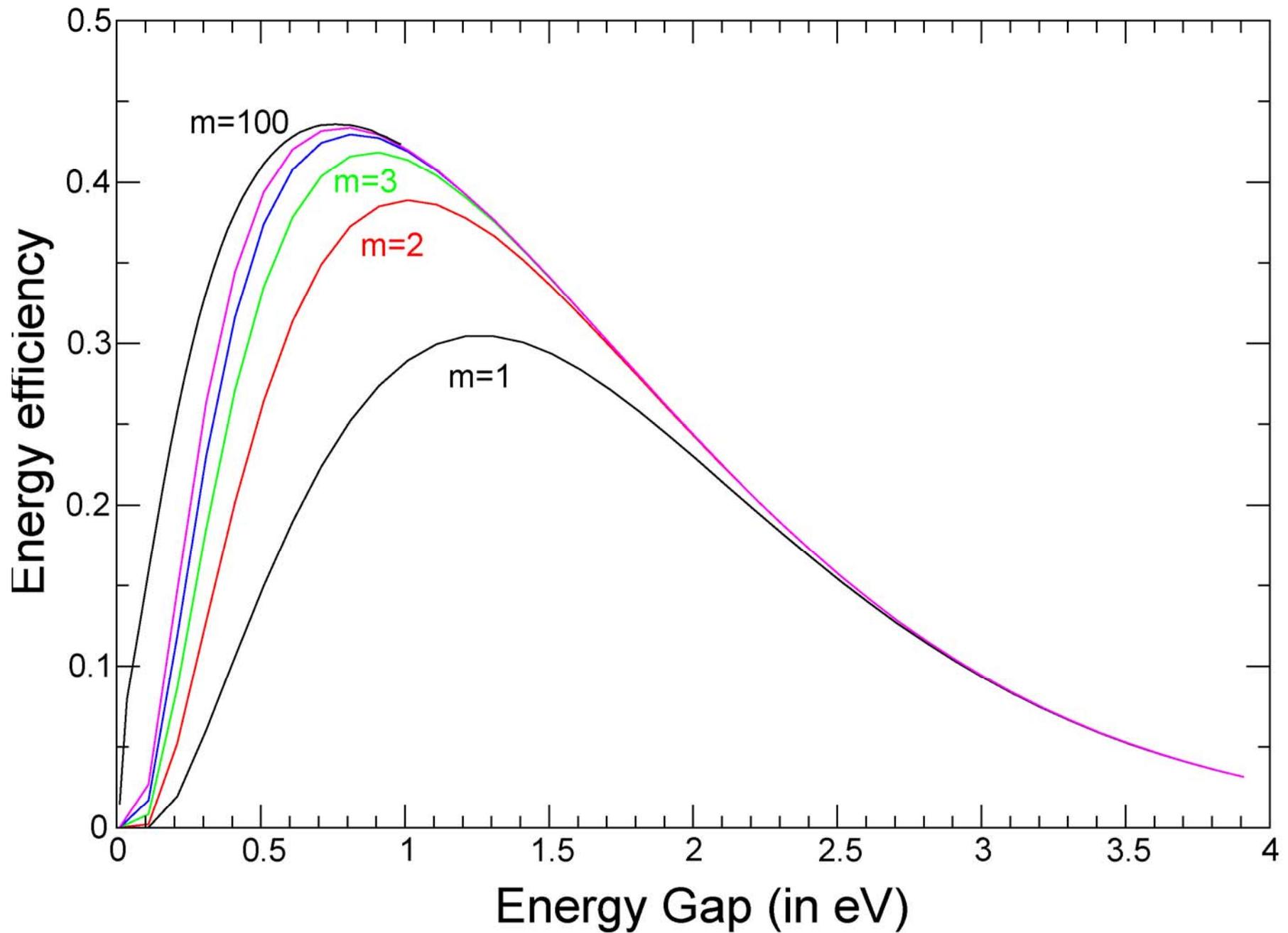
The **maximum efficiency** under these conditions can be calculated as the maximum of the following quantity  $\eta$  as a function of the voltage  $V$  and the Mott-Hubbard gap  $U^*$

$$\eta(V, U^*) = qV \frac{g(U^*) - \xi r(V, U^*)}{P_{in}}, \quad P_{in} = \int_0^{\infty} d\epsilon \frac{\epsilon^3}{\exp(\epsilon/k_B T_s) - 1}$$

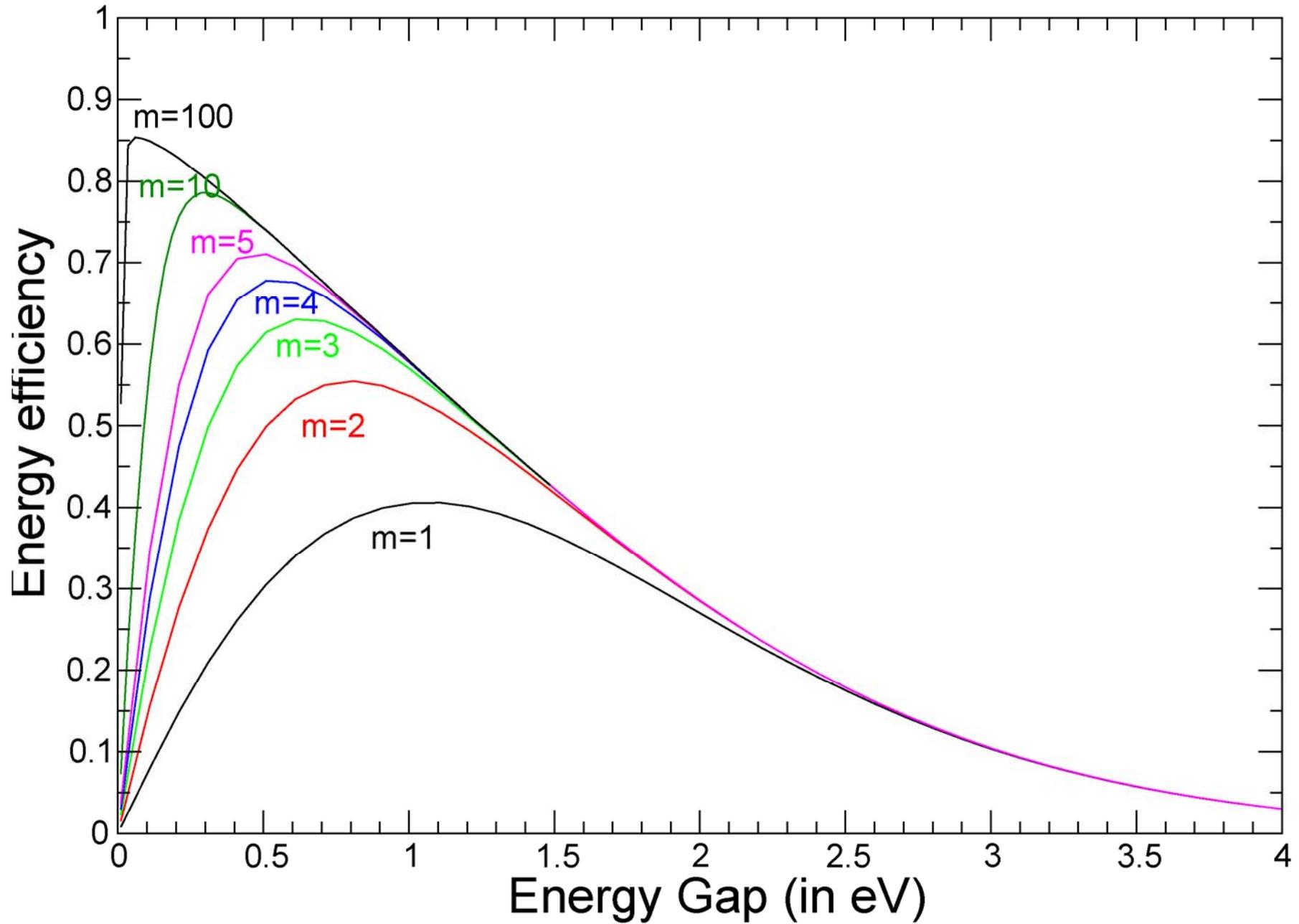
$$g(U^*) = \int_{U^*}^{\infty} d\epsilon \frac{m(\epsilon)\epsilon^2}{\exp(\epsilon/k_B T_e) - 1}, \quad r(V, U^*) = \int_{U^*}^{\infty} d\epsilon \frac{m(\epsilon)\epsilon^2}{\exp((\epsilon - qV)/k_B T_e) - 1}$$

$m(\epsilon)$  is the smaller of the following (a) integer part of  $\epsilon/U^*$ , and (b) a maximum allowed value  $M_{max}$ , and  $\xi = \pi/\Omega_s$  or 1 for fully concentrated sunlight.

# Un-concentrated sunlight

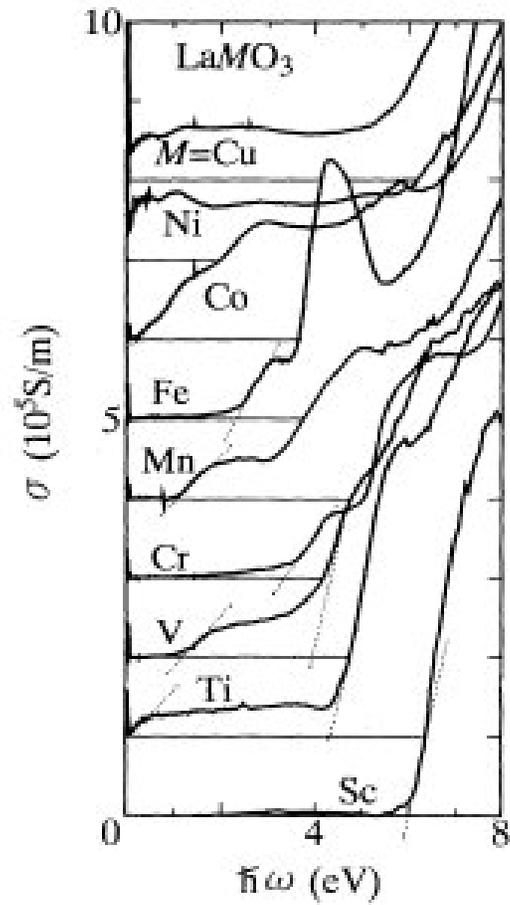


# Fully concentrated sunlight



# Photovoltaic effect on Perovskite heterojunctions

- J. R. Sun, C. M. Xiong, B. G. Shen, P. Y. Wang, and Y. X. Weng, Appl. Phys. Lett. **84**, 2611 (2004).
- J. Qiu, H.-B. Lu, K.-J. Jin, M. He, and J. Xing, Physica B, **400**, 66 (2007).
- H. Liu, K. Zhao, N. Zhou, H. Lu, N. He, Y. Huang, K.-J. Jin, Y. Zhou, G. Yang, S. Zhao, and A. Wang, Appl. Phys. Lett. **93**, 171911 (2008).
- Z. Luo, J. Gao, A. B. Djurisc, C. T. Yip, and G. B. Zhang, Appl. Phys. Lett. **92**, 182501 (2008).
- H.-B Lu, K.J Jin, Y.-H. Huang, M. He, K. Zhao, B.L. Cheng, Z.-H. Chen, Y.-L. Zhou, S.-Y. Dai, and G.-Z. Yang, Appl. Phys. Lett. **86**, 241915 (2005).
- K. Zhao, K.-J. Jin, H. Lu, Y. Huang, Q. Zhou, M. He, Z. Chen, Y. Zhou, and G. Yang, Appl. Phys. Lett. **88**, 141914 (2006).
- Y. Muraoka, T. Muramatsu, J. Yamaura, and Z. Hiroi, Appl. Phys. Lett. **85**, 2950 (2004).
- J. Xing, K. Zhao, G. Z. Liu, M. He, K. J. Jin, and H. B. Lu, J. Phys. D: Appl. Phys. **40**, 5892 2007;
- N. Zhou, K. Zhao, H. Liu, H. B. Lu, M. He, S. Q. Zhao, W. X. Leng, A. J. Wang, Y. H. Huang, K. J. Jin, Y. L. Zhou, and G. Z. Yang, *ibid.* J. Phys. D: Appl. Phys. **41**, 155414 2008.



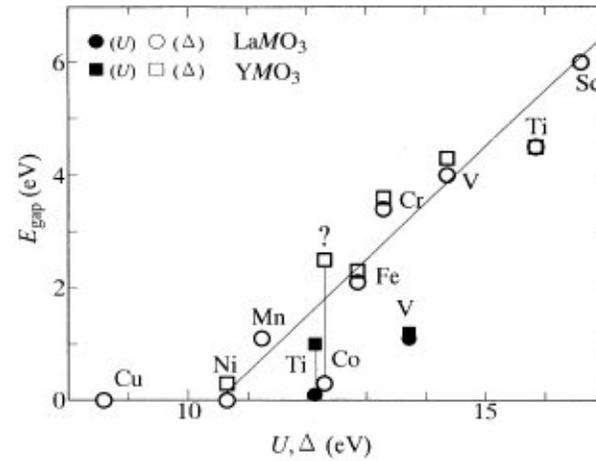
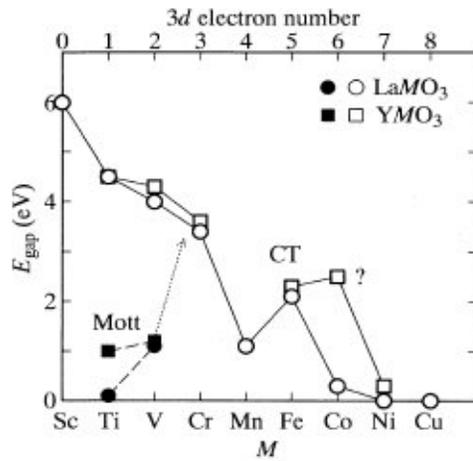
Optical conductivity  
and optical gaps in some  
perovskite materials

# Gaps in Perovskites

17 008

T. ARIMA, Y. TOKURA, AND J. B. TORRANCE

48



We are focusing our attention to the following class of materials



We would like to study optical properties of these materials  
Using hybrid functionals and many-body perturbation theory.

Maitri Warusawithana of NHMFL using MBE is growing  
epitaxially  $\text{LaVO}_3$  films on  $\text{SrTiO}_3$

We plan to carry out photoconductivity studies of films of such materials.

Fast optical pump and probe techniques should probe the decay products  
of photo-excited electron/hole pairs.

Density functional theory studies of films of the above class of materials.

# Conclusions

- When a solar photon excites an electron-hole pair above the Mott-Hubbard gap in a narrow-gap **narrow-band** Mott insulator the photo-excited electron or hole which temporarily lands on a **different high energy band** can decay quickly into  $n+1$  quasi-electrons plus  $n$  quasi-holes from the Mott band or  $n+1$  quasi-holes plus  $n$  quasi-electrons. The time scale for this energy conversion is shorter than the one required for electron/hole relaxation through phonons or spin wave excitations.
- In the case of narrow-gap **wide-band** Mott insulators the excess energy of the photo-excited electron (or hole) is converted into  $(n+1)$ -quasi-electrons plus  $n$  quasi-holes (or  $n+1$  quasi-holes plus  $n$  quasi-electrons) also within short time scales before the electron or hole loses its excess energy via phonon or other excitations.
- Therefore, a solar cell made using a narrow gap Mott insulator for the illuminated side of the p-n junction should lead to an efficient solar cell through the production of a high current.
- E. Manousakis, Phys. Rev. B **82**, 125109 (2010).