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## NEOTωIST – DESIGN STUDY OF A KINETIC IMPACTOR DEMONSTRATION MISSION FEATURING NEO SPIN CHANGE AND OBSERVER SUB-SPACECRAFT

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NEOShield-2 Science and Technology for Near-Earth Object Impact Prevention

## Outline

- Introduction
- Mission Architecture and Implementation Options
- Observation geometry
- Impactor concept
- Flyby module concept
- Chaser concept
- Concluding remarks



## Introduction

### **NEOTωIST Mission Feasibility Study**

Alternative Impactor Demonstration Mission that promises reduced cost, programmatic flexibility & high value



Higher probability of implementation

## Scientific/programmatic objectives and measurement principle

as explained by previous speaker

#### **Execute and NEOShield-2 Project**

EC-funded activity to study various aspects of Asteroid threat mitigation executed by Europe-wide consortium of Industry & Science Institutes



Tunguska, 1908



Chelyabinsk, 2013



#### Somewhere, sometime

### **Mission Architecture**

#### Concept

- One impactor spacecraft also functioning as interplanetary carrier
- One flyby sub-spacecraft (Flyby Module / FMB) ejected prior to impact
- FBM passes at safe distance functioning as observation platform and communications node
- One or two Chaser sub-spacecraft ejected prior to impact
- Chasers follow impactor trajectory and are potentially destroyed after performing observation tasks
- Observation of impact from different vantage points
- Spin change detection by observation from Earth





## Mission Architecture & Implementation Options

			Vehicle	Functions	Information generated
01–	02-	03-	Impactor	<ul><li>End game GNC</li><li>Interplanetary carrier</li></ul>	<ul> <li>Impact location (Navigation data)</li> </ul>
			Flyby module	<ul> <li>Observation of ejecta geometry (camera + radiometer)</li> <li>Reception &amp; downlink of data from all S/C</li> </ul>	<ul> <li>Direction of overall ejecta cloud vector</li> <li>Ejecta dynamics &amp; optical density</li> <li>Radiometry of blast</li> </ul>
			Chaser(s)	<ul> <li>Observation of impact crater</li> <li>(Long range observation of ejecta)</li> </ul>	<ul> <li>Constraint on ejecta volume / crater volume</li> </ul>



### AIRBUS

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## Observation geometry of flyby

Flyby observation geometry defined by offset distance and timing of passage with respect to impact time

#### Constraints

- Sensitivity of field of view to uncertainty of size of ejecta cloud
- Variation of apparent size of ejecta cloud over flyby
- Image tracking azimuth rate
- Sensitivity of observation parameters to trajectory errors
- Size of required optics

#### Selected parameters

- d-offset: ~70 km
- Time of closest passage @ impact + 2s
- FOV: 8.35°
- Max azimuth rate: 7deg/sec
- Pointing targets "middle" of predicted ejecta cloud (offset from Impact point)



FOV sensitivity to ejecta velocity assumptions



## Visualisation of Flyby Observation







Impact + 0.8 s

Impact + 2 s (Closest approach, perfect observability in one plane) Impact + 10 s (partial observability of second plane)

- **Green** = Near side of ejecta cone
- **Red** = Far side of ejecta cone

## Visualisation of flyby observation



Green = Near side of ejecta cone Red =

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Far side of ejecta cone

## Impactor and stack overview

- Small Mission Module based on updated Don Quijote design
- Propulsion using slightly modified existing propulsion module from Lisa Pathfinder
- Propulsion module contributes to impacting mass
- Visual GNC for final approach
- Launch on VEGA

Vehicle stack mass synopsis			
	Dry [kg]	Wet [kg]	
Fly-by Module	104.1	104.6	
Chaser	25.7	26.2	
Impactor Mission Module	263	263	
Propulsion module (incl. resid.)	280	1420	
Stack	673	1814	
Stack at launch incl. adapter		1929	
Delivered mass w/o prop. mod.		394	
Min. impacting mass	543		



# Flyby module drivers & payload

### **Design drivers**

- High quality imaging of ejecta cloud during high velocity pass
- Imaging geometry in light of uncertainties in approach trajectory
- Data buffering & transmission for entire constellation
- Size and mass constraints imposed by mission concept
- Mission cost
- Required reliability
- Deep space environment



# **Flyby Module Navigation**

- <u>Challenge</u>: Point the FMB camera at the target accurately despite uncertainties in the flyby trajectory (mainly offset distance)
- <u>Sources of uncertainty</u>: Errors in impactor trajectory at separation; and errors in separation delta-V
- <u>Solution</u>: relative navigation during approach using payload camera



## Flyby module configuraton





Overview

Equipment accommodation



# Flyby module preliminary specs

Item	Description
Mass	Dry/ Wet: 104.1 / 104.6 kg
Power	DC: 100 W
	S/A : 0.75 m <sup>2</sup>
	Battery: 150 Wh (usable)
Dimensions	Bus structure: 100 x 100 x 25 cm <sup>3</sup>
Power and data	PROBA-NEXT avionics integrated power and data handling, >1 Gbit mass
handling	memory
Comms.	Earth-link: X-band, 1m HGA, 20 W RF, 50 kb/s
	Impactor link: X-band, MGA, Rx only, < 10 Mbit/s
	Chaser link: X-band, MGA, Rx only, ~ 20 Mbit/s
AOCS	3-axis stabilised
	2 x STR (DTU Micro ASC)
	2x Coarse sun sensors
	3 x Micro-wheels, 0.42 Nms (e.g. SSTL)
	IMU (DTU, int. with STR electronics)
	RCS: 4 x cold gas thrusters for momentum management (0.5 kg of $N_2$ )

# Chaser concept

#### **Design drivers**

- Small size of main target (crater)
- Uncertainty in imaging geometry
- Real-time transmission of imaging data before vehicle destruction
- Large delta-V for separation from Impactor
- Mission cost and mass constraints
- Deep space environment



# Chaser concept

Item	Description	
Mass	Dry/Wet: 25.7 / 26.2 kg	
Power	DC: 40 W	
	Battery: 60 Wh (usable)	
Dimensions	20 x 30 x 10 cm <sup>3</sup> (TBC)	
Payload	Medium angle camera	
	Field of view: 16°	
	Aperture: still to be determined	
	Detector: 2048 x 2048 pixel	
	Max. resolution: 4m @ 30 km, 2m @ 15 km	
	No active target tracking	
Power & data	data Cube-sat/ small sat equipment (details still to be selected)	
Comms.	Link to FBM: X-band, MGA, TX only, ~ 20 Mbit/s	
AOCS	Stabilization along velocity direction with single uncontrolled momentum	
	wheel, spun up before ejection from Impactor	
Propulsion	Hydrazine, single thruster in anti-velocity direction, stabilization with	
	momentum wheel, dV capability 70 m/s	
Thermal	Heaters & radiators, non-stationary design for terminal phase possible	

# **Concluding remarks**

NEOTωIST Mission Concept addresses objectives of Kinetic Impactor Demo Mission at reduced cost and with flexible implementation options

#### Advantages achieved by

- New measurement principle for quantifying the achieved momentum transfer to target
- Replacing large rendezvousing reconnaissance spacecraft with small sub-spacecraft

#### Current snapshot of feasibility work shows no show stopper

#### Future work & iterations on

- Refinement of the payload
- Refinement of combined AOCS image targeting concept
- Equipment selection and overall vehicle design
- Refinement of communications and data management





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# Thank you for your attention! Questions?

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### Backup: Demonstration mission objectives

#	Objective	Derived functionality
01	Technology demonstration Kinetic Impactor	Impact target NEO with a spacecraft in hypervelocity regime with sufficient accuracy
02	Technology demonstration of an observer spacecraft for impact verification	Demonstrate observation, from a flyby vehicle, of the impact event with sufficient quality to verify that the impact took place as required for deflection
03	Deflection validation	Measure target NEO orbit or rotation before and after impact to prove transfer of (angular) momentum.
04	β determination / quantification of momentum transfer augmentation from ejecta	Quantify the magnitude of momentum carried by the ejecta
05	Observational data to validate/ improve impact modelling	Measurement of the dynamics and effects of the impact event





Augmentation of imparted momentum caused by ejecta

 $\beta = \frac{p_{KI} + p_{Ej}}{p_{KI} + p_{Ej}}$  $p_{KI}$ 

### **Backup: Measurement principles**



Itokawa target with impact geometry

(O4)  $\beta$  Quantification = Determination of magnitude of ejecta momentum vector (O4)

- Spacecraft strikes target NEO off-center inducing spin rate change
- Spin rate change determined by ground-based observation of brightness curves
- **Geometry of impact and ejecta** determined from apriory knowledge and spacecraft observations
- **NEO mass properties and targeting information** available from previous characterisation (by Hayabusa)

Combination of these parameters allows reconstruction of ejecta momentum vector reconstruction.

(O5) Impact phenomenon observation also achieved by observations by spacecraft

(O1, O2, O3) Technology demonstration & deflection validation achieved implicitly by above observation