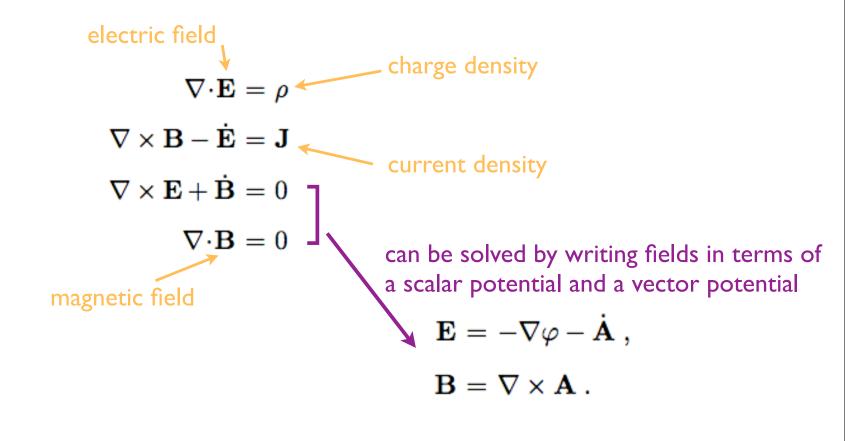
Maxwell's equations

based on S-54

Our next task is to find a quantum field theory description of spin-I particles, e.g. photons.

Classical electrodynamics is governed by Maxwell's equations:



$$\begin{split} \mathbf{E} &= -\nabla \varphi - \dot{\mathbf{A}} \,, \\ \mathbf{B} &= \nabla \times \mathbf{A} \,. \end{split}$$
The potentials uniquely determine the fields, but the fields do not uniquely determine the potentials, e.g. arbitrary function of spacetime $\varphi' &= \varphi + \dot{\Gamma} \,, \\ \mathbf{A}' &= \mathbf{A} - \nabla \Gamma \,, \end{aligned}$
result in the same electric and magnetic fields. The field strength fields the field strength field strength $A^{\mu} \equiv (\varphi, \mathbf{A})$
the field strength $F^{\mu\nu} \equiv \partial^{\mu}A^{\nu} - \partial^{\nu}A^{\mu}$
in components: $F^{0i} = E^{i} \,, \\ F^{ij} &= \varepsilon^{ijk}B_{k} \,. \end{split}$

$$egin{aligned}
abla \cdot \mathbf{E} &=
ho \ \nabla \cdot \mathbf{E} &=
ho \ \nabla imes \mathbf{B} - \dot{\mathbf{E}} &= \mathbf{J} \ \nabla imes \mathbf{B} - \dot{\mathbf{E}} &= \mathbf{J} \ \nabla imes \mathbf{E} + \dot{\mathbf{B}} &= 0 \ \nabla \cdot \mathbf{B} &= 0 \ \end{aligned}$$

The first two Maxwell's equations can be written as:

$$\partial_{\nu}F^{\mu\nu} = J^{\mu}$$
 charge-current density 4-vector $J^{\mu} \equiv (\rho, \mathbf{J})$

taking the four-divergence:

$$\partial_{\mu}\partial_{\nu}F^{\mu
u} = \partial_{\mu}J^{\mu}$$
 $F^{\mu
u} = -F^{
u\mu}$

we find that the electromagnetic current is conserved:

$$\dot{\rho} + \nabla \cdot \mathbf{J} = 0$$

 $\partial_{\mu}J^{\mu} = 0$

The last two Maxwell's equations can be written as:

$$\varepsilon_{\mu\nu\rho\sigma}\partial^{\rho}F^{\mu\nu} = 0$$

automatically satisfied!

The gauge transformation in four-vector notation: $\varphi' = \varphi + \dot{\Gamma}$, $A^{\mu} \equiv (\varphi, \mathbf{A})$ $A'^{\mu} = A^{\mu} - \partial^{\mu}\Gamma$ $\mathbf{A}' = \mathbf{A} - \nabla\Gamma$,

The field strength transforms as:

 $F^{\prime\mu
u} = \partial^{\mu}A^{\prime
u} - \partial^{
u}A^{\prime\mu}$

$$F^{\prime\mu\nu} = F^{\mu\nu} - (\partial^{\mu}\partial^{\nu} - \partial^{\nu}\partial^{\mu})\Gamma$$

 $F'^{\mu\nu} = F^{\mu\nu}$

= 0 (derivatives commute)

the field strength is gauge invariant!

Next we want to find an action that results in Maxwell's equations as the equations of motion; it should be Lorentz invariant, gauge invariant, parity and time-reversal invariant and no more than second order in derivatives; the only candidate is:

$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu}F_{\mu\nu} + J^{\mu}A_{\mu}$$

 $S = \int d^4\!x\, {\cal L}$

we will treat the current as an external source

$$S = \int d^{4}x \mathcal{L}$$

$$\mathcal{L} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} + J^{\mu} A_{\mu}$$
obviously gauge invariant
$$J^{\mu} (A'_{\mu} - A_{\mu}) = -J^{\mu} \partial_{\mu} \Gamma$$

$$= -(\partial_{\mu} J^{\mu}) \Gamma - \partial_{\mu} (J^{\mu} \Gamma)$$

$$\partial_{\mu} J^{\mu} = 0$$
total divergence
In terms of the gauge field:
$$F^{\mu\nu} = \partial^{\mu} A^{\nu} \partial_{\mu} A_{\nu} + \frac{1}{2} \partial^{\mu} A^{\nu} \partial_{\nu} A_{\mu} + J^{\mu} A_{\mu}$$

$$= +\frac{1}{2} A_{\mu} (g^{\mu\nu} \partial^{2} - \partial^{\mu} \partial^{\nu}) A_{\nu} + J^{\mu} A_{\mu} - \partial^{\mu} K_{\mu}$$

$$K_{\mu} = \frac{1}{2} A^{\nu} (\partial_{\mu} A_{\nu} - \partial_{\nu} A_{\mu})$$
equations of motion:
$$(g^{\mu\nu} \partial^{2} - \partial^{\mu} \partial^{\nu}) A_{\nu} + J^{\mu} = 0$$
equivalent to the first two Maxwell's equations!
$$\partial_{\nu} F^{\mu\nu} = \partial_{\nu} (\partial^{\mu} A^{\nu} - \partial^{\nu} A^{\mu}) = (\partial^{\mu} \partial^{\nu} - g^{\mu\nu} \partial^{2}) A_{\nu}$$

Electrodynamics in Coulomb gauge

based on S-55

Next step is to construct the hamiltonian and quantize the electromagnetic field ...

$$\begin{aligned} \mathcal{L} &= -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} + J^{\mu} A_{\mu} \\ &= -\frac{1}{2} \partial^{\mu} A^{\nu} \partial_{\mu} A_{\nu} + \frac{1}{2} \partial^{\mu} A^{\nu} \partial_{\nu} A_{\mu} + J^{\mu} A_{\mu} \end{aligned}$$

Which A_{μ} should we quantize?

too much freedom due to gauge invariance

There is no time derivative of A^0 and so this field has no conjugate momentum (and no dynamics).

To eliminate the gauge freedom we choose a gauge, e.g.

$$\nabla \cdot \mathbf{A}(x) = 0$$

Coulomb gauge

an example of a manifestly relativistic gauge is Lorentz gauge:

 $\partial^{\mu}A_{\mu} = 0$

 $\nabla \cdot \mathbf{A}(x) = 0$

We can impose the Coulomb gauge by acting with a projection operator:

$$A_i(x) \rightarrow \left(\delta_{ij} - \frac{\nabla_i \nabla_j}{\nabla^2}\right) A_j(x)$$

in the momentum space it corresponds to multiplying $\tilde{A}_i(k)$ by the matrix $\delta_{ij} - k_i k_j / \mathbf{k}^2$, that projects out the longitudinal component.

(also known as transverse gauge)

the lagrangian in terms of scalar and vector potentials:

$$egin{aligned} \mathcal{L} &= -rac{1}{4}F^{\mu
u}F_{\mu
u} + J^{\mu}A_{\mu} \ &= -rac{1}{2}\partial^{\mu}A^{
u}\partial_{\mu}A_{
u} + rac{1}{2}\partial^{\mu}A^{
u}\partial_{
u}A_{\mu} + J^{\mu}A_{\mu} \ &\mathcal{L} &= rac{1}{2}\dot{A}_{i}\dot{A}_{i} - rac{1}{2}
abla_{j}A_{i}
abla_{j}A_{i} + J_{i}A_{i} \end{aligned}$$

$$egin{aligned} &+rac{1}{2}
abla_iA_j
abla_jA_i+\dot{A}_i
abla_iarphi\ &+rac{1}{2}
abla_iarphi
abla_iarphi-
hoarphi\ . \end{aligned}$$

$$\mathcal{L} = \frac{1}{2}\dot{A}_{i}\dot{A}_{i} - \frac{1}{2}\nabla_{j}A_{i}\nabla_{j}A_{i} + J_{i}A_{i} + \frac{1}{2}\nabla_{i}A_{j}\nabla_{j}A_{i} + \dot{A}_{i}\nabla_{i}\varphi$$
integration by parts
$$+\frac{1}{2}\nabla_{i}\varphi\nabla_{i}\varphi - \rho\varphi$$
integration by parts
$$\nabla_{i}\dot{A}_{i} \longrightarrow 0$$
equation of motion
$$\nabla_{i}A_{i} = 0$$
Poisson's equation unique solution:
$$\varphi(\mathbf{x}, t) = \int d^{3}y \frac{\rho(\mathbf{y}, t)}{4\pi |\mathbf{x} - \mathbf{y}|}$$
we get the lagrangian:
$$\mathcal{L} = \frac{1}{2}\dot{A}_{i}\dot{A}_{i} - \frac{1}{2}\nabla_{j}A_{i}\nabla_{j}A_{i} + J_{i}A_{i} + \mathcal{L}_{coul}$$

$$\mathcal{L}_{coul} = -\frac{1}{2}\int d^{3}y \frac{\rho(\mathbf{x}, t)\rho(\mathbf{y}, t)}{4\pi |\mathbf{x} - \mathbf{y}|}$$

$$\mathcal{L} = \frac{1}{2}\dot{A}_i\dot{A}_i - \frac{1}{2}\nabla_j A_i\nabla_j A_i + J_i A_i + \mathcal{L}_{\text{coul}}$$

the equation of motion for a free field $(J_i = 0)$:

$$-\partial^2 A_i(x) = 0$$

massless Klein-Gordon equation

the general solution:

$$\mathbf{A}(x) = \sum_{\lambda=\pm} \int \widetilde{dk} \begin{bmatrix} \boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k})a_{\lambda}(\mathbf{k})e^{ikx} + \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k})a_{\lambda}^{\dagger}(\mathbf{k})e^{-ikx} \end{bmatrix} \\ k^{0} = \omega = |\mathbf{k}| \\ \widetilde{dk} = d^{3}k/(2\pi)^{3}2\omega$$
polarization vectors (orthogonal to k)

we can choose the polarization vectors to correspond to right- and left-handed circular polarizations:

$$\begin{split} \boldsymbol{\varepsilon}_{+}(\mathbf{k}) &= \frac{1}{\sqrt{2}}(1, -i, 0) & \mathbf{k} = (0, 0, k) \\ \boldsymbol{\varepsilon}_{-}(\mathbf{k}) &= \frac{1}{\sqrt{2}}(1, +i, 0) \\ \mathbf{k} \cdot \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) &= 0 , \\ \boldsymbol{\varepsilon}_{\lambda'}(\mathbf{k}) \cdot \boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k}) &= \delta_{\lambda'\lambda} , \\ \boldsymbol{\varepsilon}_{\lambda'}(\mathbf{k}) \cdot \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) &= \delta_{ij} - \frac{k_{i}k_{j}}{\mathbf{k}^{2}} . \end{split}$$

in general:

$$\mathbf{A}(x) = \sum_{\lambda=\pm} \int \widetilde{dk} \left[\boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k}) a_{\lambda}(\mathbf{k}) e^{ikx} + \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) a_{\lambda}^{\dagger}(\mathbf{k}) e^{-ikx} \right]$$

following the procedure used for a scalar field we can express the operators in terms of fields:

$$egin{aligned} a_\lambda(\mathbf{k}) &= +i\,oldsymbol{arepsilon}_\lambda(\mathbf{k})\cdot\int d^3\!x\;e^{-ikx}\,\overleftrightarrow{\partial_0}\mathbf{A}(x)\ a^\dagger_\lambda(\mathbf{k}) &= -i\,oldsymbol{arepsilon}_\lambda^*(\mathbf{k})\cdot\int d^3\!x\;e^{+ikx}\,\overleftrightarrow{\partial_0}\mathbf{A}(x)\ f\,\overleftrightarrow{\partial_\mu}g &= f(\partial_\mu g) - (\partial_\mu f)g \end{aligned}$$

to find the hamiltonian we start with the conjugate momenta:

$$\mathcal{L} = \frac{1}{2}\dot{A}_{i}\dot{A}_{i} - \frac{1}{2}\nabla_{j}A_{i}\nabla_{j}A_{i} + J_{i}A_{i} + \mathcal{L}_{coul}$$
$$\Pi_{i} = \frac{\partial\mathcal{L}}{\partial\dot{A}_{i}} = \dot{A}_{i}$$
$$\nabla_{i}A_{i} = 0 \longrightarrow \nabla_{i}\Pi_{i} = 0$$

the hamiltonian density is then

$$\mathcal{H} = \Pi_i \dot{A}_i - \mathcal{L}$$

= $\frac{1}{2} \Pi_i \Pi_i + \frac{1}{2} \nabla_j A_i \nabla_j A_i - J_i A_i + \mathcal{H}_{coul}$

$$\mathcal{H}_{\mathrm{coul}} = -\mathcal{L}_{\mathrm{coul}}$$

we impose the canonical commutation relations:

$$A_i(x)
ightarrow igg(\delta_{ij} - rac{
abla_i
abla_j}{
abla^2}igg) A_j(x)$$

with the projection operator

$$\begin{split} [A_i(\mathbf{x},t),\Pi_j(\mathbf{y},t)] &= i \bigg(\delta_{ij} - \frac{\nabla_i \nabla_j}{\nabla^2} \bigg) \delta^3(\mathbf{x} - \mathbf{y}) \\ &= i \int \frac{d^3k}{(2\pi)^3} \, e^{i\mathbf{k} \cdot (\mathbf{x} - \mathbf{y})} \bigg(\delta_{ij} - \frac{k_i k_j}{\mathbf{k}^2} \bigg) \end{split}$$

$$[A_i, A_j] = [\Pi_i, \Pi_j] = 0$$

these correspond to the canonical commutation relations for creation and annihilation operators: (the same procedure as for the scalar field)

$$egin{aligned} &[a_\lambda(\mathbf{k}),a_{\lambda'}(\mathbf{k}')]=0\ ,\ &[a_\lambda^\dagger(\mathbf{k}),a_{\lambda'}^\dagger(\mathbf{k}')]=0\ ,\ &[a_\lambda(\mathbf{k}),a_{\lambda'}^\dagger(\mathbf{k}')]=(2\pi)^32\omega\,\delta^3(\mathbf{k}'-\mathbf{k})\delta_{\lambda\lambda'}\ . \end{aligned}$$

creation and annihilation operators for photons with helicity +1 (right-circular polarization) and -1 (left-circular polarization) now we can write the hamiltonian in terms of creation and annihilation operators: $\mathcal{H} = \Pi_i \dot{A}_i - \mathcal{L}$

$$= \frac{1}{2}\Pi_i\Pi_i + \frac{1}{2}\nabla_j A_i \nabla_j A_i - J_i A_i + \mathcal{H}_{\text{coul}}$$

(the same procedure as for the scalar field)

$$\begin{split} H &= \sum_{\lambda = \pm} \int \widetilde{dk} \; \omega \; a_{\lambda}^{\dagger}(\mathbf{k}) a_{\lambda}(\mathbf{k}) + 2\mathcal{E}_{0}V - \int d^{3}x \; \mathbf{J}(x) \cdot \mathbf{A}(x) + H_{\mathrm{coul}} \\ H_{\mathrm{coul}} &= \frac{1}{2} \int d^{3}x \; d^{3}y \; \frac{\rho(\mathbf{x}, t)\rho(\mathbf{y}, t)}{4\pi |\mathbf{x} - \mathbf{y}|} \\ \text{2-times the zero-point} \\ \text{energy of a scalar field} \\ \mathcal{E}_{0} &= \frac{1}{2} (2\pi)^{-3} \int d^{3}k \; \omega \end{split}$$

this form of the hamiltonian of electrodynamics is used in calculations of atomic transition rates, in particle physics the hamiltonian doesn't play a special role; we start with the lagrangian with specific interactions, calculate correlation functions, plug them into LSZ to get transition amplitudes ...

LSZ reduction for photons

based on S-56

Next step is to get the LSZ formula for the photon. The derivation closely follows the scalar field case; the only difference is due to the presence of polarization vectors: For a scalar field we found that in order to obtain a transition amplitude we simply replace the creation and annihilation operators in the transition amplitude by:

$$a^{\dagger}(\mathbf{k})_{\mathrm{in}} \rightarrow i \int d^4 z_1 \ e^{+ikz_1}(-\partial^2 + m^2)\varphi(z_1)$$

 $a(\mathbf{k}')_{\mathrm{out}} \rightarrow i \int d^4 z_2 \ e^{-ik'z_2}(-\partial^2 + m^2)\varphi(z_2)$

similarly, for an incoming and outgoing photon we simply replace:

$$\begin{split} \mathbf{A}(x) &= \sum_{\lambda=\pm} \int \widetilde{dk} \left[\boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k}) a_{\lambda}(\mathbf{k}) e^{ikx} + \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) a_{\lambda}^{\dagger}(\mathbf{k}) e^{-ikx} \right] \\ \boldsymbol{\varepsilon}_{\lambda}^{0}(\mathbf{k}) &\equiv 0 & a_{\lambda}^{\dagger}(\mathbf{k}) = -i \, \boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k}) \cdot \int d^{3}x \, e^{+ikx} \overleftrightarrow{\partial_{0}} \mathbf{A}(x) \\ a_{\lambda}(\mathbf{k}) &= +i \, \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) \cdot \int d^{3}x \, e^{-ikx} \overleftrightarrow{\partial_{0}} \mathbf{A}(x) \\ a_{\lambda}(\mathbf{k})_{\mathrm{in}} &\to i \, \boldsymbol{\varepsilon}_{\lambda}^{\mu*}(\mathbf{k}) \int d^{4}x \, e^{+ikx} (-\partial^{2}) A_{\mu}(x) \\ a_{\lambda}(\mathbf{k})_{\mathrm{out}} &\to i \, \boldsymbol{\varepsilon}_{\lambda}^{\mu}(\mathbf{k}) \int d^{4}x \, e^{-ikx} (-\partial^{2}) A_{\mu}(x) \end{split}$$

$$\mathbf{A}(x) = \sum_{\lambda=\pm} \int \widetilde{dk} \left[oldsymbol{arepsilon}_{\lambda}^{*}(\mathbf{k}) a_{\lambda}(\mathbf{k}) e^{ikx} + oldsymbol{arepsilon}_{\lambda}(\mathbf{k}) a_{\lambda}^{\dagger}(\mathbf{k}) e^{-ikx}
ight]$$

the LSZ formula is then valid if the field is normalized according to the free field formulae:

$$egin{aligned} &\langle 0|A^i(x)|0
angle &= 0 \ , \ &\langle k,\lambda|A^i(x)|0
angle &= arepsilon_\lambda^i({f k})ar e^{ikx} \end{aligned}$$

where a single photon state is normalized according to:

$$\langle k', \lambda' | k, \lambda \rangle = (2\pi)^3 2\omega \, \delta^3 (\mathbf{k}' - \mathbf{k}) \delta_{\lambda\lambda'}$$

and the renormalization of fields results in the Z-factors in the lagrangian:

$$\mathcal{L} = -\frac{1}{4}Z_3 F^{\mu\nu}F_{\mu\nu} + Z_1 J^{\mu}A_{\mu}$$

we will discuss this next semester...

Now we want to calculate correlation functions (the derivation again closely follows the scalar field case).

$$\mathbf{A}(x) = \sum_{\lambda=\pm} \int \widetilde{dk} \left[\boldsymbol{\varepsilon}_{\lambda}^{*}(\mathbf{k}) a_{\lambda}(\mathbf{k}) e^{ikx} + \boldsymbol{\varepsilon}_{\lambda}(\mathbf{k}) a_{\lambda}^{\dagger}(\mathbf{k}) e^{-ikx} \right]$$

the propagator for a free field theory:

$$\langle 0|\mathrm{T}A^{i}(x)A^{j}(y)|0
angle = rac{1}{i}\Delta^{ij}(x-y)$$

$$\Delta^{ij}(x-y) = \int \frac{d^4k}{(2\pi)^4} \, \frac{e^{ik(x-y)}}{k^2 - i\epsilon} \sum_{\lambda=\pm} \varepsilon_{\lambda}^{i*}(\mathbf{k}) \varepsilon_{\lambda}^j(\mathbf{k})$$

correlation functions of more fields given in terms of propagators...

Next we want to calculate the path integral for the free EM field:

$$Z_0(J) \equiv \langle 0|0\rangle_J = \int \mathcal{D}A \ e^{i\int d^4x \left[-\frac{1}{4}F^{\mu\nu}F_{\mu\nu} + J^{\mu}A_{\mu}\right]}$$

we will treat the current as an external source

$$Z_0(J)\equiv \langle 0|0
angle_J=\int {\cal D}\!A\;e^{i\int d^4\!x\left[-rac{1}{4}F^{\mu
u}F_{\mu
u}+J^\mu\!A_\mu
ight]}$$

In the Coulomb gauge we integrate over those field configurations that satisfy $\nabla \cdot \mathbf{A}(x) = 0$; in addition the zero's component is not dynamical we can replace it by the solution of the equation of motion

$$\mathcal{L}_{
m coul} = -rac{1}{2} \int d^3y \; rac{
ho({f x},t)
ho({f y},t)}{4\pi |{f x}-{f y}|} \ S_{
m coul} = -rac{1}{2} \int d^4x \; d^4y \; \delta(x^0 - y^0) \; rac{J^0(x)J^0(y)}{4\pi |{f x}-{f y}|}$$

and for the rest of the path integral we will guess the result based on the result we got for a scalar field:

$$Z_0(J) = \exp\left[iS_{\text{coul}} + \frac{i}{2}\int d^4x \, d^4y \, J_i(x)\Delta^{ij}(x-y)J_j(y)\right]$$
propagator

$$Z_0(J) = \exp\left[iS_{\text{coul}} + \frac{i}{2}\int d^4x \, d^4y \, J_i(x)\Delta^{ij}(x-y)J_j(y)\right]$$
$$S_{\text{coul}} = -\frac{1}{2}\int d^4x \, d^4y \, \delta(x^0 - y^0) \, \frac{J^0(x)J^0(y)}{4\pi|\mathbf{x} - \mathbf{y}|} \qquad \Delta^{ij}(x-y) = \int \frac{d^4k}{(2\pi)^4} \, \frac{e^{ik(x-y)}}{k^2 - i\epsilon} \sum_{\lambda = \pm} \varepsilon_{\lambda}^{i*}(\mathbf{k})\varepsilon_{\lambda}^j(\mathbf{k})$$

we can make it look better:

$$Z_0(J)=\expiggl[rac{i}{2}\int d^4x\,d^4y\,J_\mu(x)\Delta^{\mu
u}(x-y)J_
u(y)iggr]$$

where

$$\Delta^{\mu
u}(x-y) \equiv \int \frac{d^4k}{(2\pi)^4} e^{ik(x-y)} \tilde{\Delta}^{\mu
u}(k)$$

 $\tilde{\Delta}^{\mu
u}(k) \equiv -\frac{1}{\mathbf{k}^2} \delta^{\mu 0} \delta^{
u 0} + \frac{1}{k^2 - i\epsilon} \sum_{\lambda=\pm} \epsilon^{\mu*}_{\lambda}(\mathbf{k}) \epsilon^{
u}_{\lambda}(\mathbf{k})$

 $arepsilon_\lambda^0({f k})\equiv 0$

and the Coulomb term is reproduced thanks to:

$$\int_{-\infty}^{+\infty} \frac{dk^0}{2\pi} e^{-ik^0(x^0 - y^0)} = \delta(x^0 - y^0)$$
$$\int \frac{d^3k}{(2\pi)^3} \frac{e^{i\mathbf{k} \cdot (\mathbf{x} - \mathbf{y})}}{\mathbf{k}^2} = \frac{1}{4\pi |\mathbf{x} - \mathbf{y}|}$$

$$ilde{\Delta}^{\mu
u}(k) \equiv -rac{1}{\mathbf{k}^2} \delta^{\mu 0} \delta^{
u 0} + rac{1}{k^2 - i\epsilon} \sum_{\lambda=\pm} \varepsilon^{\mu*}_{\lambda}(\mathbf{k}) \varepsilon^{
u}_{\lambda}(\mathbf{k})$$

 $\hat{t} \cdot k = -k^0$

We can simplify the propagator further... Let's define:

$$\hat{t}^{\mu} = (1, \mathbf{0})$$

and \hat{z}^{μ} as a unit vector in the **k** direction:

$$\hat{z}^{\mu} = \frac{k^{\mu} + (\hat{t} \cdot k)\hat{t}^{\mu}}{[k^2 + (\hat{t} \cdot k)^2]^{1/2}} \qquad (0, \mathbf{k}) = k^{\mu} + (\hat{t} \cdot k)\hat{t}^{\mu}$$
$$\hat{t}^2 = -1$$
$$\mathbf{k}^2 = k^2 + (\hat{t} \cdot k)^2$$

now we can replace:

and thus we get:

$$ilde{\Delta}^{\mu
u}(k) = -\,rac{\hat{t}^{\mu}\hat{t}^{
u}}{k^2 + (\hat{t}\cdot k)^2} + rac{g^{\mu
u} + \hat{t}^{\mu}\hat{t}^{
u} - \hat{z}^{\mu}\hat{z}^{
u}}{k^2 - i\epsilon}$$

$$\begin{split} \tilde{\Delta}^{\mu\nu}(k) &= -\frac{\hat{t}^{\mu}\hat{t}^{\nu}}{k^{2} + (\hat{t}\cdot k)^{2}} + \frac{g^{\mu\nu} + \hat{t}^{\mu}\hat{t}^{\nu} - \hat{z}^{\mu}\hat{z}^{\nu}}{k^{2} - i\epsilon} \\ \hat{t}^{\mu} &= (1, \mathbf{0}) \\ \hat{z}^{\mu} &= \underbrace{k^{\mu} + (\hat{t}\cdot k)\hat{t}^{\mu}}_{[k^{2} + (\hat{t}\cdot k)^{2}]^{1/2}} \end{split}$$

this looks better but we can simplify the propagator further.

 $Z_{0}(J) = \exp\left[\frac{i}{2} \int d^{4}x \, d^{4}y \, J_{\mu}(x) \Delta^{\mu\nu}(x-y) J_{\nu}(y)\right]$ the momentum can be replaced $\Delta^{\mu\nu}(x-y) \equiv \int \frac{d^{4}k}{(2\pi)^{4}} e^{ik(x-y)} \tilde{\Delta}^{\mu\nu}(k)$ by the derivative with respect to x^{μ} acting on the exponential, and then integrate by parts to obtain $\partial^{\mu}J_{\mu}(x)$ which vanishes. $\hat{z}^{\mu} \rightarrow \frac{(\hat{t} \cdot k)\hat{t}^{\mu}}{[k^{2} + (\hat{t} \cdot k)^{2}]^{1/2}}$

and we get:

$$\begin{split} \tilde{\Delta}^{\mu\nu}(k) &= \frac{1}{k^2 - i\epsilon} \left[g^{\mu\nu} + \left(-\frac{k^2}{k^2 + (\hat{t} \cdot k)^2} + 1 - \frac{(\hat{t} \cdot k)^2}{k^2 + (\hat{t} \cdot k)^2} \right) \hat{t}^{\mu} \hat{t}^{\nu} \right] \\ &= 0 \end{split}$$

We obtained a very simple formula for the photon propagator:

$$\Delta^{\mu
u}(x-y) \equiv \int rac{d^4k}{(2\pi)^4} e^{ik(x-y)} \, ilde{\Delta}^{\mu
u}(k) \ ilde{\Delta}^{\mu
u}(k) = rac{g^{\mu
u}}{k^2 - i\epsilon}$$

Feynman gauge

(it would still be in the Coulomb gauge if we had kept the terms proportional to momenta)